

# Van der Waals interactions in Hadron Resonance Gas: From nuclear matter to lattice QCD

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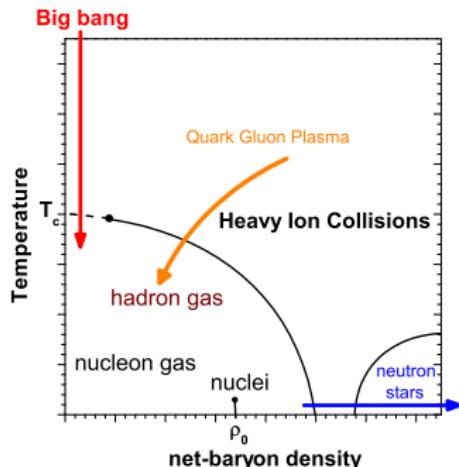
FIAS Frankfurt Institute  
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HGS-HIRe *for FAIR*  
Helmholtz Graduate School for Hadron and Ion Research

# Strongly interacting matter

- Theory of strong interactions: Quantum Chromodynamics (QCD)
- Basic degrees of freedom: quarks and gluons
- At smaller energies confined into hadrons (baryons and mesons)



Where is it relevant?

- Early universe
- Neutron stars
- Heavy-ion collisions

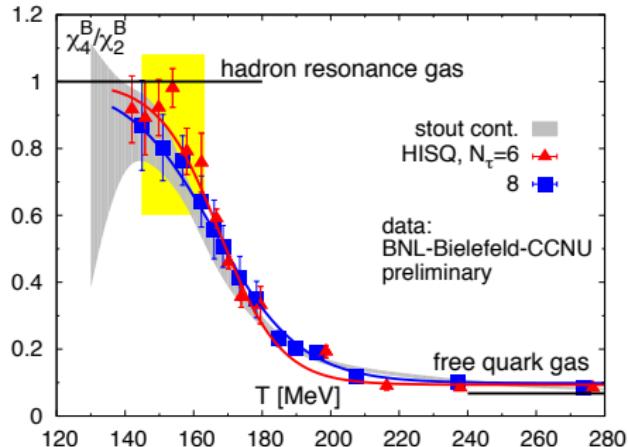
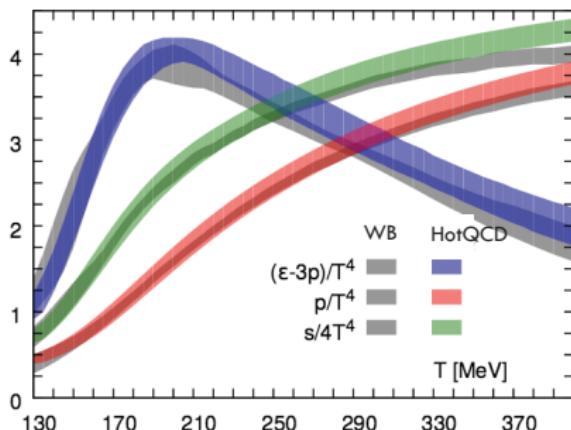
First principles of QCD are rather established,  
but direct calculations are problematic

Phenomenological tools are very useful

Experiment: **heavy-ion collisions**

# QCD equation of state at $\mu = 0$

Lattice simulations provide EoS at  $\mu = 0^1$



Common model for confined phase is **ideal HRG**: non-interacting gas of known hadrons and resonances

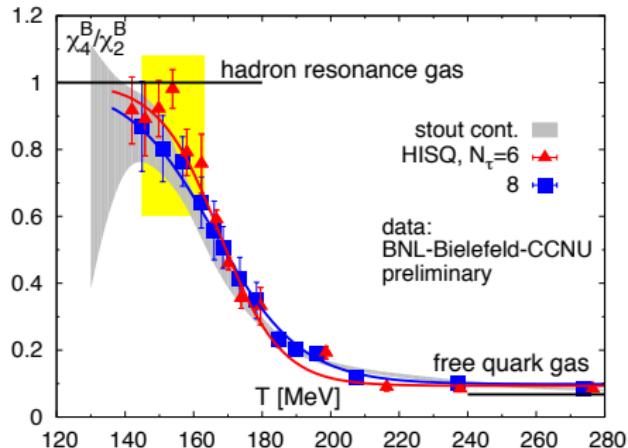
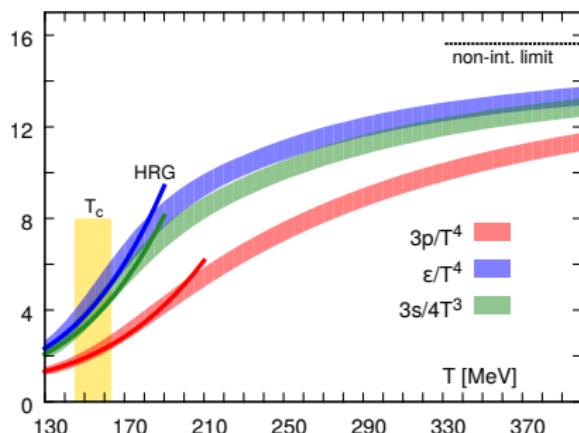
- Good description of thermodynamic functions up to 180 MeV
- Rapid **breakdown** in crossover region for description of **susceptibilities**<sup>2</sup>
- Often interpreted as clear signal of deconfinement...
- **But what is the role of hadronic interactions** beyond normal HRG?

<sup>1</sup>Bazavov et al., PRD 90, 094503 (2014); Borsanyi et al., PLB 730, 99 (2014)

<sup>2</sup>Ding, Karsch, Mukherjee, IJMPE 24, 1530007 (2015)

# QCD equation of state at $\mu = 0$

Lattice simulations provide EoS at  $\mu = 0^1$



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# Van der Waals equation

## Van der Waals equation

$$P(T, V, N) = \frac{NT}{V - bN} - a \frac{N^2}{V^2}$$



Formulated in  
1873.

Simplest model which contains attractive and repulsive interactions

Contains 1st order phase transition and critical point

Can elucidate role of fluctuations in phase transitions



Nobel Prize in  
1910.

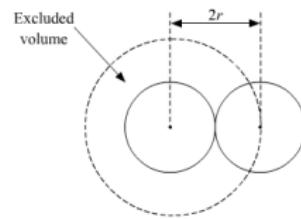
Two ingredients:

1) Short-range repulsion: excluded volume procedure,

$$V \rightarrow V - bN, \quad b = 4 \frac{4\pi r^3}{3}$$

2) Attractive interactions in mean-field approximation,

$$P \rightarrow P - a n^2$$

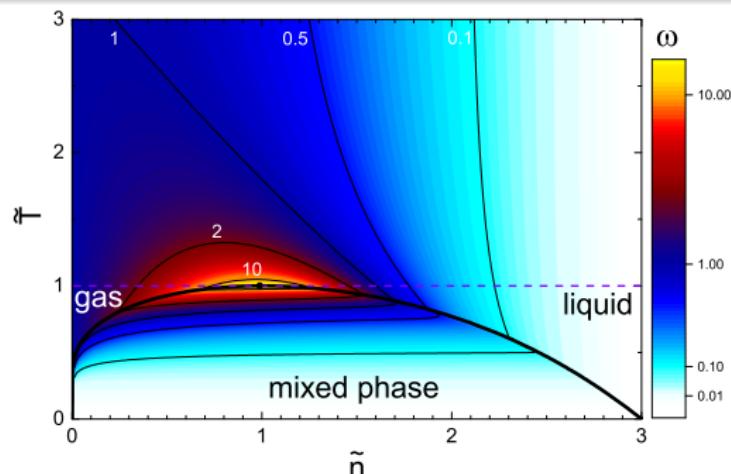


# Scaled variance for classical VDW equation

Transformation of VDW eq. to GCE leads to new physical applications<sup>1</sup>

Scaled variance of particle number fluctuations in VDW gas (pure phases)

$$\omega[N] = \frac{\sigma^2}{\langle N \rangle} = \left[ \frac{1}{(1 - bn)^2} - \frac{2an}{T} \right]^{-1}$$



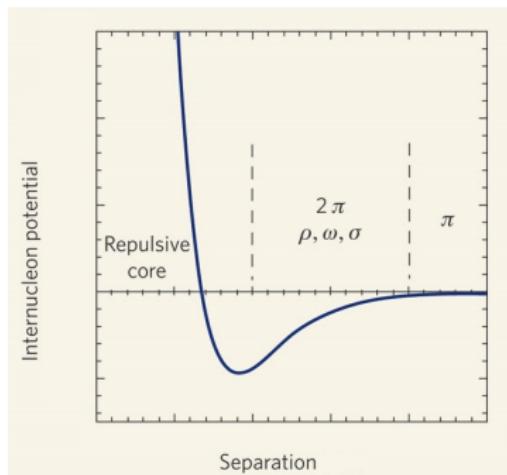
- Repulsive interactions suppress N-fluctuations
- Attractive interactions enhance N-fluctuations

<sup>1</sup>Vovchenko et al., J. Phys. A 305001, 48 (2015) and D. Anchishkin's talk on Monday

# Nucleon-nucleon interaction

Nucleon-nucleon potential:

- Repulsive core at small distances
- Attraction at intermediate distances
- Suggestive similarity to VDW interactions
- Could nuclear matter described by VDW equation?



Standard VDW equation is for **canonical ensemble** and **Boltzmann** statistics

Nucleons are fermions, obey Pauli exclusion principle

Unlike for classical fluids, **quantum statistics** is important

# VDW equation with quantum statistics in GCE

## Requirements for VDW equation with quantum statistics

- 1) Reduce to **ideal quantum gas** at  $a = 0$  and  $b = 0$
- 2) Reduce to **classical VDW** when quantum statistics are negligible
- 3)  $s \geq 0$  and  $s \rightarrow 0$  as  $T \rightarrow 0$

**Ansatz:** Take pressure in the following form<sup>1,2</sup>

$$p(T, \mu) = p^{\text{id}}(T, \mu^*) - an^2, \quad \mu^* = \mu - bp - abn^2 + 2an$$

where  $p^{\text{id}}(T, \mu^*)$  is pressure of ideal **quantum** gas.

$$n(T, \mu) = \left( \frac{\partial p}{\partial \mu} \right)_T = \frac{n^{\text{id}}(T, \mu^*)}{1 + b n^{\text{id}}(T, \mu^*)}$$

## Algorithm for GCE

- 1) Solve system of eqs. for  $p$  and  $n$  at given  $(T, \mu)$
- 2) Choose the solution with **largest** pressure

<sup>1</sup>V. Vovchenko, D. Anchishkin, M. Gorenstein, Phys. Rev. C 91, 064314 (2015)

<sup>2</sup>**Alternative derivation:** K. Redlich, K. Zalewski, arXiv:1605.09686 (2016)

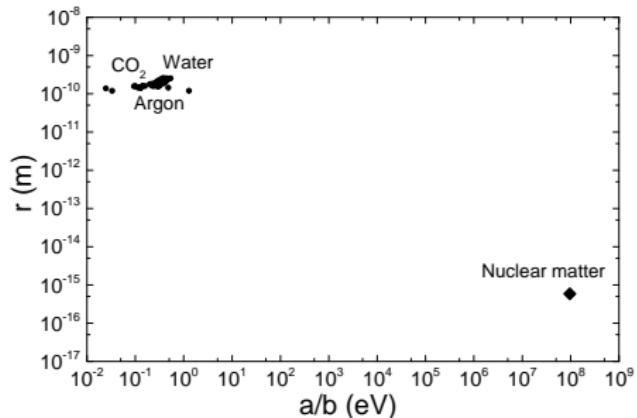
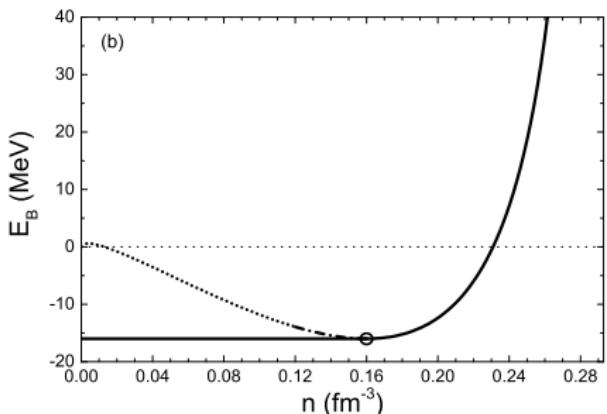
<sup>3</sup> $a=0 \Rightarrow$  **excluded-volume** model, D. Rischke et al., Z.Phys. C51, 485 (1991)

# VDW gas of nucleons: zero temperature

How to fix  $a$  and  $b$ ? For classical fluid usually tied to CP location.  
Different approach: Reproduce **saturation density** and **binding energy**

From  $E_B = E/A \cong -16$  MeV and  $n = n_0 \cong 0.16 \text{ fm}^{-3}$  at  $T = 0$  and  $p = 0$

$$a \cong 329 \text{ MeV fm}^3 \text{ and } b \cong 3.42 \text{ fm}^3$$



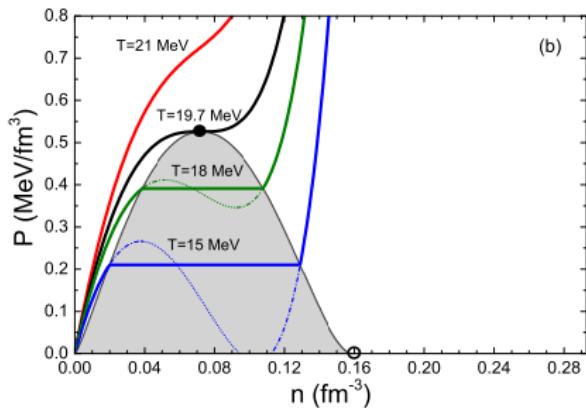
Mixed phase at  $T = 0$  is specific:  
A mix of vacuum ( $n = 0$ ) and liquid  
at  $n = n_0$

VDW eq. now at very different scale!

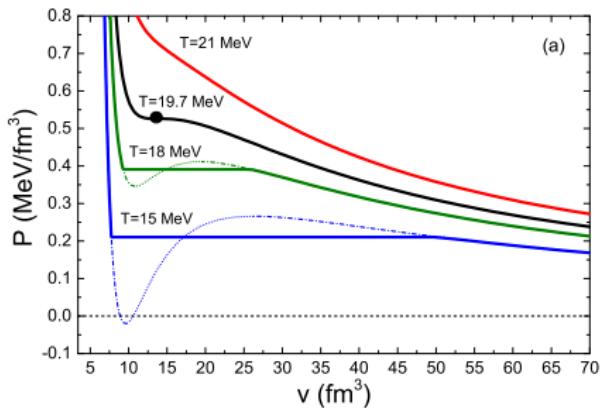
# VDW gas of nucleons: pressure isotherms

CE pressure

$$p = p^{\text{id}} \left[ T, \mu^{\text{id}} \left( \frac{n}{1 - bn}, T \right) \right] - an^2$$



(b)



(a)

Behavior qualitatively **same** as for Boltzmann case

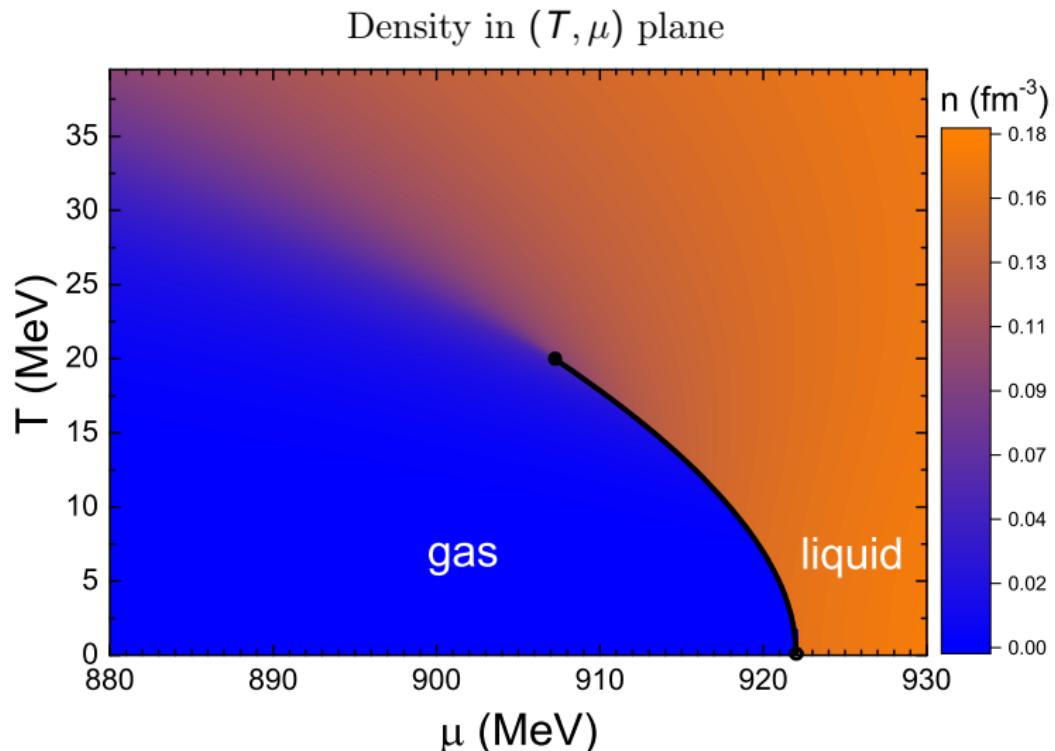
Mixed phase results from **Maxwell construction**

**Critical point** at  $T_c \cong 19.7$  MeV and  $n_c \cong 0.07$  fm $^{-3}$

**Experimental estimate<sup>1</sup>:**  $T_c = 17.9 \pm 0.4$  MeV,  $n_c = 0.06 \pm 0.01$  fm $^{-3}$

<sup>1</sup>J.B. Elliot, P.T. Lake, L.G. Moretto, L. Phair, Phys. Rev. C 87, 054622 (2013)

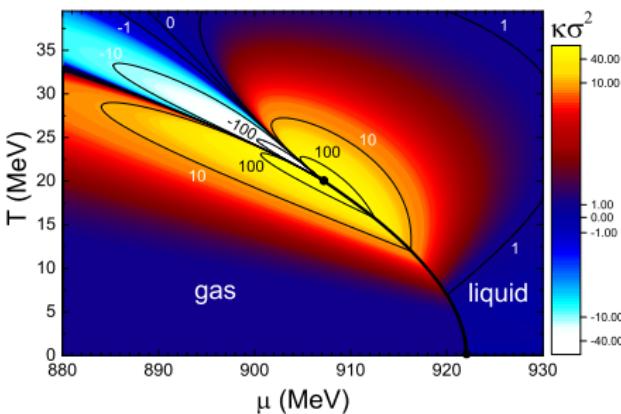
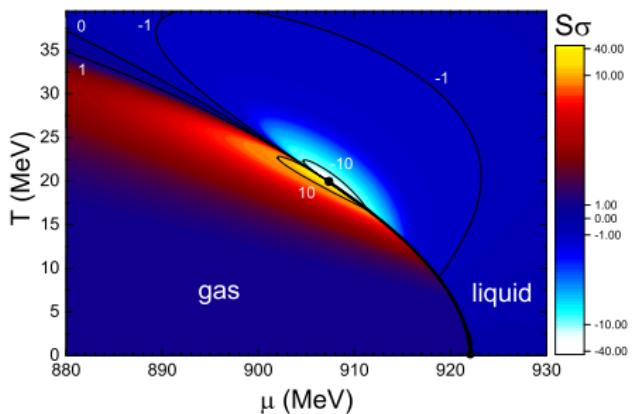
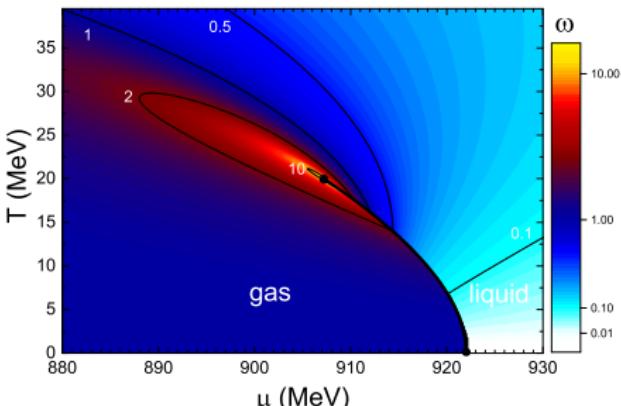
# VDW gas of nucleons: $(T, \mu)$ plane



Crossover region at  $\mu < \mu_C \cong 908$  MeV is clearly seen

# VDW-HRG gas of nucleons: fluctuations

- Results for  $\sigma^2/\langle N \rangle$ ,  $S\sigma$ , and  $\kappa\sigma^2$
- Fluctuations diverge at CP (as expected)
- For higher moments singularity is specific: sign depends on path of approach

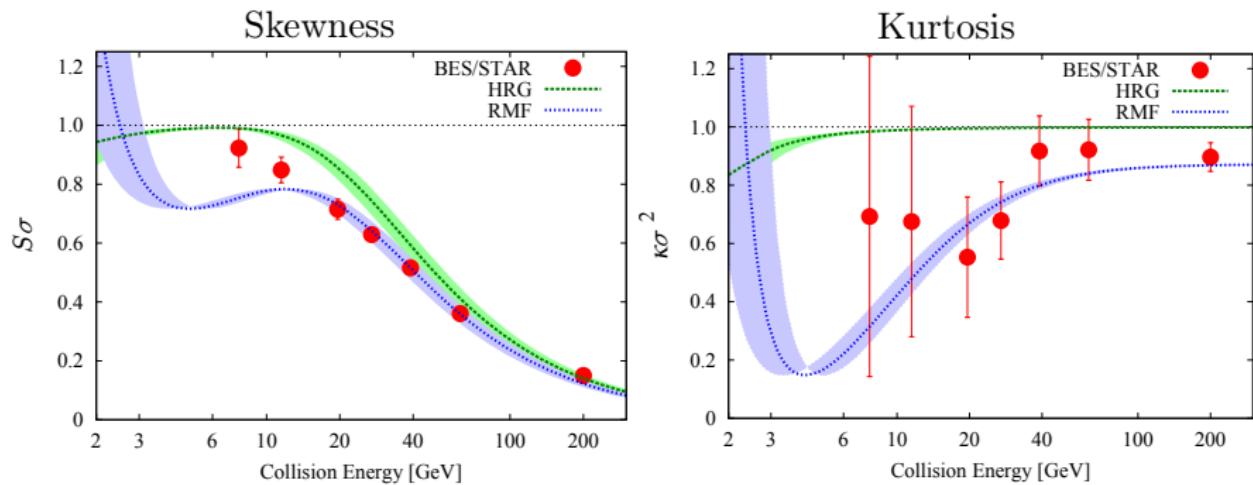


Vovchenko, Anchishkin, Gorenstein, Poberezhnyuk, PRC 91, 064314 (2015)

# Net-baryon fluctuations and nuclear matter

Are NN interactions relevant for observables in HIC region?

Net-nucleon fluctuations within RMF ( $\sigma$ - $\omega$  model) of nuclear matter along line of “chemical freeze-out”



K. Fukushima, PRC 91, 044910 (2015)

A notable effect in fluctuations even at  $\mu \simeq 0$

Reconciliation of HRG with nuclear matter can be interesting

# Van der Waals interactions in HRG

Simplest generalization of VDW nuclear matter model to full HRG:

- Same VDW interactions between all baryons
- The baryon-antibaryon, meson-meson, and meson-baryon VDW interactions are neglected
- Baryon VDW parameters extracted from ground state of nuclear matter ( $a = 329 \text{ MeV fm}^3$ ,  $b = 3.42 \text{ fm}^3$ )

Three independent subsystems: **mesons** + **baryons** + **antibaryons**

$$p(T, \mu) = P_M(T, \mu) + P_B(T, \mu) + P_{\bar{B}}(T, \mu),$$

$$P_M(T, \mu) = \sum_{j \in M} p_j^{\text{id}}(T, \mu_j) \quad \text{and} \quad P_B(T, \mu) = \sum_{j \in B} p_j^{\text{id}}(T, \mu_j^{B*}) - a n_B^2$$

$$n_B(T, \mu) = (1 - b n_B) \sum_{j \in B} n_j^{\text{id}}(T, \mu_j^{B*}).$$

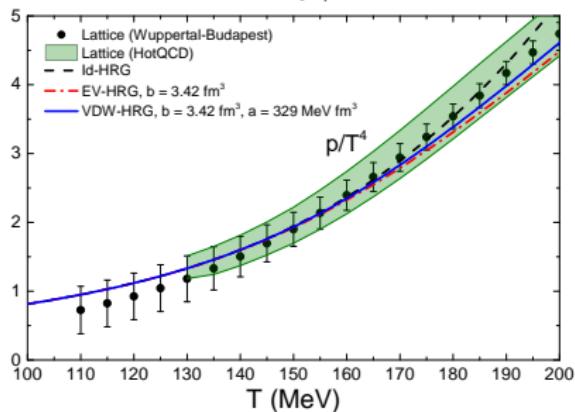
In this simplest setup model is essentially “**parameter-free**”

Transcendental equations for  $P_B$  and  $n_B$

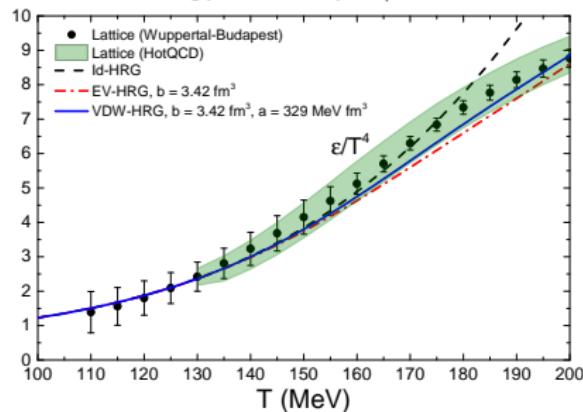
# VDW-HRG at $\mu = 0$ : thermodynamic functions

Comparison of VDW-HRG with lattice QCD at  $\mu = 0$

Pressure  $p/T^4$

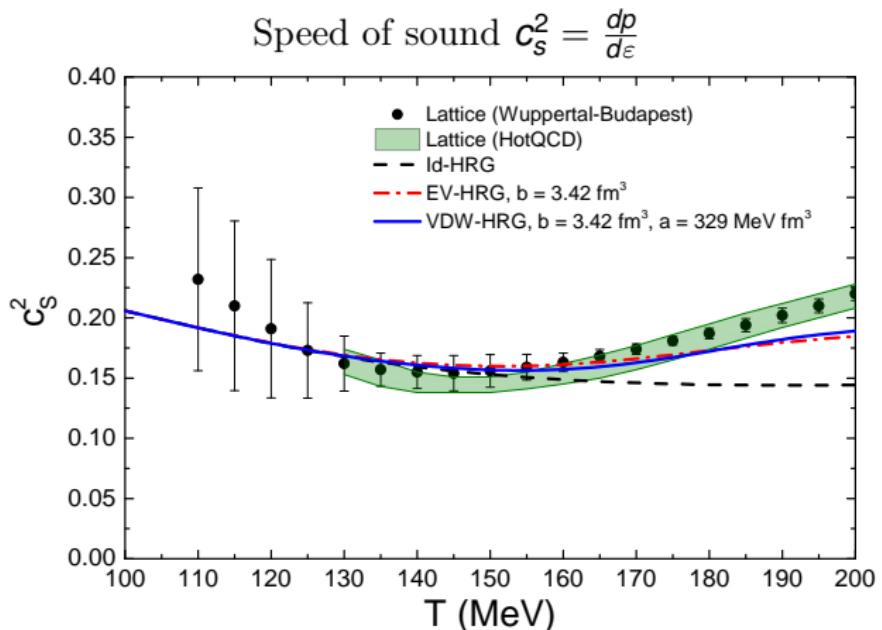


Energy density  $\varepsilon/T^4$



- VDW-HRG **does not spoil** existing agreement of Id-HRG with lQCD despite significant EV interactions between baryons
- Not surprising: matter **meson-dominated** at  $\mu = 0$

# VDW-HRG at $\mu = 0$ : speed of sound

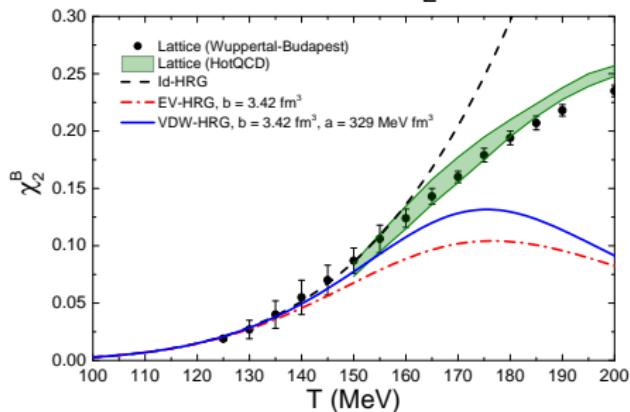


- Monotonic decrease in Id-HRG, at odds with lattice
- **Minimum** for EV-HRG/VDW-HRG at 150-160 MeV
- **No acausal** behavior, often an issue in models with eigenvalues

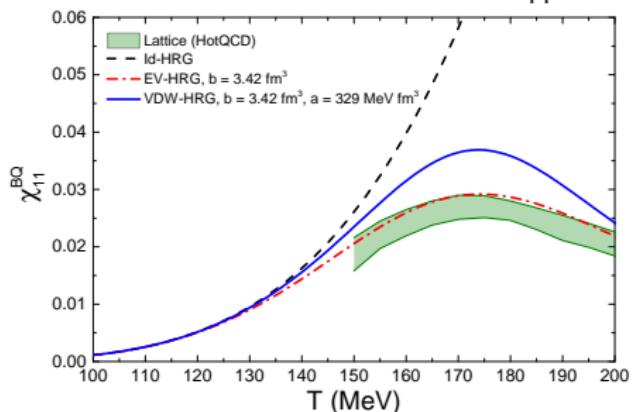
# VDW-HRG at $\mu = 0$ : baryon number fluctuations

Susceptibilities:  $\chi_{lmn}^{BSQ} = \frac{\partial^{l+m+n} p/T^4}{\partial(\mu_B/T)^l \partial(\mu_S/T)^m \partial(\mu_Q/T)^n}$

Net-baryon  $\chi_2^B$



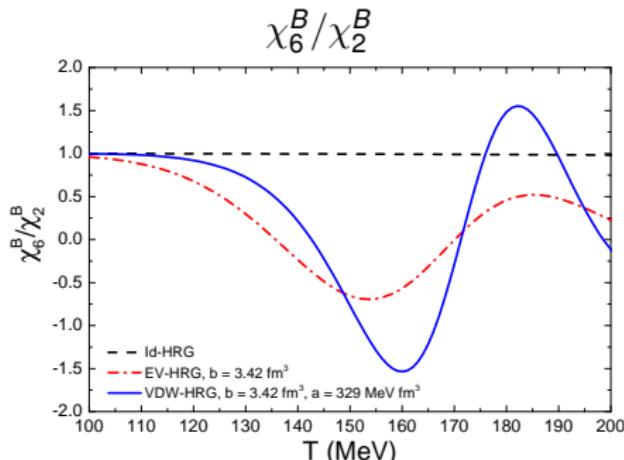
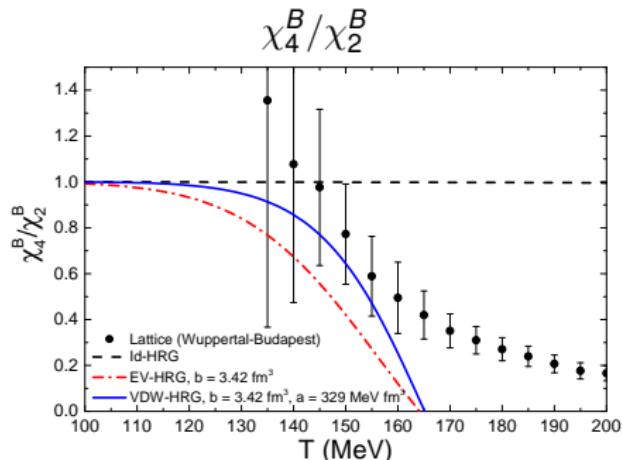
Baryon-charge correlator  $\chi_{11}^{BQ}$



- Very **different qualitative** behavior between Id-HRG and VDW-HRG
- For  $\chi_2^B$  lattice data is **between** Id-HRG and VDW-HRG at high T
- For  $\chi_{11}^{BQ}$  lattice data is below all models, closer to EV-HRG

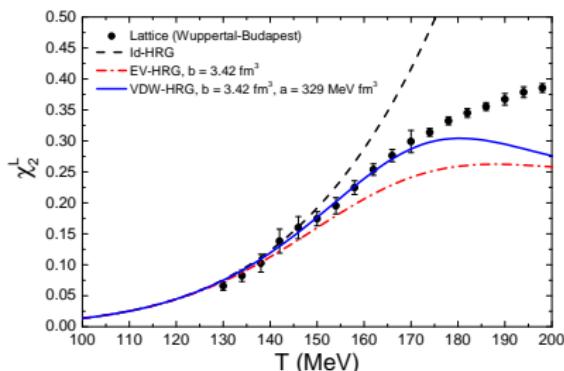
# VDW-HRG at $\mu = 0$ : baryon number fluctuations

Higher-order of fluctuations are expected to be even more sensitive



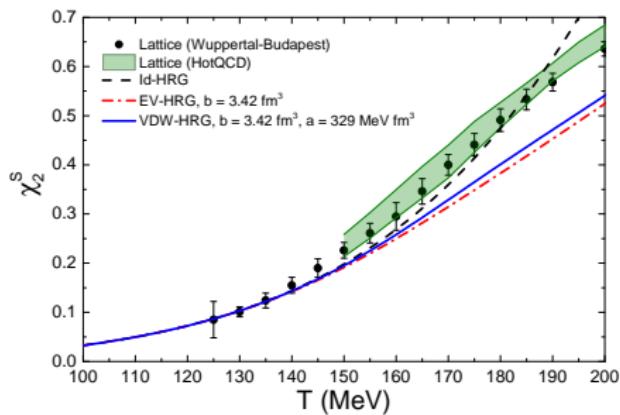
- $\chi_4^B$  deviates from  $\chi_2^B$  at high enough  $T$ , stays equal in Id-HRG
- Cannot be related only to onset of deconfinement
- VDW-HRG predicts strong **non-monotonic** behavior for  $\chi_6^B / \chi_2^B$

# VDW-HRG at $\mu = 0$ : net-light and net-strangeness



- Net number of light quarks  $\chi_2^L$
- $L = (u + d)/2 = (3B + S)/2$
- Improved description in VDW-HRG

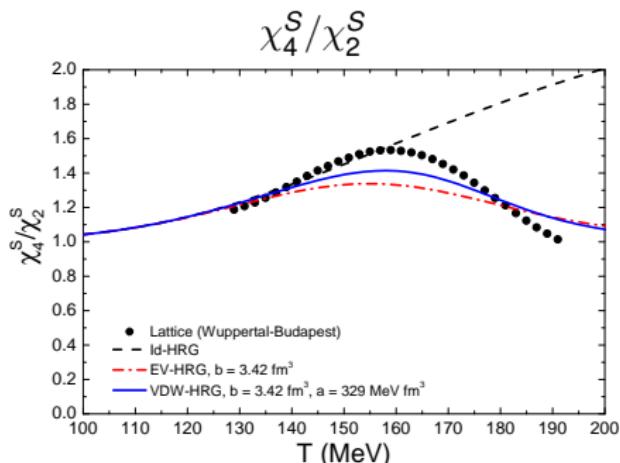
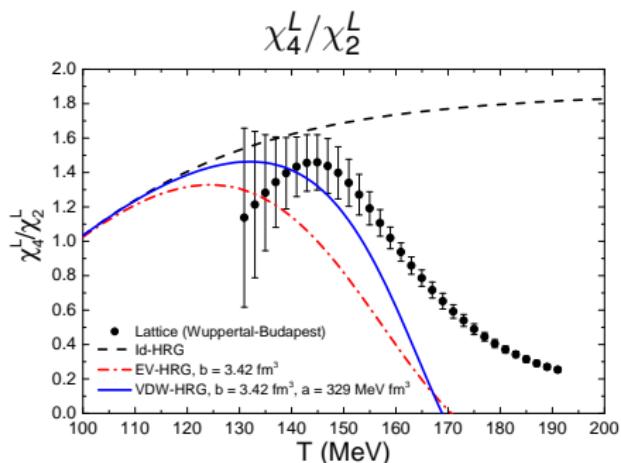
- Net-strangeness  $\chi_2^S$
- Underestimated by HRG models, similar for  $\chi_{11}^{BS}$
- Extra strange states?<sup>1</sup>
- Weaker VDW interactions for strange baryons?<sup>2</sup>



<sup>1</sup>Bazavov et al., PRL 113, 072001 (2014)

<sup>2</sup>Alba, Vovchenko, Gorenstein, Stoecker, arXiv:1606.06542

# VDW-HRG at $\mu = 0$ : net-light and net-strangeness



- Lattice shows **peaked structures** in crossover regions
- Not at all reproduced by Id-HRG, signal for deconfinement?<sup>1</sup>
- Peaks** at different  $T$  for net-L and net-S  $\Rightarrow$  **flavor hierarchy?**<sup>2</sup>
- VDW-HRG **also shows** peaks and flavor hierarchy  $\Rightarrow$  cannot be traced back directly to deconfinement

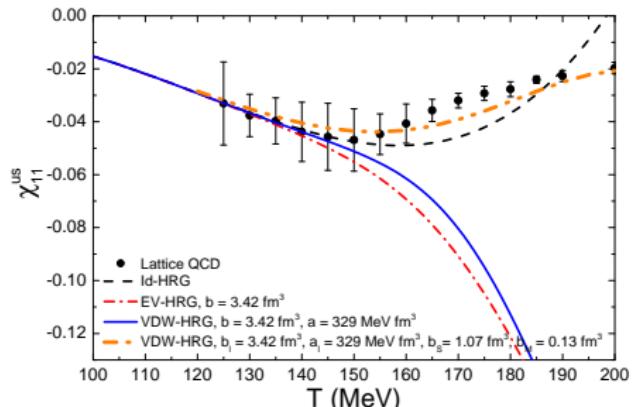
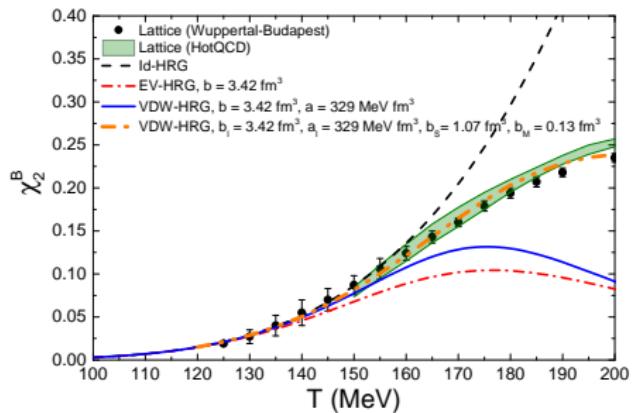
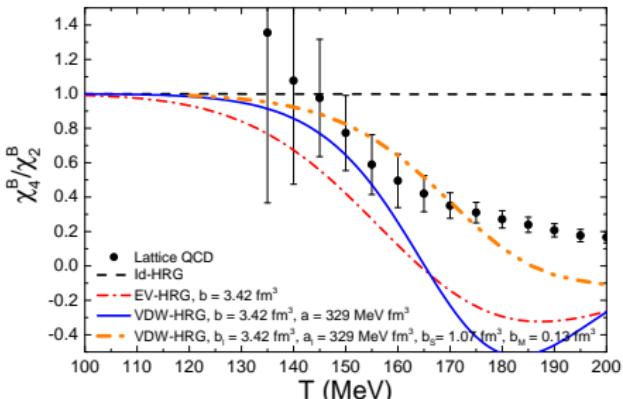
<sup>1</sup>S. Ejiri, F. Karsch, K. Redlich, PLB 633, 275 (2006)

<sup>2</sup>Bellwied et al., PRL 111, 202302 (2013)

# VDW-HRG: extensions

## Effect of reducing VDW interactions involving strange hadrons

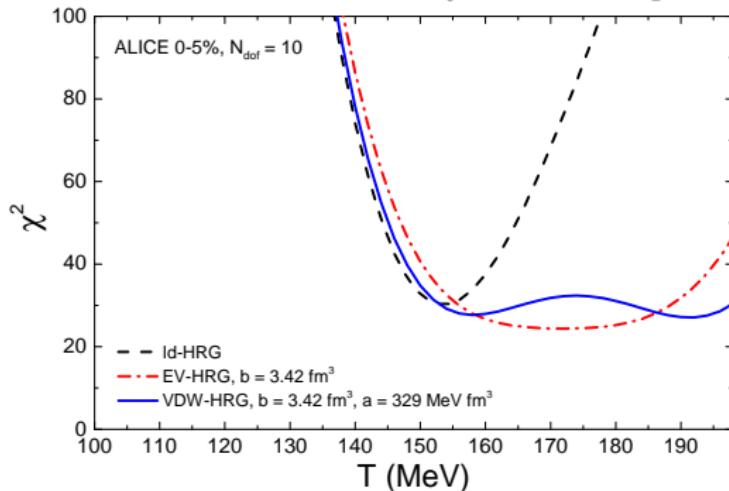
- 3 times smaller EV for strange baryons
- Small EV for mesons
- Illustrative calculation
- Most observables improved



# VDW-HRG: influence on hadron ratios

VDW interactions **change** relative **hadron yields**

Thermal fit to **ALICE** hadron yields: from pions to  $\Omega$

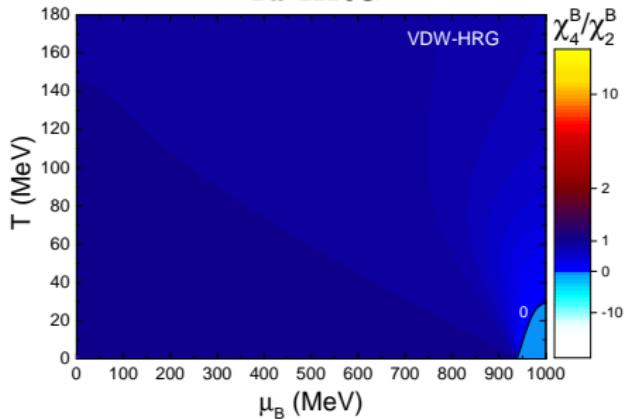


- Fit quality slightly **better** in EV-HRG/VDW-HRG vs Id-HRG but very **different picture!**
- **All** temperatures between 150 and 200 MeV yield similarly **fair** data description in VDW-HRG
- Results likely to be **sensitive** to further **modifications**, e.g for strangeness

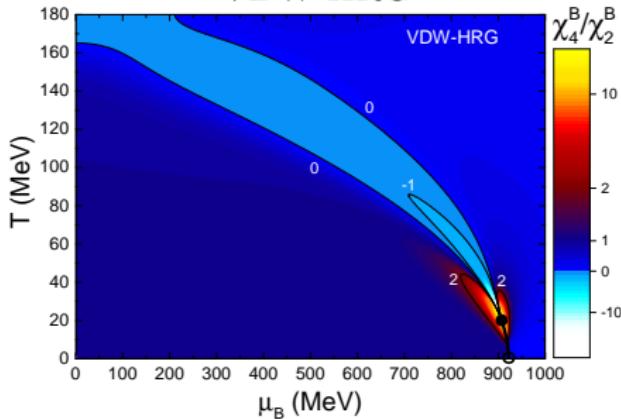
# VDW-HRG at finite $\mu_B$

Net-baryon fluctuations in  $T$ - $\mu_B$  plane:  $\chi_4^B/\chi_2^B$

Id-HRG



VDW-HRG



- Almost no effect in Id-HRG, only Fermi statistics
- Rather rich structure for VDW-HRG
- Fluctuations seen at RHIC are remnants of nuclear liquid-gas PT?
- In more involved setup interplay between nuclear liquid-gas and chiral transitions may be important<sup>1</sup>

<sup>1</sup>Mukherjee, Steinheimer, Schramm, arXiv:1611.10144

- VDW equation provides **simple insight** on fluctuations near CP
- **Nuclear matter** can be described as VDW equation with Fermi statistics
- VDW interactions between baryons have strong influence on **fluctuations of conserved charges** in the crossover region within HRG
- VDW-HRG captures **basic features** of both lattice results at  $\mu = 0$  and nuclear matter properties
- Freeze-out parameters extracted from thermal fits within ideal HRG are **not unique**, procedure sensitive to modeling of VDW interactions
- Interpretation of results obtained within standard **ideal HRG** should be done with extreme care

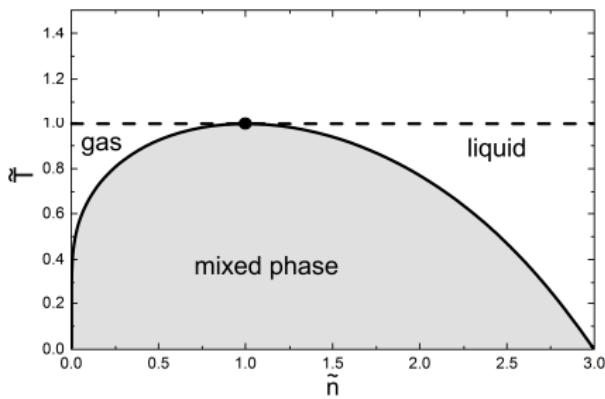
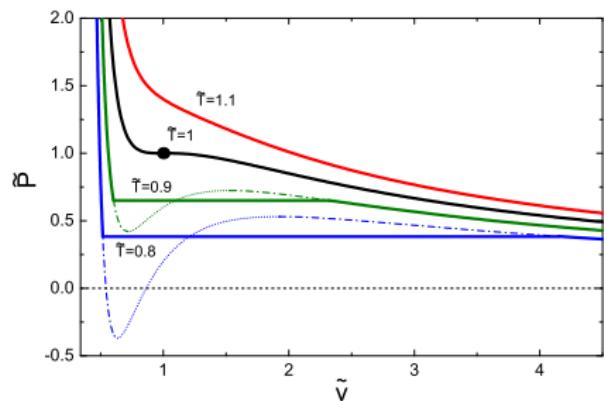
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Thanks for your attention!

# Backup slides

# Van der Waals equation

- VDW isotherms show irregular behavior below certain temperature  $T_C$
- Below  $T_C$  isotherms are corrected by Maxwell's rule of equal areas
- Results in appearance of mixed phase



Critical point

$$\frac{\partial p}{\partial v} = 0, \quad \frac{\partial^2 p}{\partial v^2} = 0, \quad v = V/N$$
$$p_C = \frac{a}{27b^2}, \quad n_C = \frac{1}{3b}, \quad T_C = \frac{8a}{27b}$$

Reduced variables

$$\tilde{p} = \frac{p}{p_C}, \quad \tilde{n} = \frac{n}{n_C}, \quad \tilde{T} = \frac{T}{T_C}$$

VDW equation originally formulated in **canonical ensemble**

How to transform **CE** pressure  $P(T, n)$  into **GCE** pressure  $P(T, \mu)$ ?

- Calculate  $\mu(T, V, N)$  from standard TD relations
- Invert the relation to get  $N(T, V, \mu)$  and put it back into  $P(T, V, N)$
- Consistency due to thermodynamic equivalence of ensembles

Result: transcendental equation for  $n(T, \mu)$

$$\frac{N}{V} \equiv n(T, \mu) = \frac{n_{\text{id}}(T, \mu^*)}{1 + b n_{\text{id}}(T, \mu^*)}, \quad \mu^* = \mu - b \frac{n T}{1 - b n} + 2 a n$$

- Implicit equation in GCE, in CE it was explicit
- May have multiple solutions below  $T_C$
- Choose one with largest pressure – equivalent to Maxwell rule in CE

**Advantages** of the GCE formulation

- 1) **Hadronic** physics applications: number of hadrons usually **not conserved**.
- 2) **CE** cannot describe particle number **fluctuations**. N-fluctuations in a small ( $V \ll V_0$ ) subsystem follow **GCE** results.
- 3) Good starting point to include effects of **quantum statistics**.

# Scaled variance in VDW equation

New application from GCE formulation: **particle number fluctuations**

Scaled variance is an **intensive** measure of N-fluctuations

$$\frac{\sigma^2}{N} = \omega[N] \equiv \frac{\langle N^2 \rangle - \langle N \rangle^2}{\langle N \rangle} = \frac{T}{n} \left( \frac{\partial n}{\partial \mu} \right)_T = \frac{T}{n} \left( \frac{\partial^2 P}{\partial \mu^2} \right)_T$$

In **ideal** Boltzmann gas fluctuations are Poissonian and  $\omega_{id}[N] = 1$ .

$\omega[N]$  in VDW gas (pure phases)

$$\omega[N] = \left[ \frac{1}{(1 - bn)^2} - \frac{2an}{T} \right]^{-1}$$

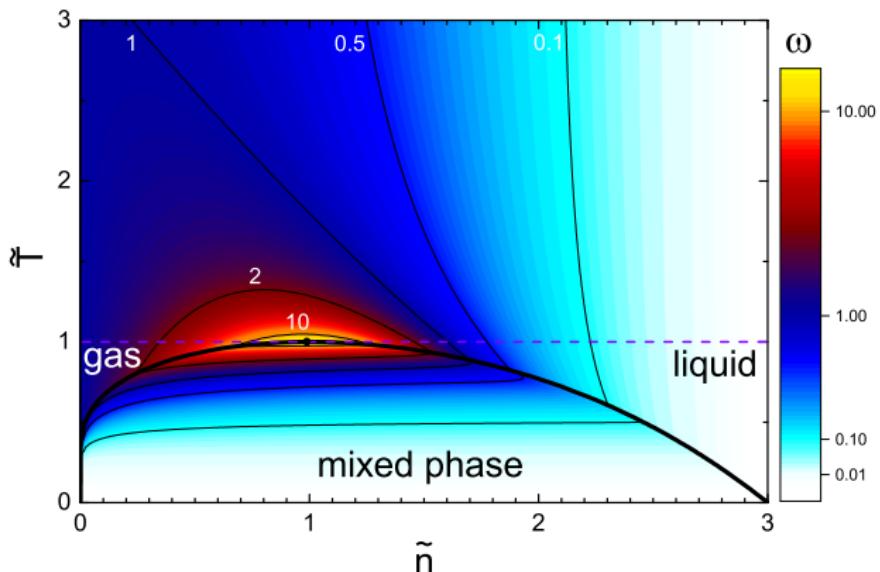
- Repulsive interactions **suppress** N-fluctuations
- Attractive interactions **enhance** N-fluctuations

N-fluctuations are useful because they

- Carry information about finer details of EoS, e.g. **phase transitions**
- Measurable **experimentally**

# Scaled variance

$$\omega[N] = \frac{1}{9} \left[ \frac{1}{(3 - \tilde{n})^2} - \frac{\tilde{n}}{4\tilde{T}} \right]^{-1}$$



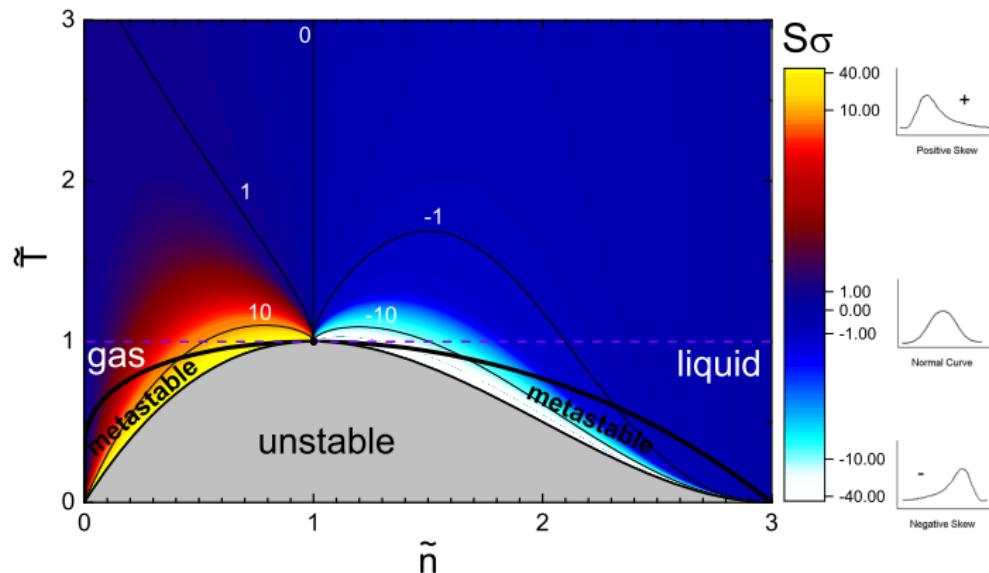
- Deviations from unity signal effects of interaction
- Fluctuations grow rapidly near critical point

V. Vovchenko et al., J. Phys. A 305001, 48 (2015)

# Skewness

Higher-order (non-gaussian) fluctuations are even more sensitive

Skewness:  $S\sigma = \frac{\langle(\Delta N)^3\rangle}{\sigma^2} = \omega[N] + \frac{T}{\omega[N]} \left( \frac{\partial \omega[N]}{\partial \mu} \right)_T$  asymmetry

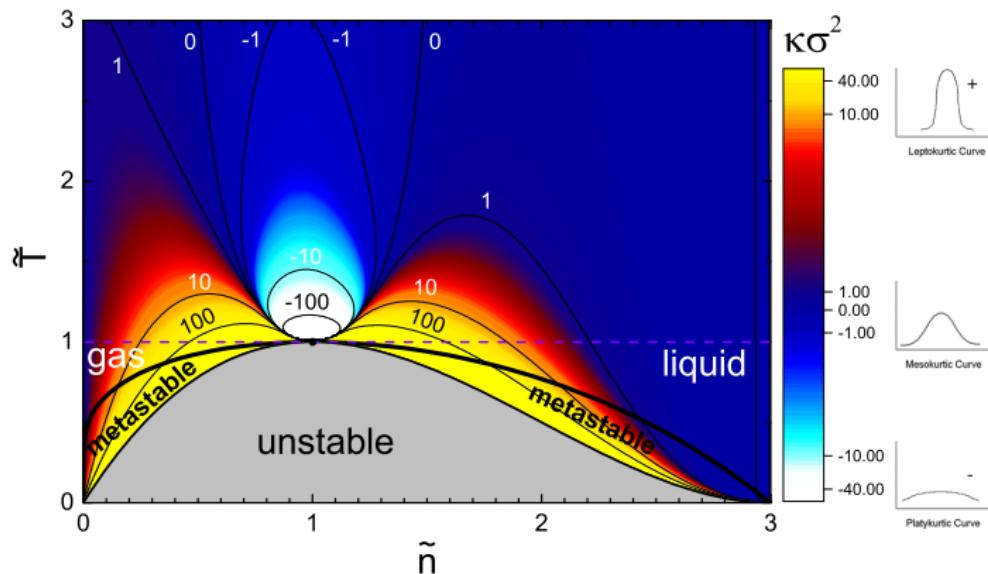


Skewness is

- Positive (right-tailed) in gaseous phase
- Negative (left-tailed) in liquid phase

# Kurtosis

$$\text{Kurtosis: } \kappa\sigma^2 = \frac{\langle(\Delta N)^4\rangle - 3\langle(\Delta N)^2\rangle^2}{\sigma^2} \quad \text{peakedness}$$



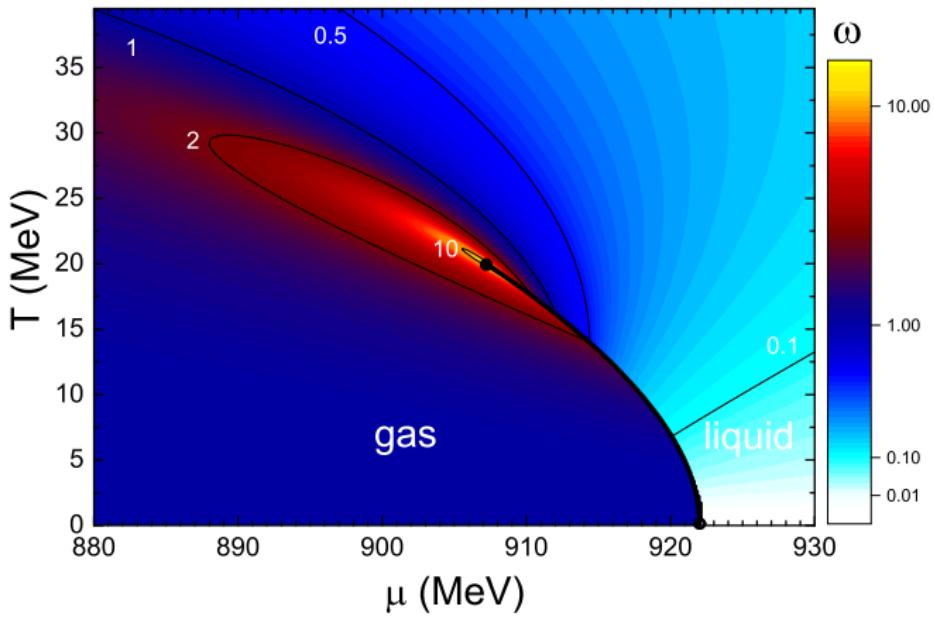
Kurtosis is **negative** (flat) above critical point (crossover), **positive** (peaked) elsewhere and very **sensitive** to the **proximity** of the critical point

V. Vovchenko et al., J. Phys. A 015003, 49 (2016)

# VDW gas of nucleons: scaled variance

Scaled variance in quantum VDW:

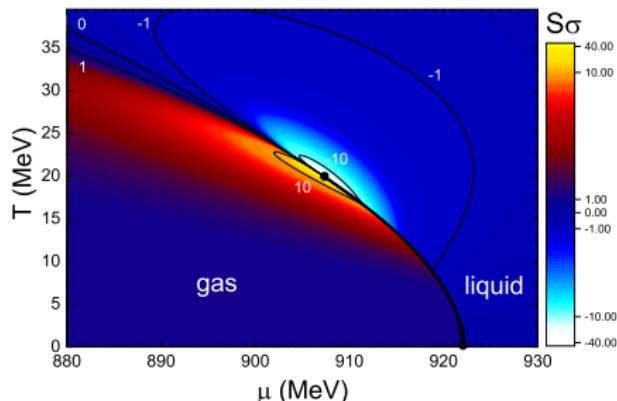
$$\omega[N] = \omega_{\text{id}}(T, \mu^*) \left[ \frac{1}{(1 - bn)^2} - \frac{2an}{T} \omega_{\text{id}}(T, \mu^*) \right]^{-1}$$



# VDW gas of nucleons: skewness and kurtosis

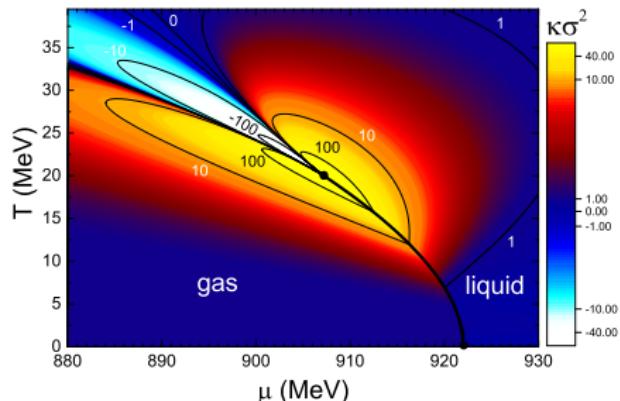
Skewness

$$S\sigma = \omega[N] + \frac{T}{\omega[N]} \left( \frac{\partial \omega[N]}{\partial \mu} \right)_T$$



Kurtosis

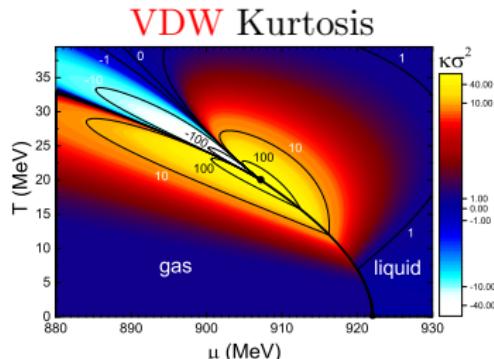
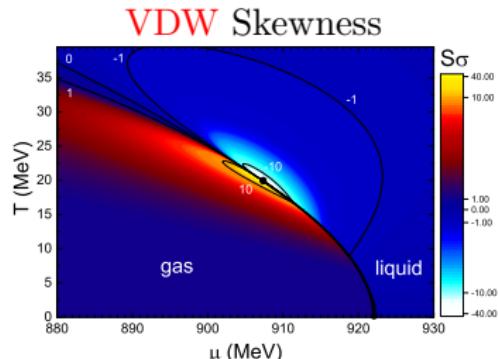
$$\kappa\sigma^2 = (S\sigma)^2 + T \left( \frac{\partial [S\sigma]}{\partial \mu} \right)_T$$



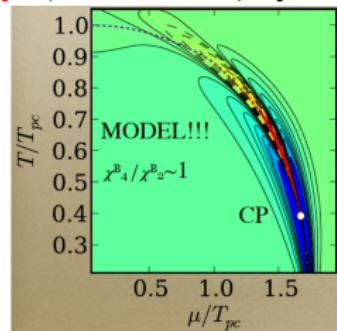
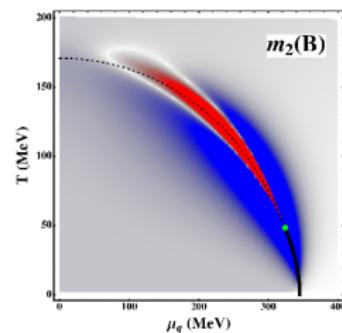
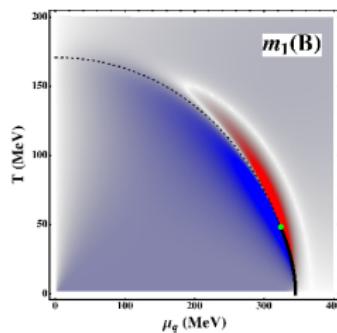
For skewness and kurtosis singularity is rather specific: sign depends on the path of approach

V. Vovchenko et al., Phys. Rev. C 92, 054901 (2015)

# VDW gas of nucleons: skewness and kurtosis



NJL, J.W. Chen et al., PRD 93, 034037 (2016) PQM, V. Skokov, QM2012



Fluctuation patterns in VDW very similar to effective QCD models

# VDW-HRG at $\mu = 0$ : baryon number fluctuations

$$\chi_6^B/\chi_2^B$$

