



Statistical-thermal model: Applications using Thermal-FIST

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3rd EMMI Workshop: Anti-matter, hyper-matter and exotica production at the LHC

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FIAS Frankfurt Institute
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1. Short description of **Thermal-FIST**
2. Recent applications to light nuclei and exotica
 - Canonical suppression
 - The Saha equation approach
 - Feeddown contributions from excited nuclear states
3. Summary

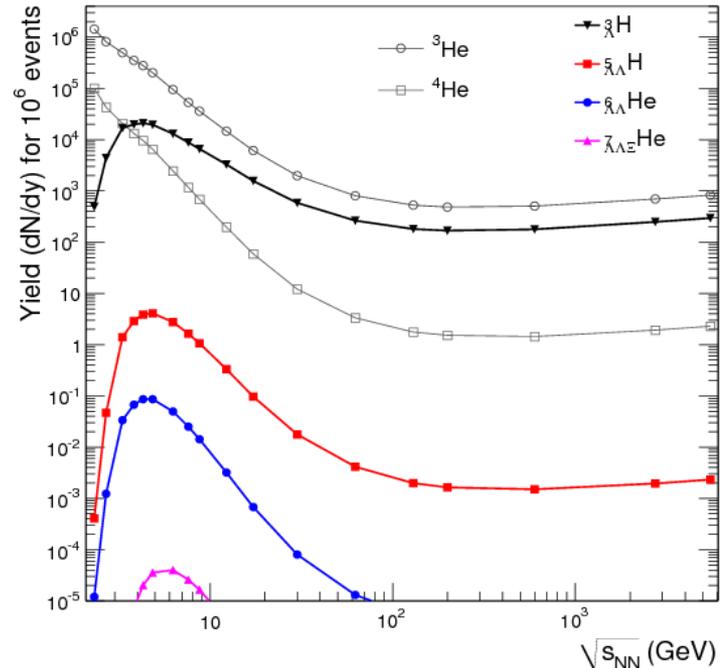
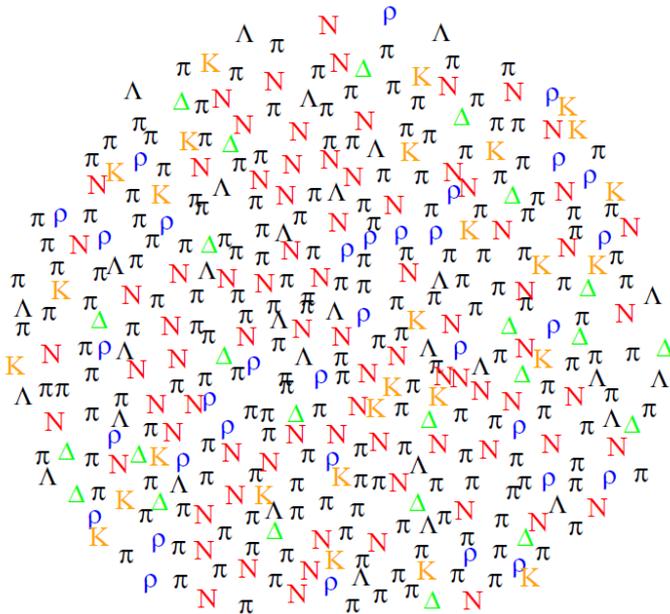
Hadron resonance gas (HRG) at freeze-out



HRG: Equation of state of hadronic matter as a multi-component (non-)interacting gas of known hadrons, resonances, and *light nuclei*

$$\ln Z \approx \sum_{i \in M, B} \ln Z_i^{id} = \sum_{i \in M, B} \frac{d_i V}{2\pi^2} \int_0^\infty \pm p^2 dp \ln \left[1 \pm \exp \left(\frac{\mu_i - E_i}{T} \right) \right]$$

Grand-canonical ensemble: $\mu_i = b_i \mu_B + q_i \mu_Q + s_i \mu_S$ *chemical equilibrium*



Thermal-FIST



Thermal-FIST* (a.k.a. FIST or FAUST)

C++/Qt/Jupyter

open source: <https://github.com/vlvovch/Thermal-FIST>

reference: V.V., H. Stoecker, *Computer Physics Communications* **244**, 295 (2019)

Thermal-FIST 1.0

File View Help

Particle list file: C:/FIST/PDG2014/list-withnuclei.dat

Thermal model Thermal fits Equation of state Event generator Particle list editor

Data to fit: Hint: double-click on yield to edit

Name	Fit?	Exp. value	Exp. error	Model value	Deviation	Data/Model	Feeddown
1 pi+	<input checked="" type="checkbox"/>	669.5	48	605.439	-1.33461	1.106 ± 0.079	Strong+EM decays
2 pi-	<input checked="" type="checkbox"/>	668	47	605.474	-1.33034	1.103 ± 0.078	Strong+EM decays
3 K+	<input checked="" type="checkbox"/>	100	8	108.722	1.09024	0.920 ± 0.074	Strong+EM decays
4 K-	<input checked="" type="checkbox"/>	99.5	8.5	108.657	1.07724	0.916 ± 0.078	Strong+EM decays
5 p	<input checked="" type="checkbox"/>	31.5	2.5	33.4123	0.764917	0.943 ± 0.075	Strong+EM decays
6 anti-p	<input checked="" type="checkbox"/>	30.5	2.5	33.2773	1.11091	0.917 ± 0.075	Strong+EM decays
7 Lambda	<input checked="" type="checkbox"/>	24	2.5	19.3002	-1.87991	1.244 ± 0.130	Strong+EM decays

HRG model configuration:

Model: Ideal Ensemble: Grand-canonical

Statistics: Boltzmann Quantum for All particles Use quadratures

Resonance widths: eBW

Conservation laws... EV/vdW interactions... Other options...

Fit parameters:

Parameter	Fit?	Initial value	Min value	Max value
T (MeV)	<input checked="" type="checkbox"/>	155	20	500
R (fm)	<input checked="" type="checkbox"/>	8	0	25
uB (MeV)	<input checked="" type="checkbox"/>	0	-100	900

Extracted parameters:

Parameter	Value	Error
T (MeV)	154.766	1.19547
uB (MeV)	0.323424	3.94532
Yq	1	--
Ys	1	--
R (fm)	10.5875	0.263019
V (fm ³)	4971.27	370.495
chi2/dof	26.0616/19	

Plots: Yields Deviations Data/M

Equation of state... Chi2 profile...

Thermal fit result

Data/Model

A framework for general-purpose statistical-thermal model applications

**Thermal, Fast and Interactive Statistical Toolkit*

Using Thermal-FIST



The package is *cross-platform* (Linux, Mac, Windows, Android)

Installation using **git** and **cmake**

```
# Clone the repository from GitHub
git clone https://github.com/vlvovch/Thermal-FIST.git
cd Thermal-FIST

# Create a build directory, configure the project with cmake
# and build with make
mkdir build
cd build
cmake ../
make

# Run the GUI frontend
./bin/QtThermalFIST

# Run the test calculations from the paper
./bin/examples/cpc1HRGTDep
./bin/examples/cpc2chi2
./bin/examples/cpc3chi2NEQ
./bin/examples/cpc4mcHRG
```

GUI requires free [Qt5 framework](#), the rest of the package has *no external dependencies*

[Quick start guide](#)

[Documentation](#)

[Physics manual](#)

Statistical-thermal model aspects in FIST



- Extensions of the HRG model
 - finite resonance widths
 - repulsive (excluded volume) and van der Waals (*criticality*) interactions
 - particle number fluctuations and correlations
 - chemical non-equilibrium (γ_q, γ_s) a la Rafelski
 - unstable nuclear fragments
- Equation of state
- Canonical statistical model (CSM)
 - (local) (selective) exact conservation of conserved charges
 - canonical suppression of light nuclei
- Monte Carlo generator (Blast-wave, canonical ensemble,...)
- Hadronic phase and dynamical freeze-out
 - partial chemical equilibrium
 - suppression of resonance yields
 - evolution of light nuclei abundances via the Saha equation

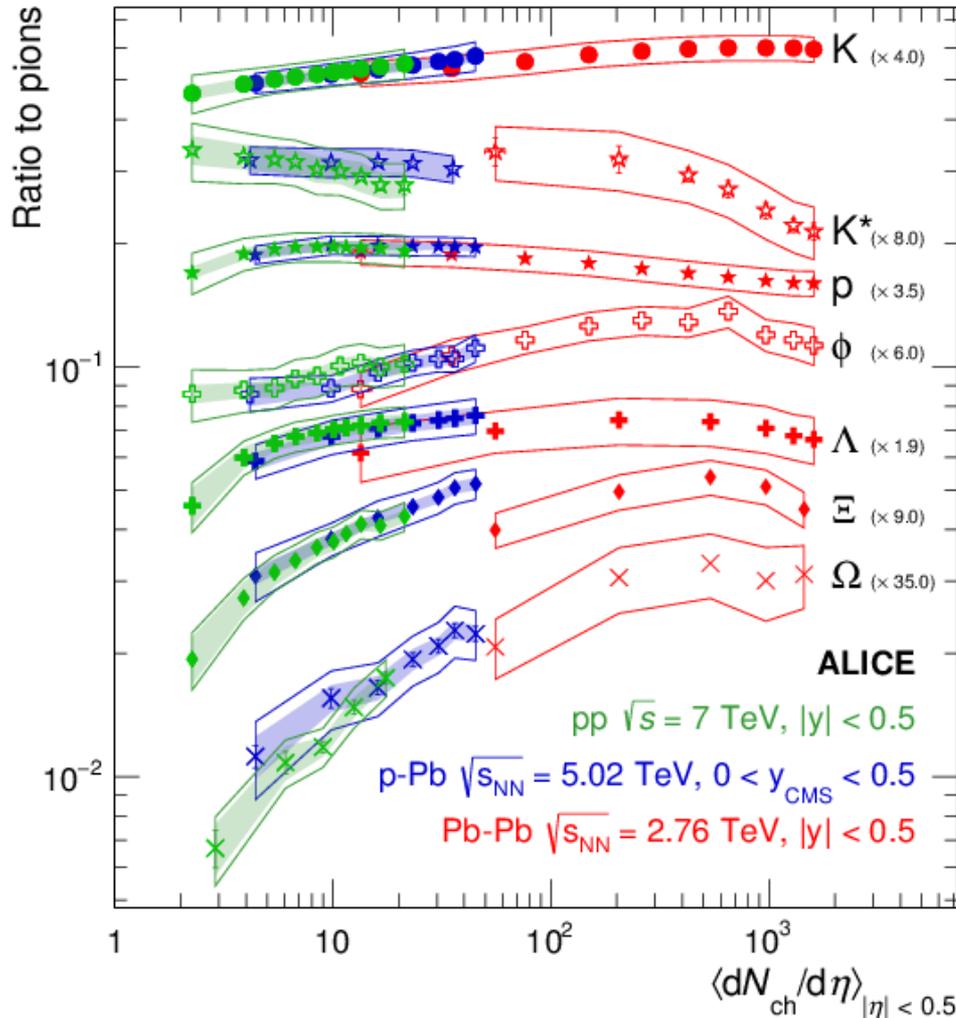
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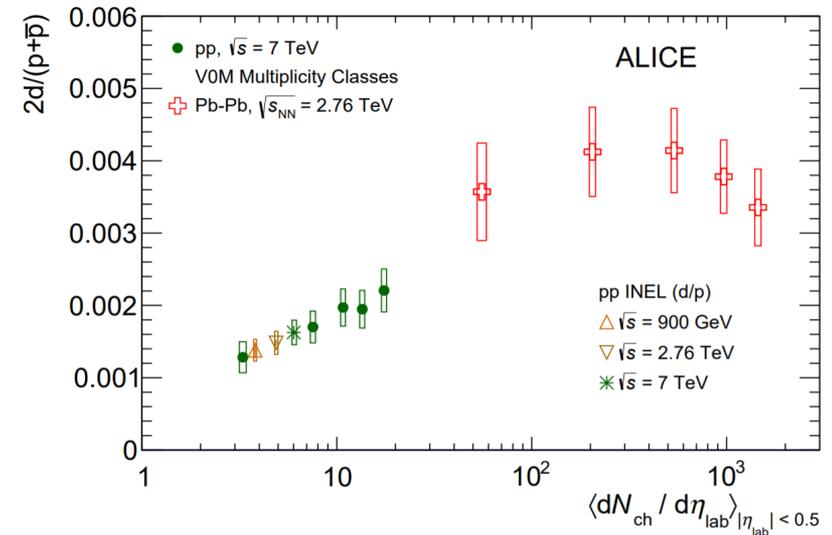
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Canonical suppression of light nuclei at the LHC

Multiplicity dependence of hadrochemistry



[ALICE collaboration, 1807.11321]



[ALICE collaboration, 1902.09290]

- Grand-canonical thermal picture predicts **no multiplicity dependence**
- Apply **canonical** statistical model?

Canonical statistical model (CSM)



Exact conservation of B , Q , S in a correlation volume V_C

[Rafelski, Danos, et al., PLB '80; Hagedorn, Redlich, ZPC '85]

$$Z(B, Q, S) = \int_{-\pi}^{\pi} \frac{d\phi_B}{2\pi} \int_{-\pi}^{\pi} \frac{d\phi_Q}{2\pi} \int_{-\pi}^{\pi} \frac{d\phi_S}{2\pi} e^{-i(B\phi_B + Q\phi_Q + S\phi_S)} \exp \left[\sum_j z_j^1 e^{i(B_j\phi_B + Q_j\phi_Q + S_j\phi_S)} \right]$$

$$z_j^1 = V_C \int dm \rho_j(m) d_j \frac{m^2 T}{2\pi^2} K_2(m/T)$$

$$\langle N_j^{\text{prim}} \rangle^{\text{ce}} = \frac{Z(B - B_j, Q - Q_j, S - S_j)}{Z(B, Q, S)} \langle N_j^{\text{prim}} \rangle^{\text{gce}}$$

[Becattini et al., ZPC '95, ZPC '97]

Implemented in Thermal-FIST for a full HRG

Canonical statistical model (CSM)



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Net-proton fluctuations affected by baryon number conservation

[Braun-Munzinger, Rustamov, Stachel, 1612.00702]

$$\frac{\kappa_2(\mathbf{p} - \bar{\mathbf{p}})}{\langle \mathbf{p} \rangle + \langle \bar{\mathbf{p}} \rangle} \simeq 1 - \frac{\langle \mathbf{p} \rangle}{k dN_B/dy}$$

Using ALICE data for net-p fluctuations [1910.14396] one obtains $k \sim 3-4$ for most centrality bins in Pb-Pb collisions

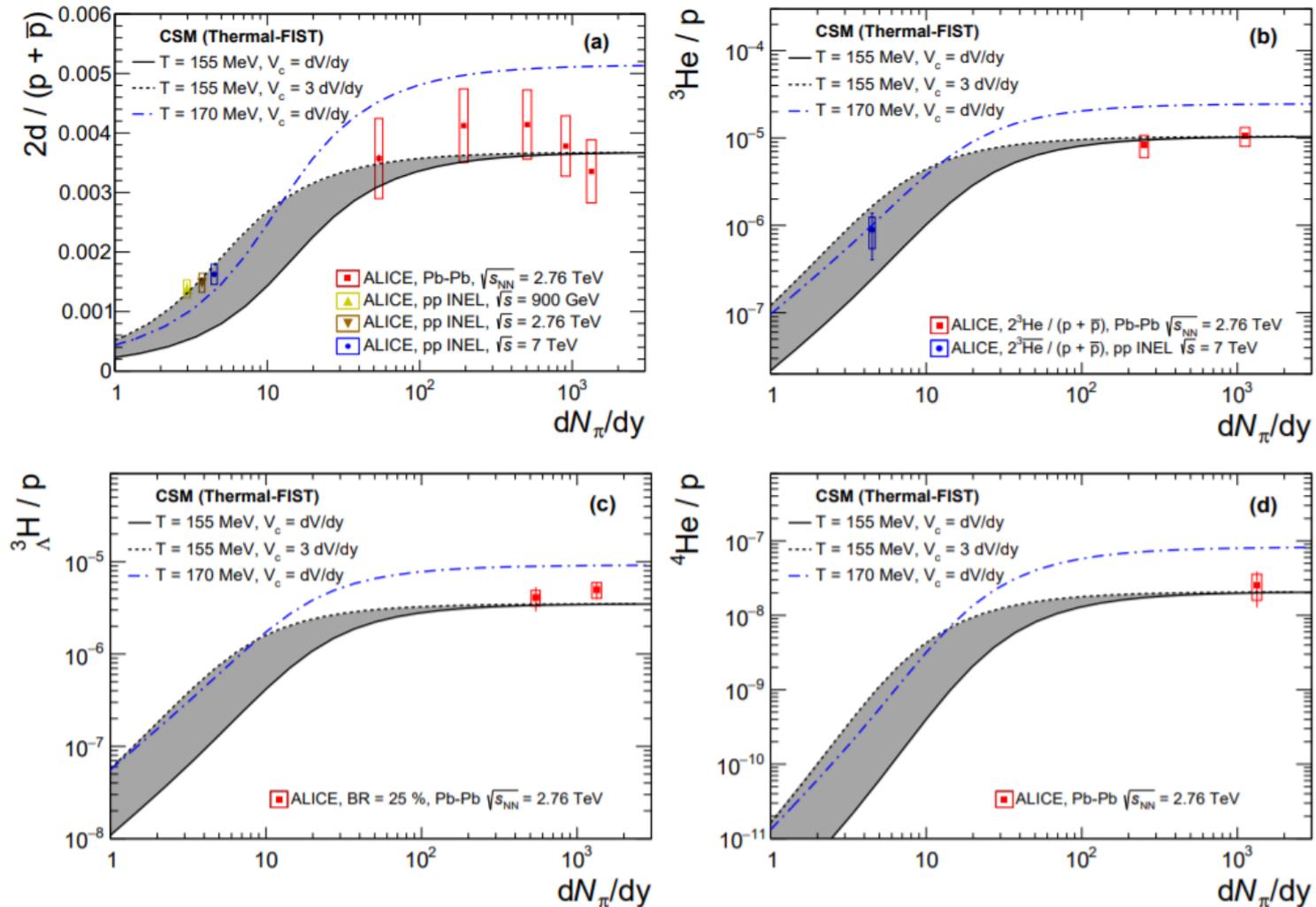
[V.V., Dönigus, Stoecker, 1906.03145, PRC '19]

“Vanilla” CSM



$T_{ch} = 155$ MeV, $V_C = 3dV/dy$, multiplicity dependence driven by V_C only

[V.V., Dönigus, Stoecker, 1808.05245, PLB '18]

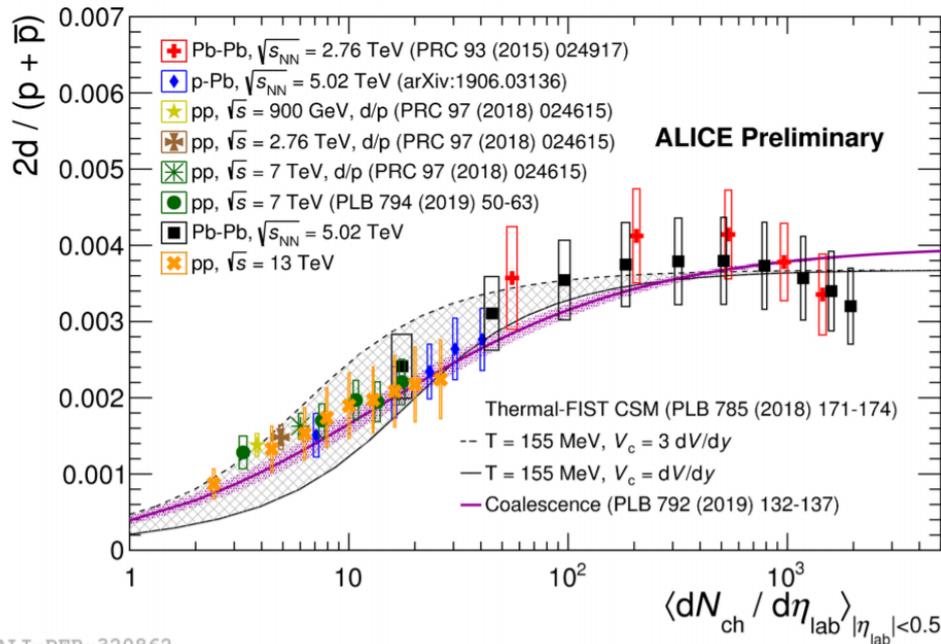


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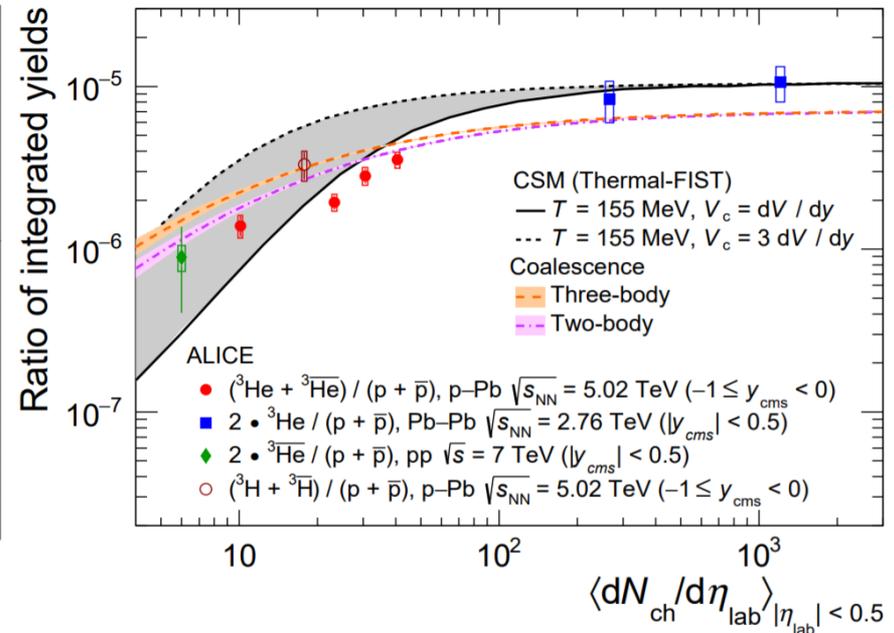
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ALI-DER-320862

[L. Barioglio, QM19]



[ALICE collaboration, 1910.14401]

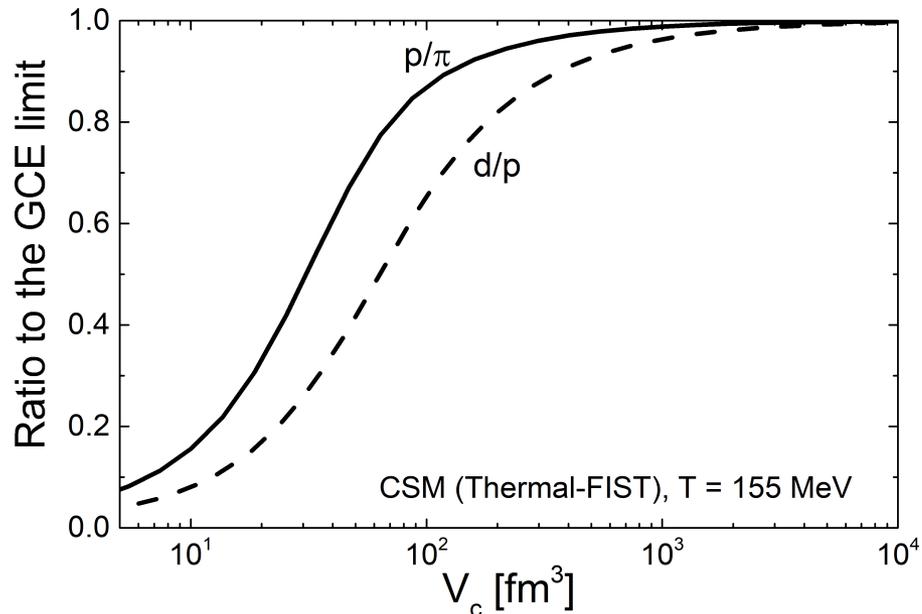
Basic CSM appears to capture trends seen in light nuclei production data

“Vanilla” CSM: nuclei vs p/π ratio



Canonical suppression affects not only nuclei, but also the p/π ratio

The effect for p/π is generally milder than d/p , but not insignificant

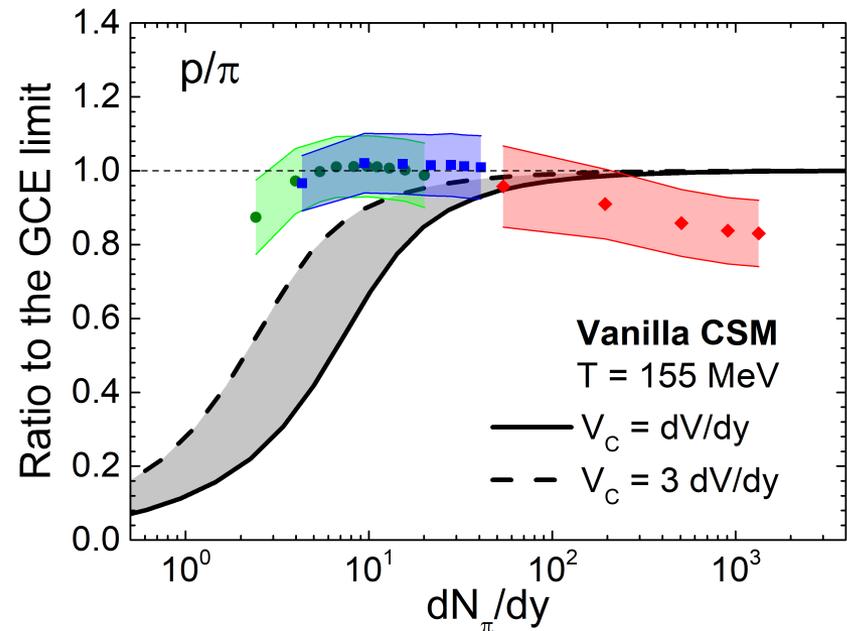
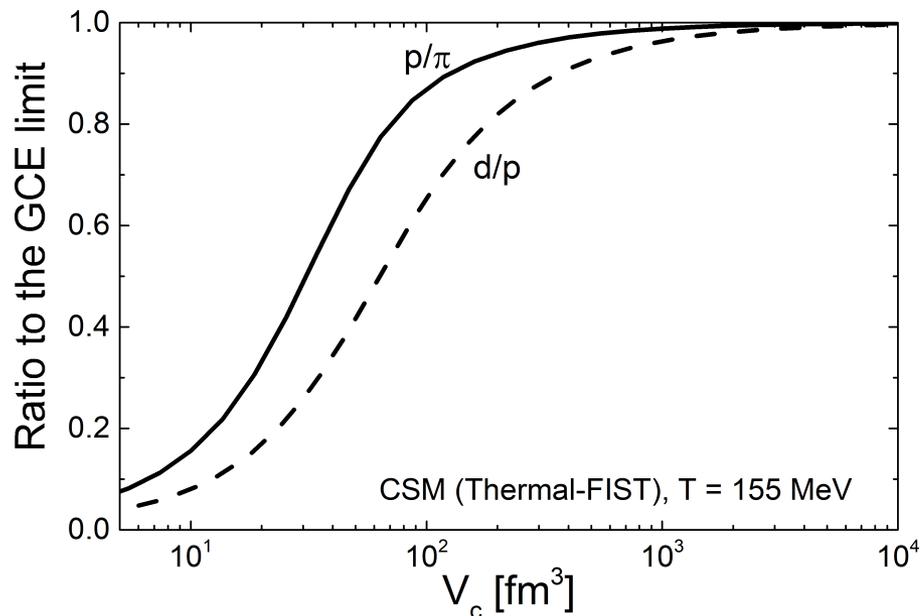


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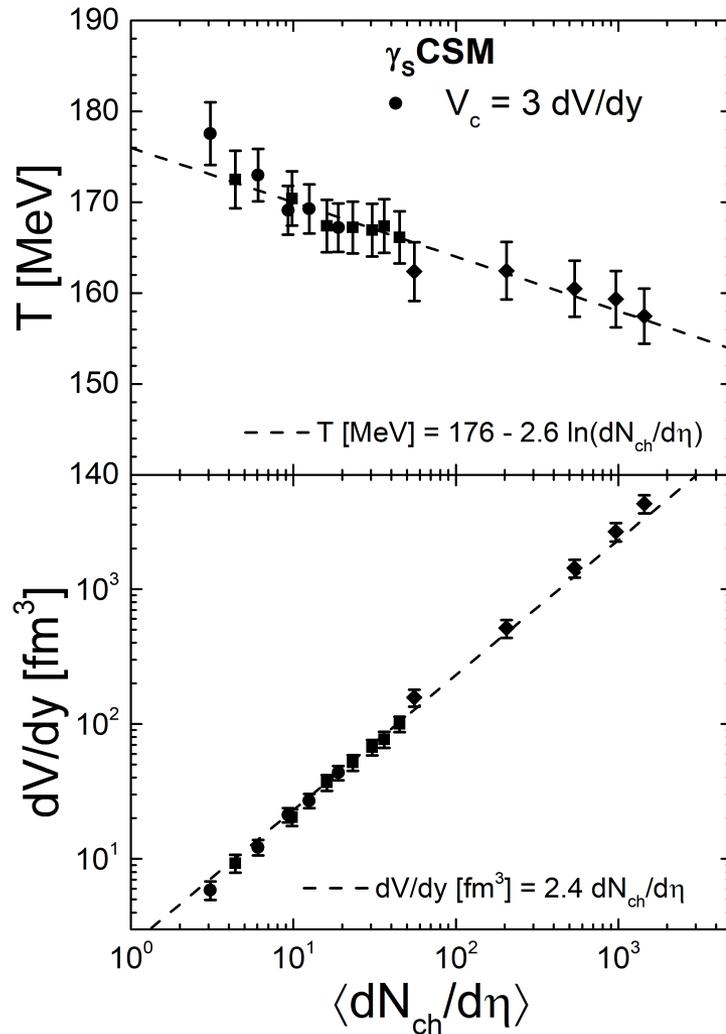


p/π suppression predicted by vanilla CSM not supported by the data

Full CSM



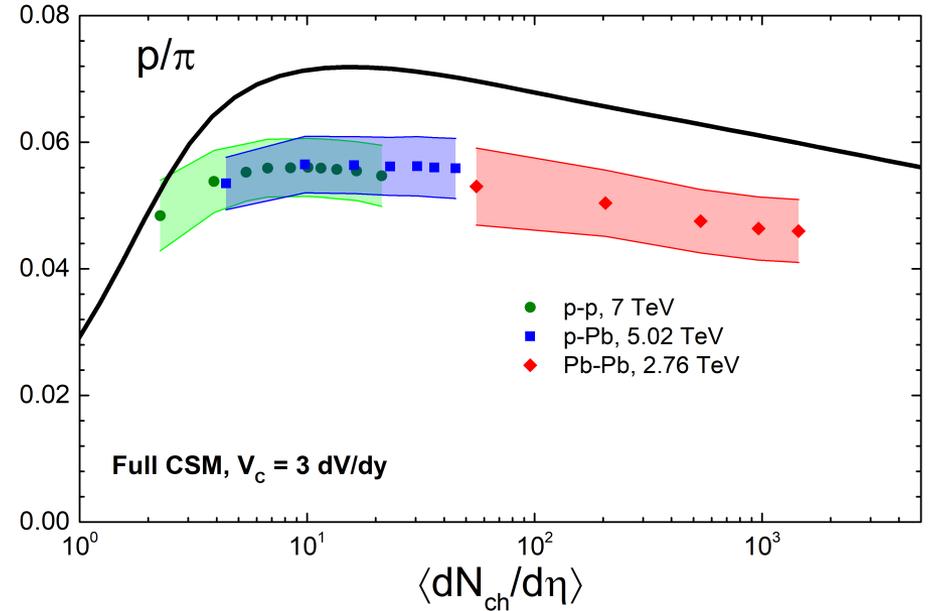
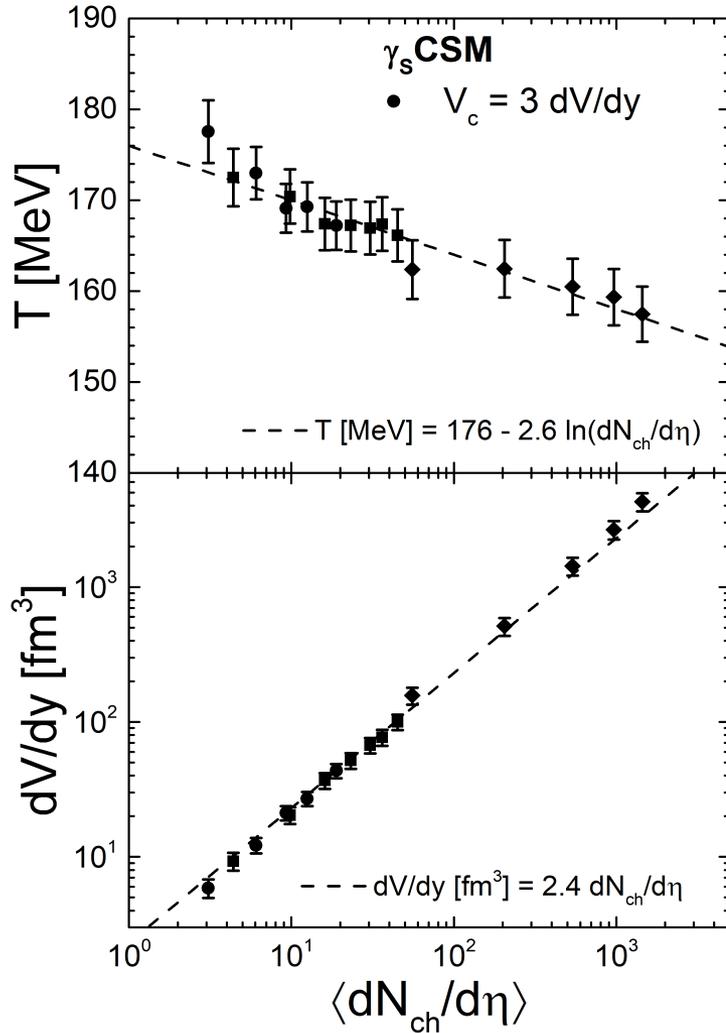
Full CSM: allow for **multiplicity-dependent** T_{ch} [V.V., Dönigus, Stoecker, 1906.03145, PRC '19]



Full CSM



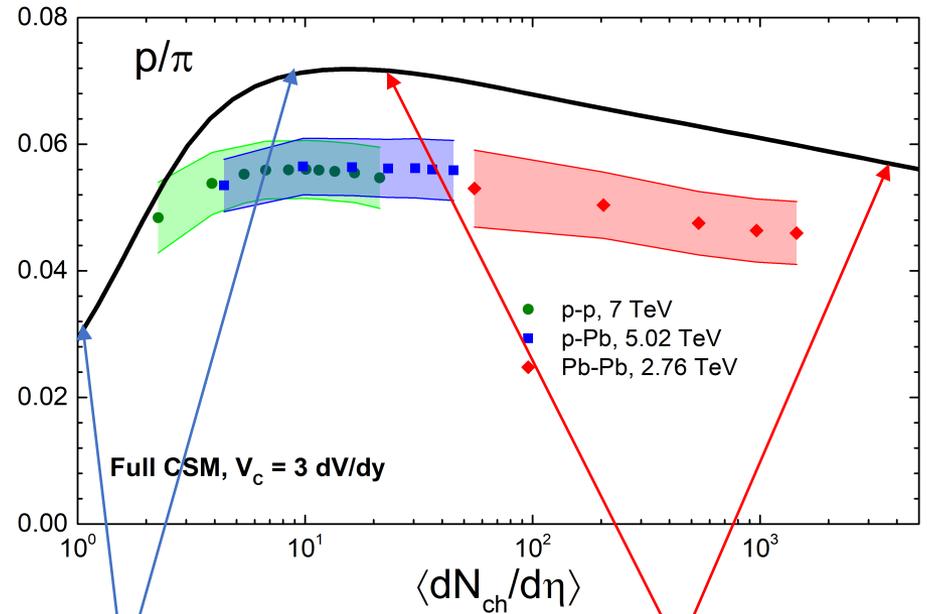
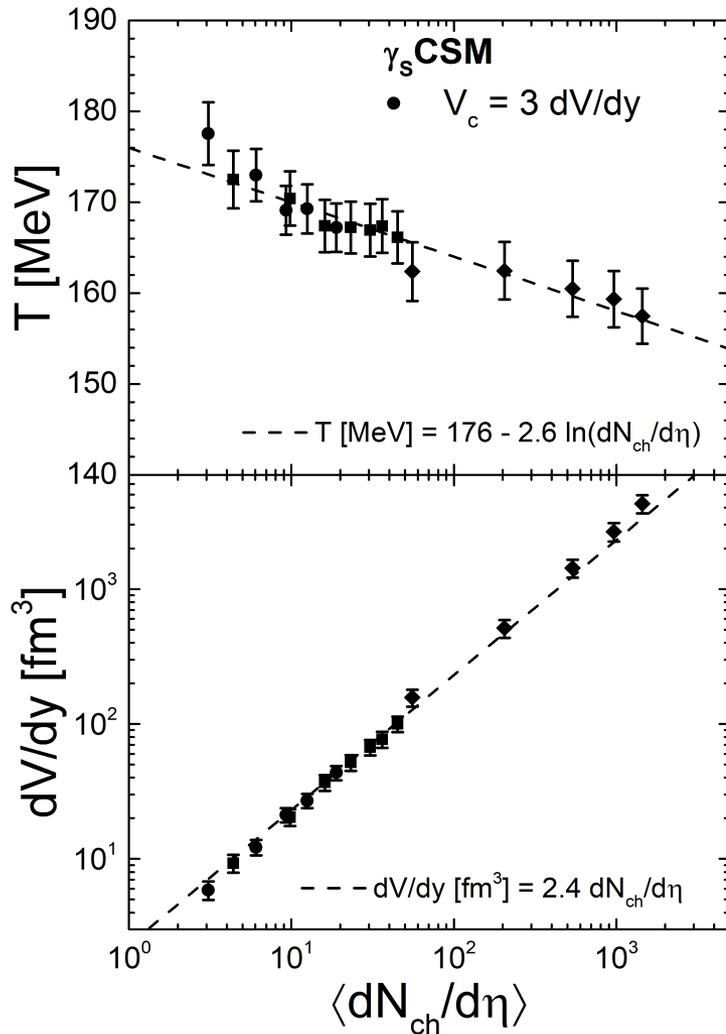
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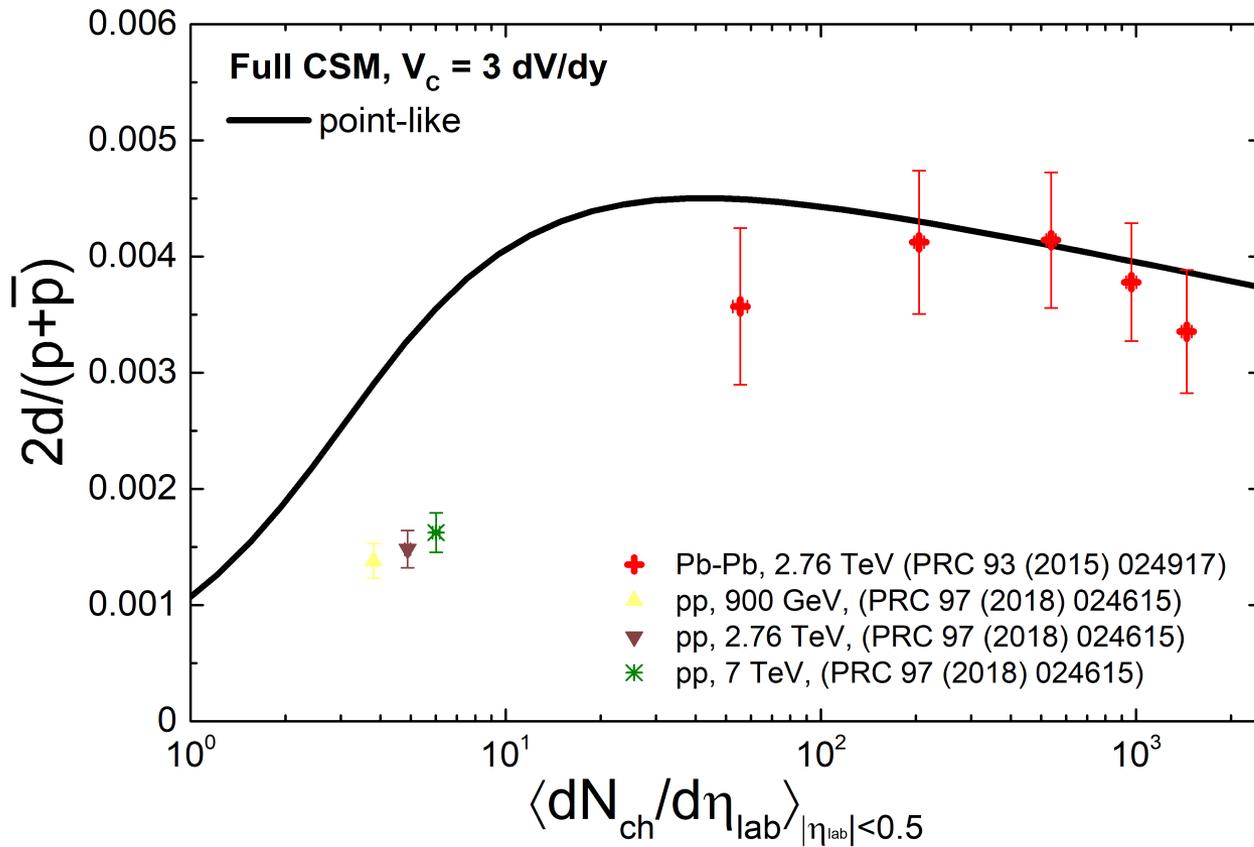
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$V_c \searrow \Rightarrow p/\pi \searrow$

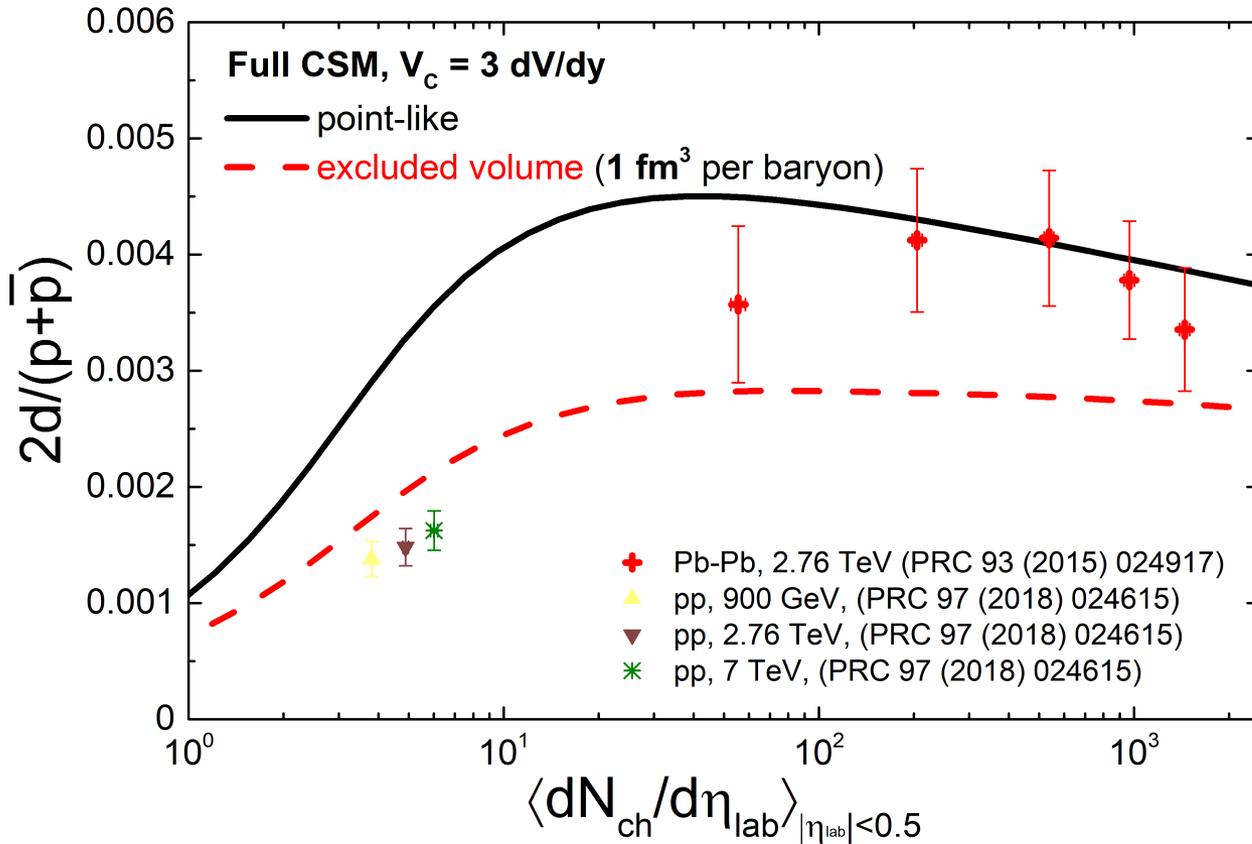
$T_{ch} \nearrow \Rightarrow p/\pi \nearrow$

Full CSM: d/p



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Full CSM: d/p



$T_{ch} \nearrow \Rightarrow d/p \nearrow$

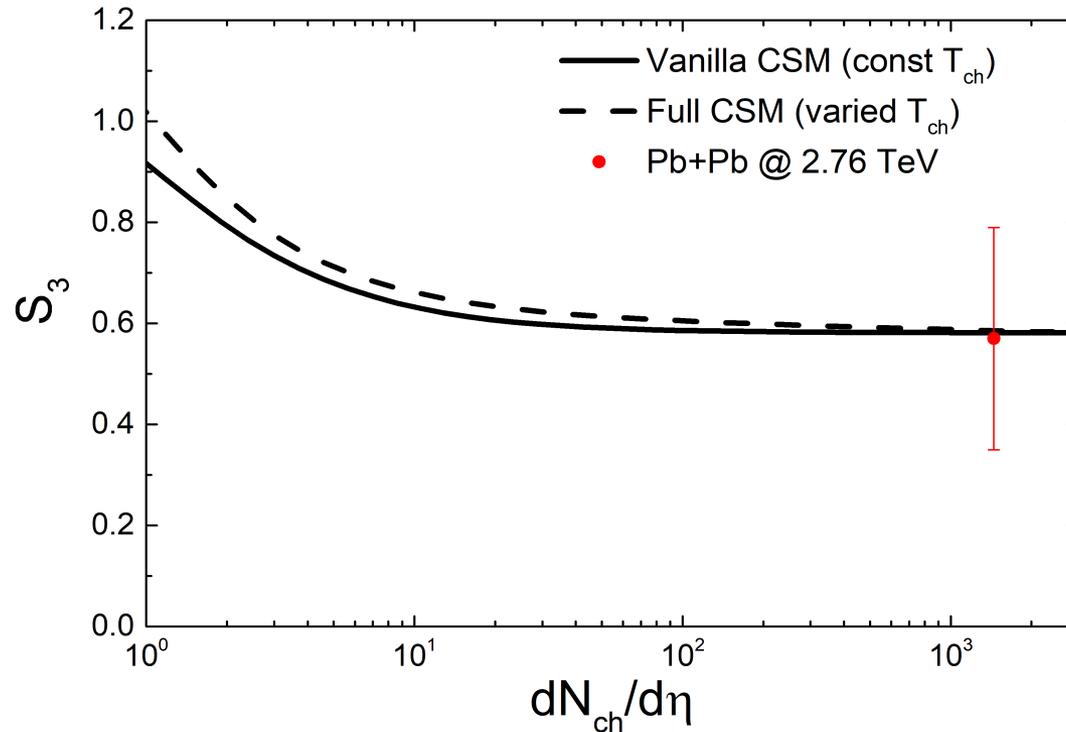
Excluded volume (schematic): $N_i \rightarrow N_i \exp\left(-\frac{v_i p}{T}\right) \Rightarrow d/p \searrow$

Simultaneous description of light nuclei and p/π ratio remains challenging



$$S_3 = \binom{3}{\Lambda} \text{H}/\text{He} / (\Lambda/p)$$

[E864 collab., PRC '04; Zhang et al., PLB '10]

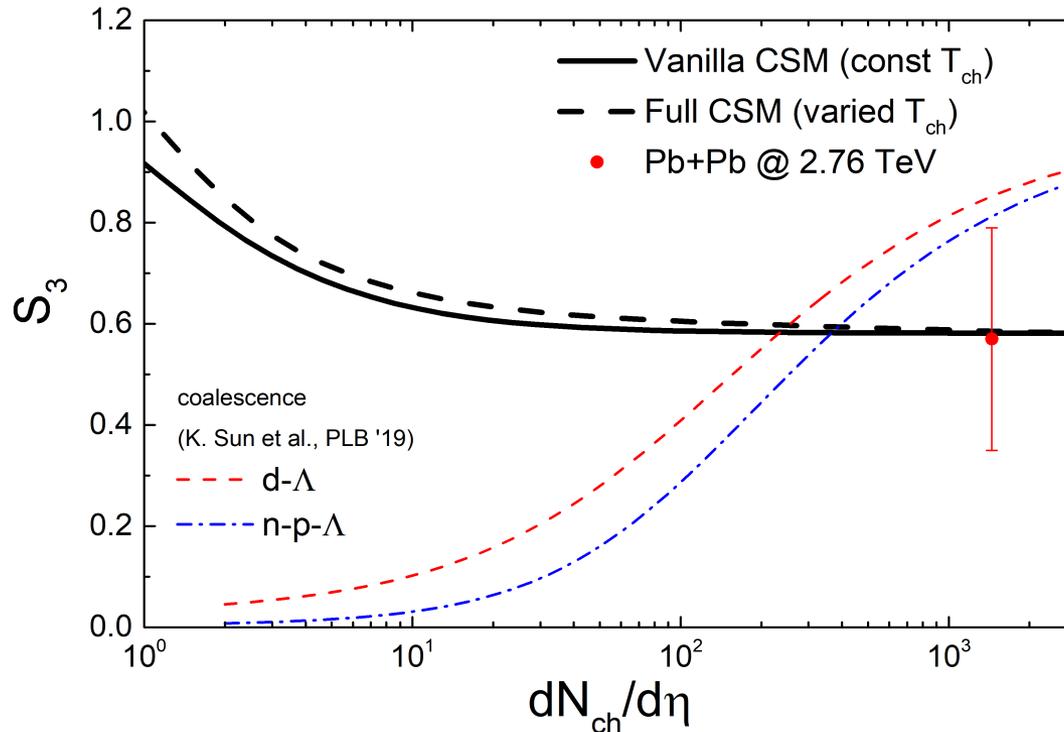


Different versions of CSM give similar predictions, mild increase of S_3 due to baryon and strangeness conservation



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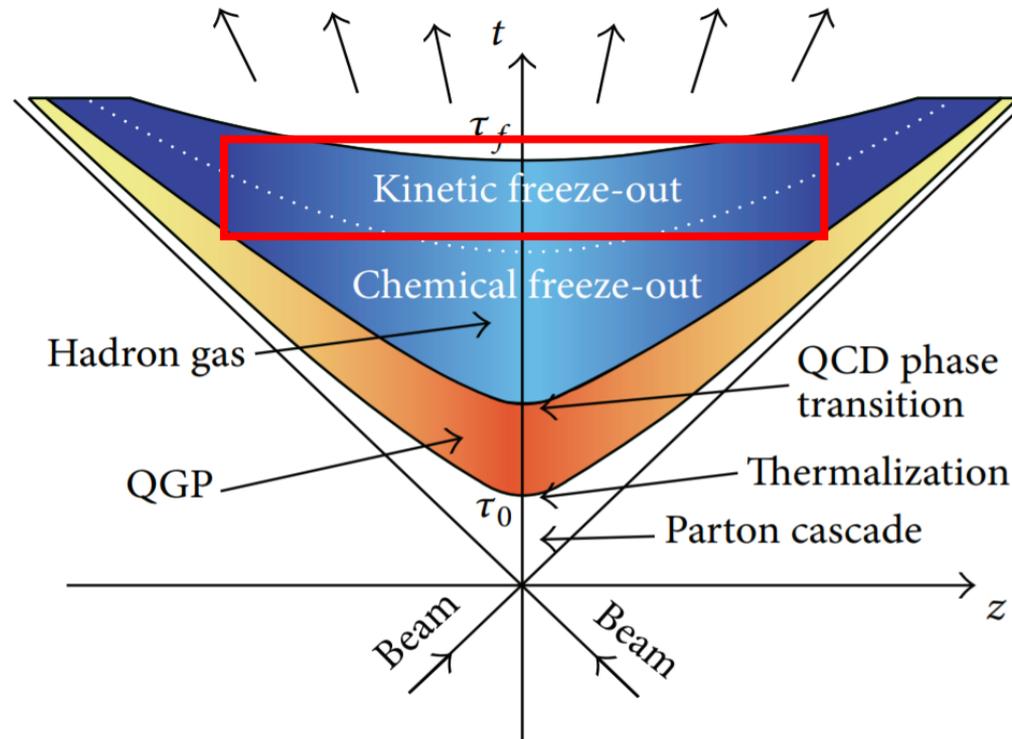
Different versions of CSM give similar predictions, mild increase of S_3 due to baryon and strangeness conservation

Coalescence [Sun, Dönigus, Ko, PLB '19] predicts **opposite trend**

Hadronic phase and the Saha equation approach to light nuclei production

V.V., *K. Gallmeister, J. Schaffner-Bielich, C. Greiner, [1903.10024](#), PLB (in print)*

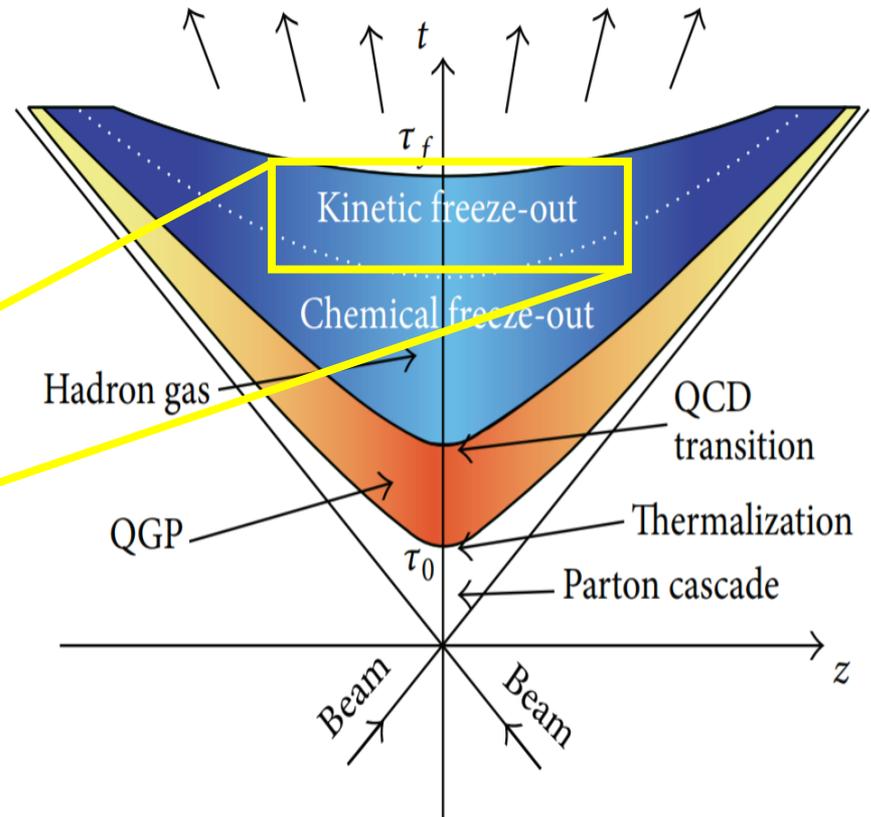
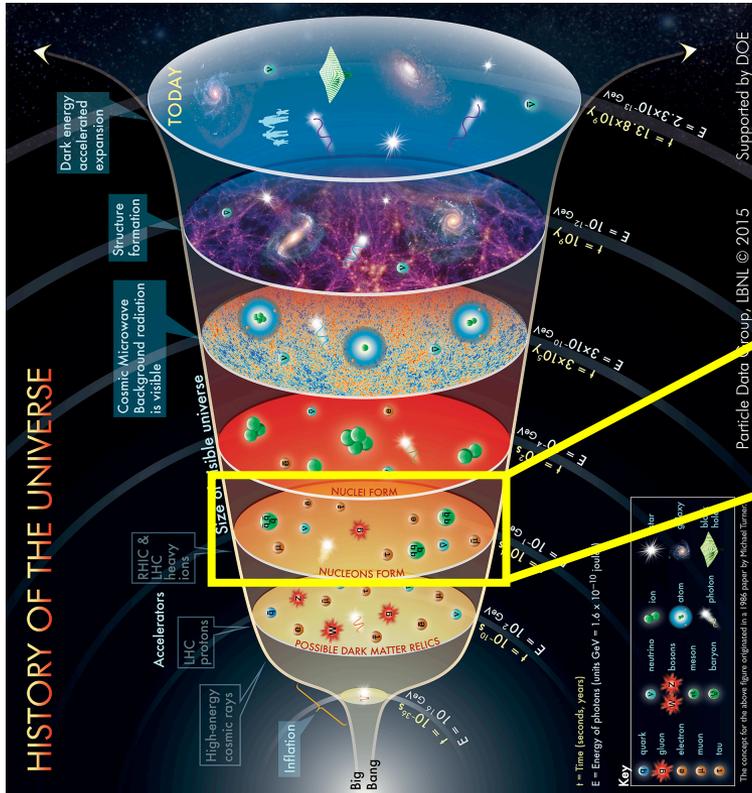
Hadronic phase in central HICs



- At $T_{ch} \approx 150 - 160 \text{ MeV}$ inelastic collisions cease, yields of hadrons frozen
- Kinetic equilibrium maintained down to $T_{kin} \approx 100 - 120 \text{ MeV}$ through (pseudo-)elastic scatterings

[e.g., E. Shuryak, Rev. Mod. Phys. 89, 035001 (2017)]

Big Bang vs “Little Bangs”



- Hadrons (nucleons) form and “freeze-out” chemically before nuclei
- Bosons (photons or pions) catalyse nucleosynthesis

e.g. $p + n \leftrightarrow d + \gamma$ vs $p + n + \pi \leftrightarrow d + \pi$

Saha equation: $\mu_A = \sum_i \mu_{A_i}$, e.g. $\mu_d = \mu_p + \mu_n$, $\mu_{3\text{He}} = 2\mu_p + \mu_n$, ...

Partial chemical equilibrium (PCE)



Expansion of hadron resonance gas in partial chemical equilibrium at $T < T_{ch}$

[H. Bebie, P. Gerber, J.L. Goity, H. Leutwyler, Nucl. Phys. B '92; C.M. Hung, E. Shuryak, PRC '98]

Chemical composition of stable hadrons is fixed, kinetic equilibrium maintained through pseudo-elastic resonance reactions $\pi\pi \leftrightarrow \rho$, $\pi K \leftrightarrow K^*$, $\pi N \leftrightarrow \Delta$, etc.

Effective chemical potentials:

$$\tilde{\mu}_j = \sum_{i \in \text{stable}} \langle n_i \rangle_j \mu_i, \quad \langle n_i \rangle_j - \text{mean number of hadron } i \text{ from decays of hadron } j, \quad j \in \text{HRG}$$

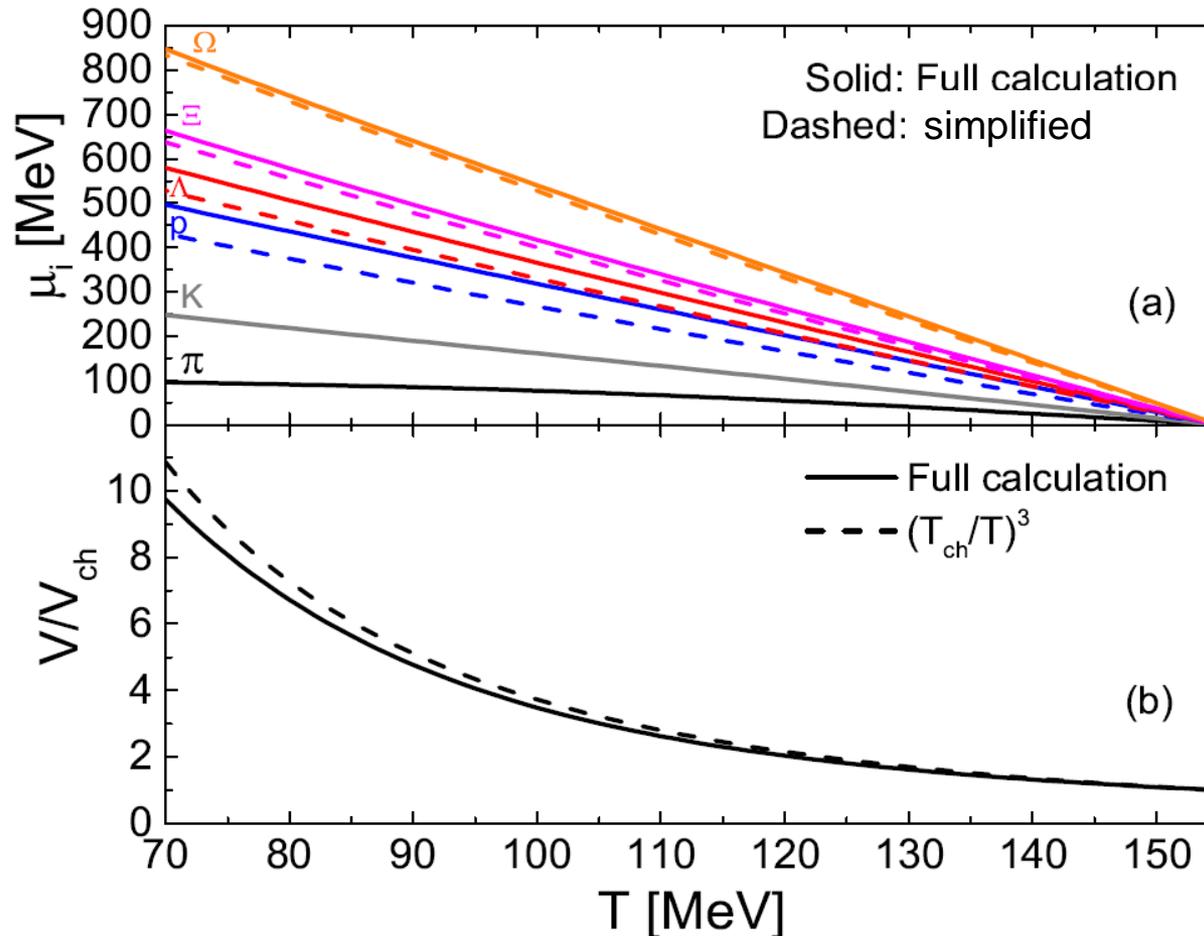
Conservation laws:

$$\sum_{j \in \text{hrg}} \langle n_i \rangle_j n_j(T, \tilde{\mu}_j) V = N_i(T_{ch}), \quad i \in \text{stable} \quad \text{numerical solution}$$
$$\sum_{j \in \text{hrg}} s_j(T, \tilde{\mu}_j) V = S(T_{ch}) \quad \{\mu_i(T)\}, V(T)$$

E.g.: $\pi + 2\rho + 3\omega + \dots = \text{const}$, $N + \Delta + N^* + \dots = \text{const}$, $K + K^* + \dots = \text{const}$

Numerical implementation of PCE in **Thermal-FIST**

Partial chemical equilibrium at the LHC



[V.V., K. Gallmeister, J. Schaffner-Bielich, C. Greiner, 1903.10024]

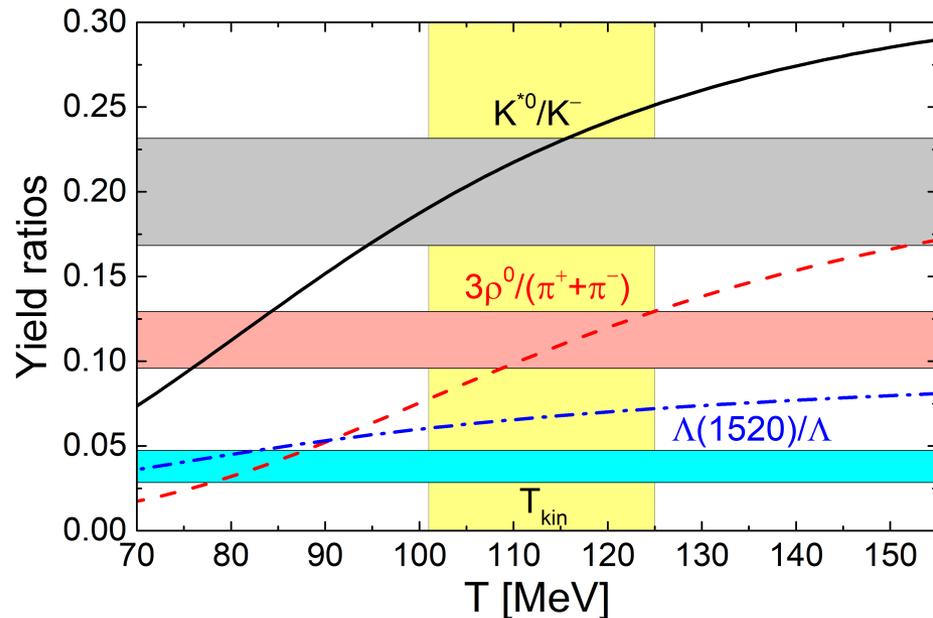
“Initial conditions”: $T_{ch} = 155$ MeV, $V_{ch} = 4700$ fm³ (0-10% Pb-Pb 2.76 TeV)

Resonance suppression in hadronic phase



Yields of **resonances** are *not* conserved in partial chemical equilibrium

E.g. K^* yield dilutes during the cooling through reactions $\pi K \leftrightarrow K^*$



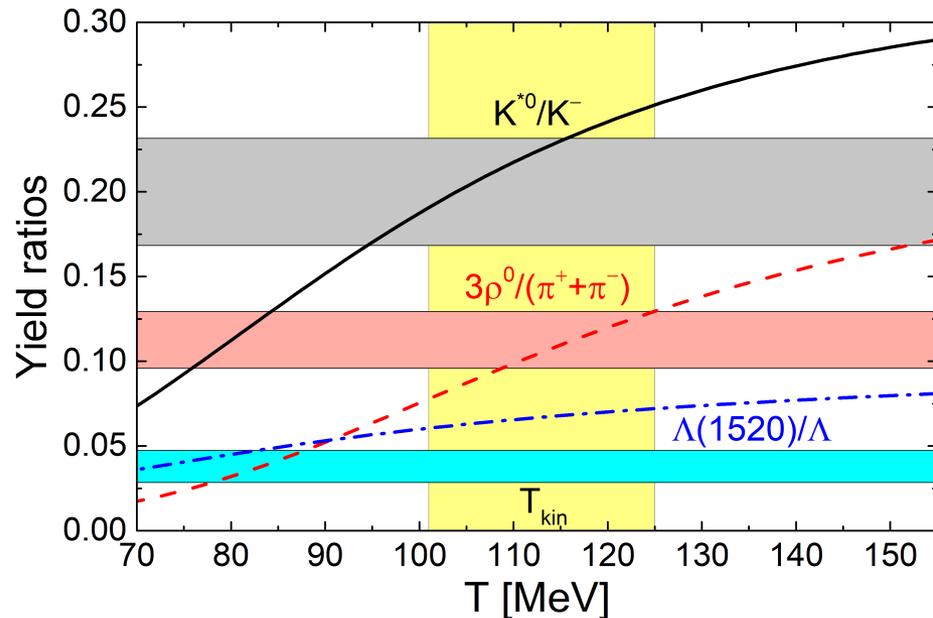
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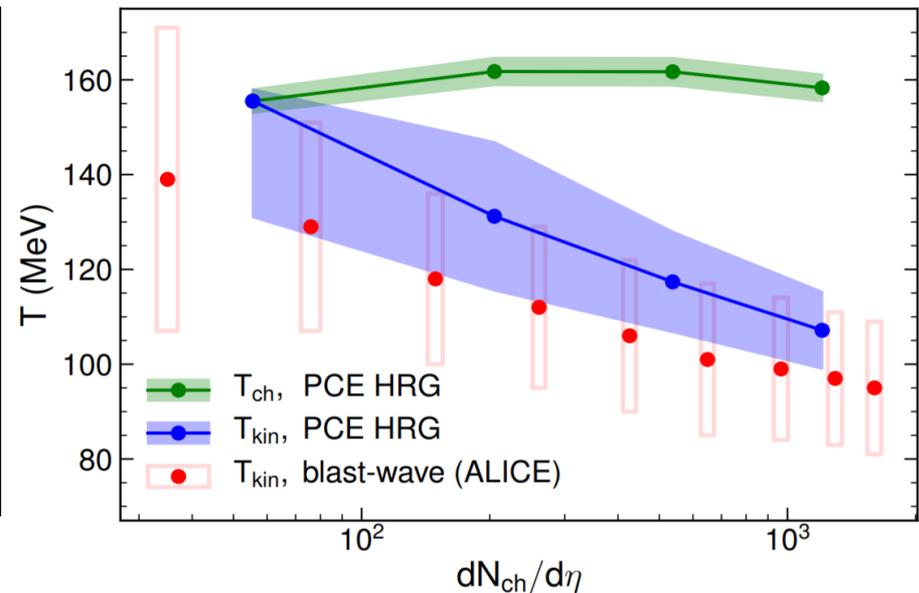


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[Motornenko, V.V., Greiner, Stoecker, 1908.11730]

Fitting the yields of **short-lived resonances** is a new way to extract the **kinetic freeze-out temperature**

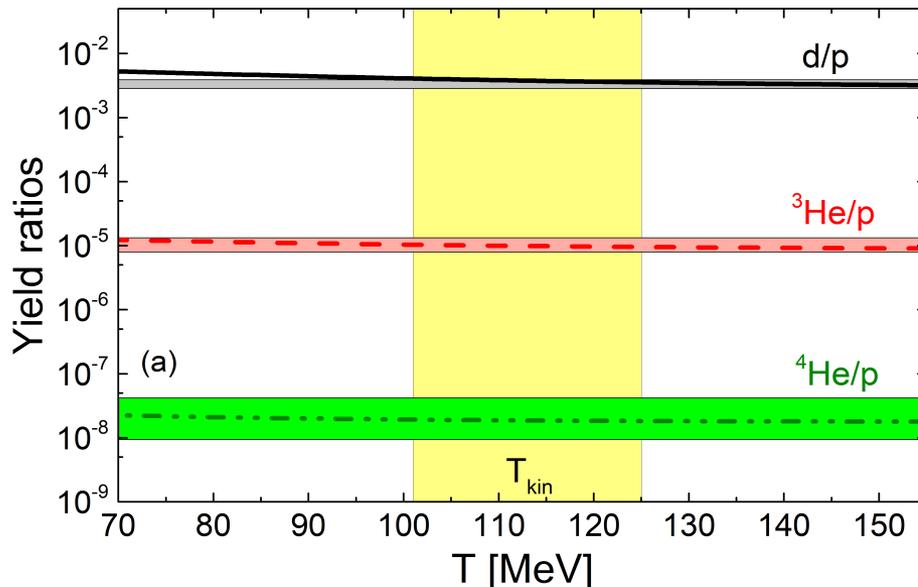
Saha equation: Light nuclei



Detailed balance for nuclear reactions, $X + A \leftrightarrow X + \sum_i A_i$, X is e.g. a pion

$$\frac{n_A}{\prod_i n_{A_i}} = \frac{n_A^{\text{eq}}}{\prod_i n_{A_i}^{\text{eq}}}, \quad \Leftrightarrow \quad \mu_A = \sum_i \mu_{A_i}, \quad \text{e.g. } \mu_d = \mu_p + \mu_n, \quad \mu_{^3\text{He}} = 2\mu_p + \mu_n, \quad \dots$$

Saha equation



Deviations from thermal model predictions are moderate despite significant cooling and dilution. *Is this the reason for why thermal model works so well?*

For $T = T_{\text{kin}}$ similar results reported in [X. Xu, R. Rapp, EPJA 55, 68 (2019)]

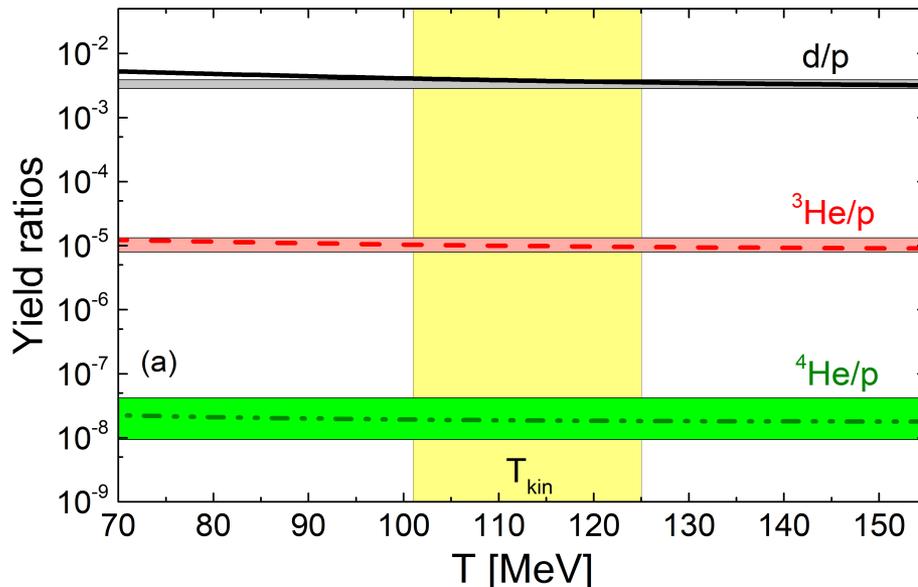
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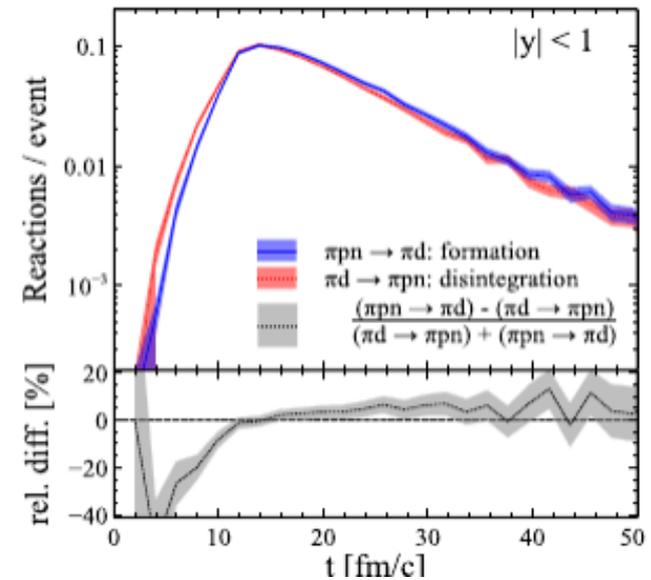
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Saha equation



Transport [Oliinychenko et al., 1809.03071]

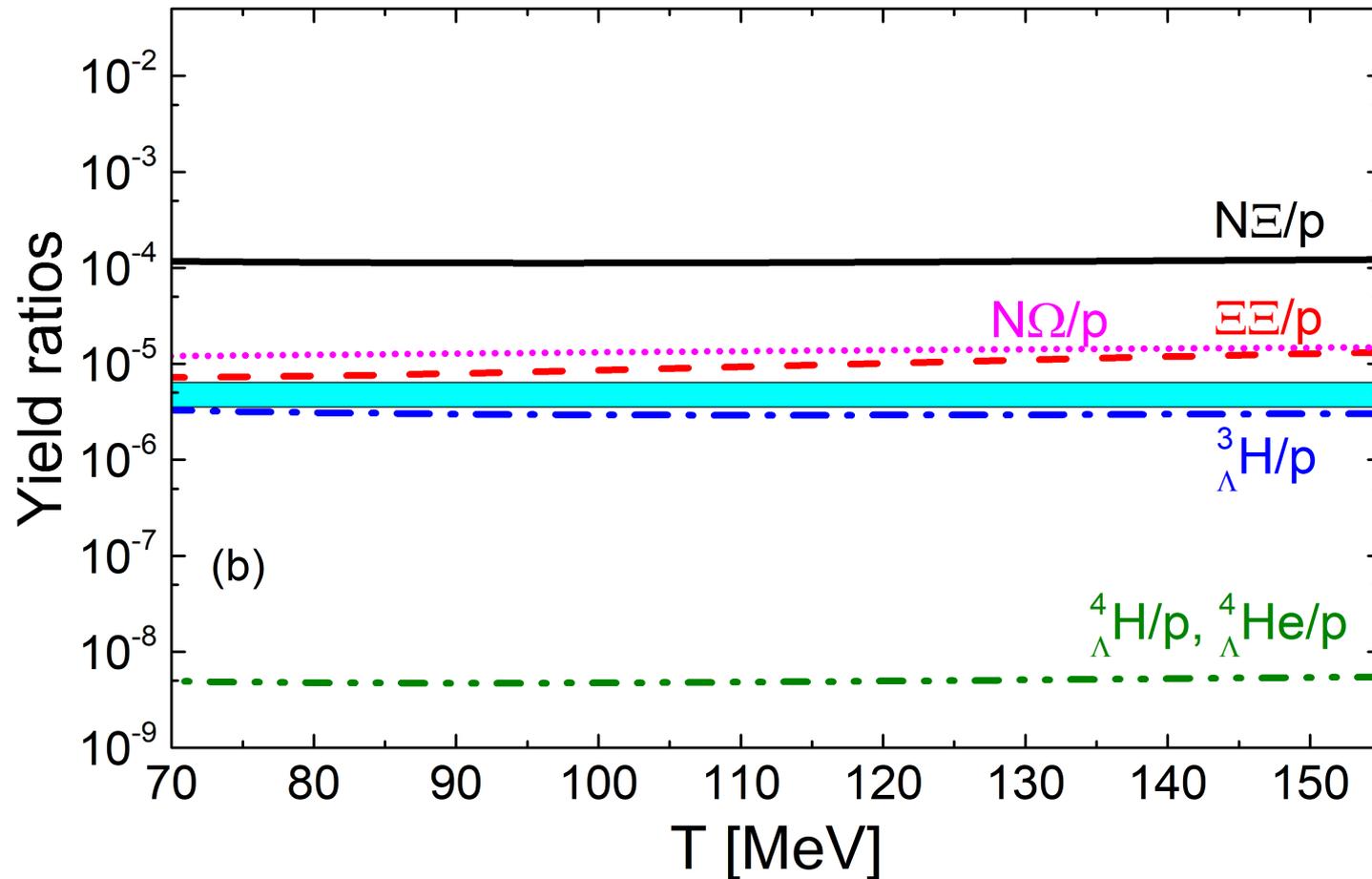


▪ **Law of mass action at work**

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Saha equation: Hypernuclei



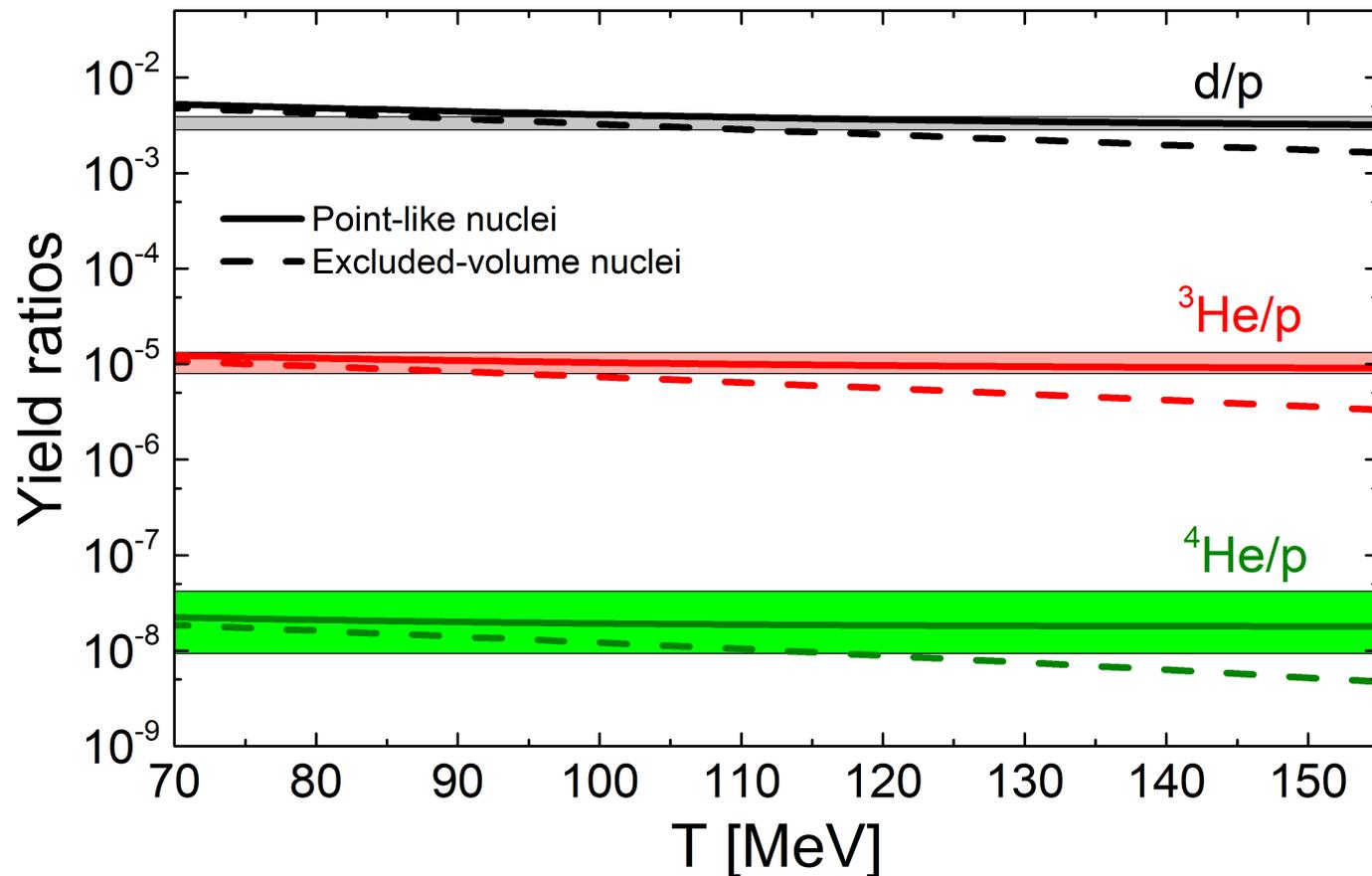
Hypernuclei stay close to the thermal model prediction. An exception is a hypothetical $\Xi\Xi$ state ← *planned measurement in Runs 3 & 4 at the LHC*

[LHC Yellow Report, 1812.06772]

Saha equation and excluded volume effects



Eigen volumes: effective mechanism for nuclei suppression at large densities



Excluded-volume effects go away as the system dilutes. At $T \cong 100$ MeV agrees with the point-particle model. Does not describe data for $T = T_{ch}$

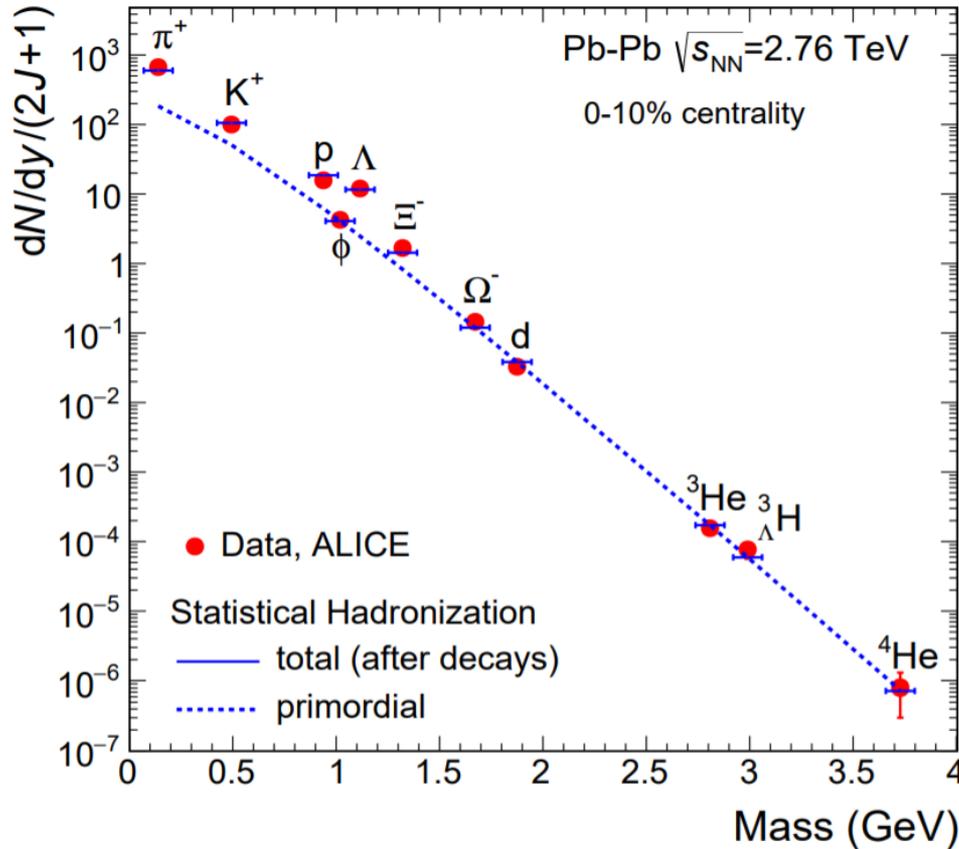
On the feeddown contributions from decays of unstable fragments

V.V., B. Dönigus, H. Stoecker, et al., in preparation

Feeddown in thermal model



Is well-known to be important for hadron yields



[Andronic et al., Nature (2018)]

Production of p

Primordial density = $0.0028648 \text{ fm}^{-3}$

$T = 155 \text{ MeV}$

Primordial yield = 11.4594

Total yield = 31.4347

Primordial + strong decays = 31.4347

Primordial + strong + EM decays = 31.4347

Primordial + strong + EM + weak decays = 48.5857

Source	Multiplicity	Fraction (%)
Primordial	11.4594	36.4545
Decays from primordial Delta(1232)++	4.86466	15.4755
Decays from primordial Delta(1232)+	3.24327	10.3175
Decays from primordial Delta(1232)0	1.62139	5.15797
Decays from primordial N(1520)0	0.5628	1.79038
Decays from primordial Delta(1600)++	0.540859	1.72058
Decays from primordial N(1520)+	0.436374	1.38819
Decays from primordial N(1440)0	0.412215	1.31134
Decays from primordial Delta(1600)+	0.3931	1.25053
Decays from primordial N(1440)+	0.367071	1.16773
Decays from primordial N(1675)+	0.362324	1.15263
Decays from primordial N(1680)0	0.352206	1.12044

[V.V., Stoecker, CPC (2019)]

Feeddown to yields of light nuclei seldom considered in HICs

Feeddown from nuclear fragments



${}^4\text{He}$

E_x (MeV)	J^π	Decay
g.s.	0^+	
20.21	0^+	p
21.01	0^-	p, n
21.84	2^-	p, n
23.33	2^-	p, n
23.64	1^-	p, n, (γ)
24.25	1^-	p, n, d
25.28	0^-	p, n
25.95	1^-	p, n, γ
27.42	2^+	p, n, d
28.31	1^+	p, n, d
28.37	1^-	(p, n), d
28.39	2^-	(p, n), d
28.64	0^-	d
28.67	2^+	d, γ
29.89	2^+	(p, n), d

${}^4\text{H}$

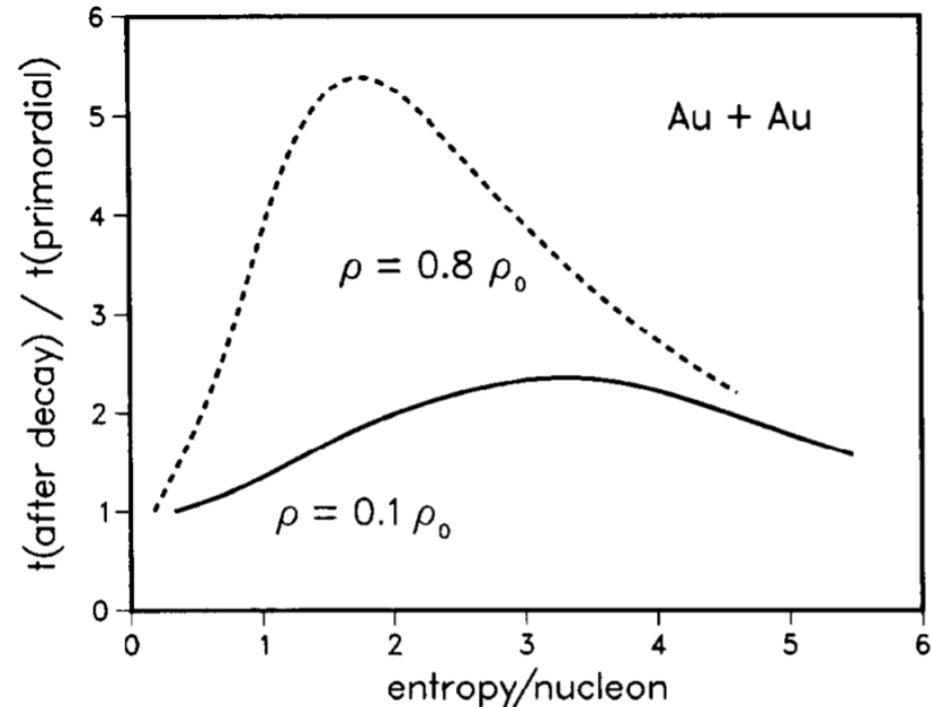
E_x (MeV)	J^π	Decay
g.s. ^a	2^-	n, ${}^3\text{H}$
0.31	1^-	n, ${}^3\text{H}$
2.08	0^-	n, ${}^3\text{H}$
2.83	1^-	n, ${}^3\text{H}$

${}^4\text{Li}$

E_x (MeV)	J^π	Decay
g.s. ^a	2^-	p, ${}^3\text{He}$
0.32	1^-	p, ${}^3\text{He}$
2.08	0^-	p, ${}^3\text{He}$
2.85	1^-	p, ${}^3\text{He}$

[Tilley, Weller, Hale, NPA '92]

See also <https://www.nndc.bnl.gov/>



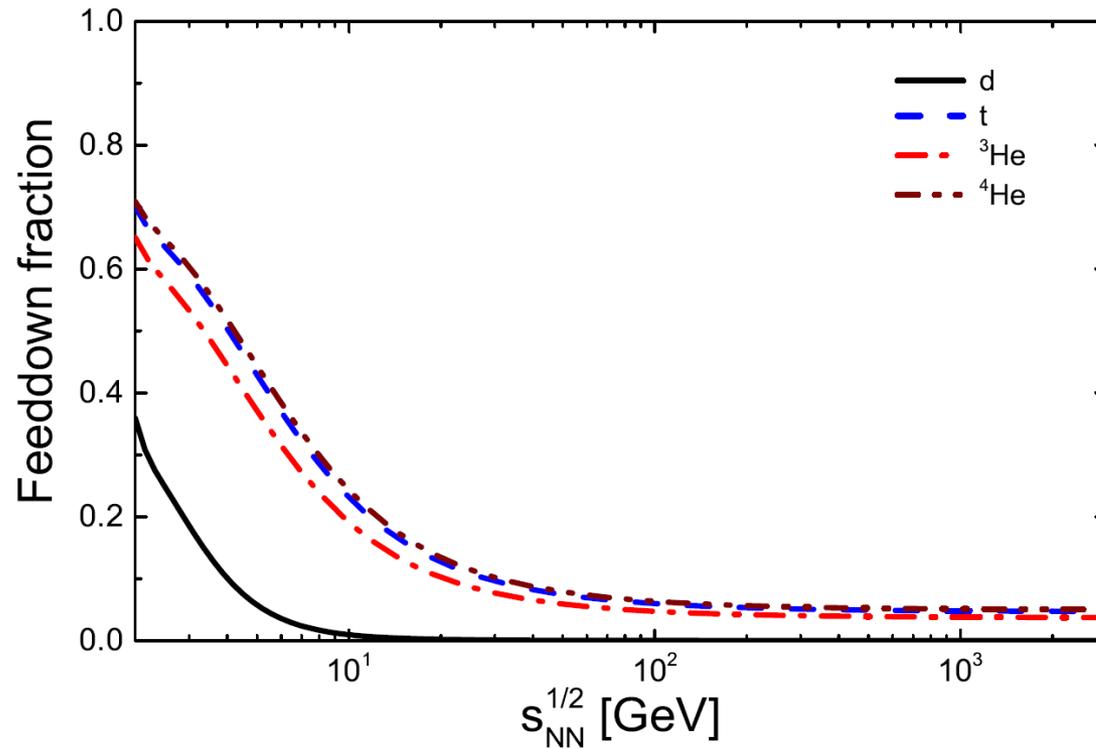
[Hahn, Stöcker, NPA '88]

In what follows feeddown from known $A = 4$ and significant $A = 5$ unstable fragments included. Fragments modeled as point particles.

Feeddown from nuclear fragments



Feeddown fraction along the phenomenological freeze-out curve



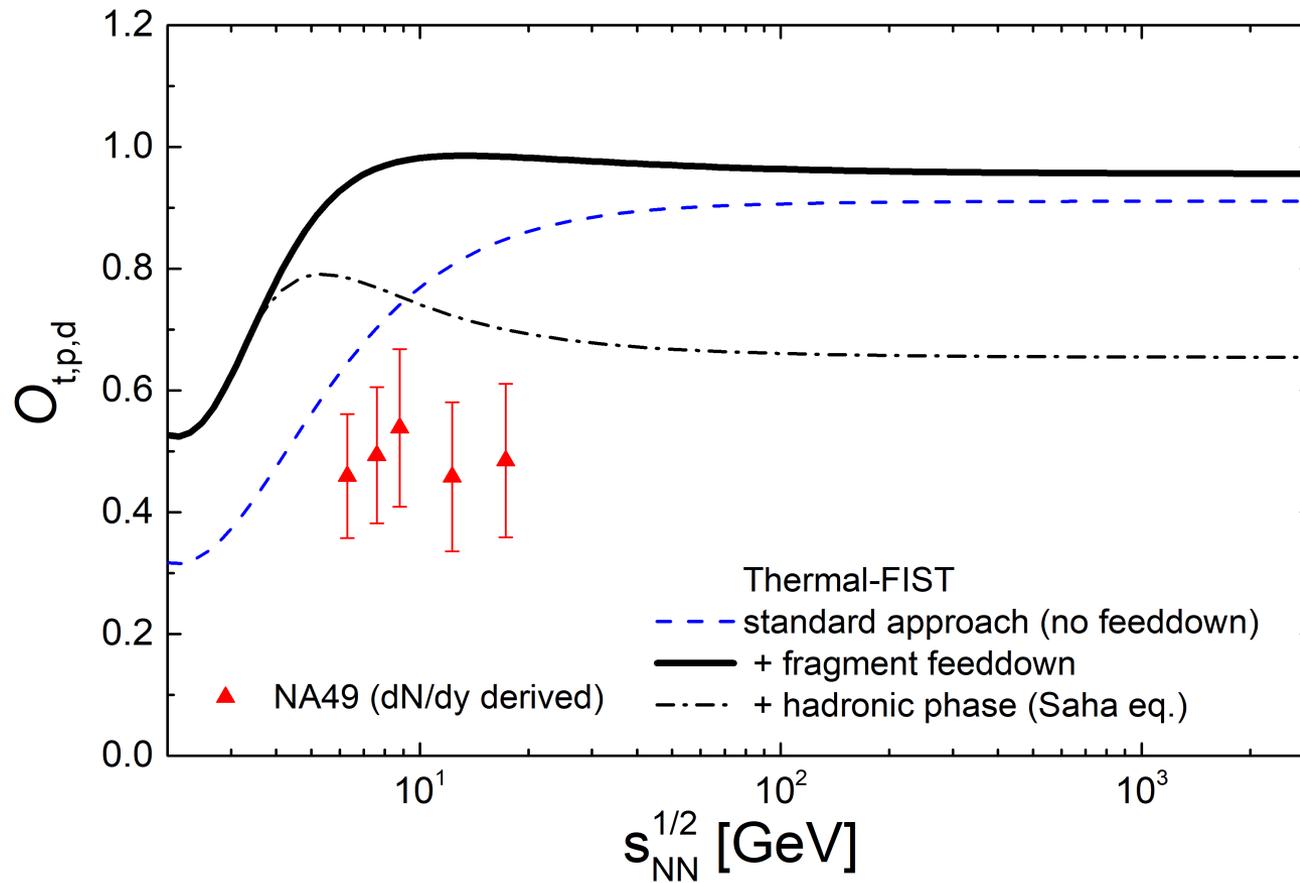
- **LHC:** 5% effect. Can be measured through p- ^3He , p- ^4He correlation?
- **RHIC/SPS:** 10-40% effect
- **GSI-HADES/FAIR:** Feeddown accounts for more than half of t, ^3He , ^4He

Feeddown from nuclear fragments: $O_{t,p,d}$



$O_{t,p,d} = N_t N_p / (N_d)^2$ suggested as a possible probe of critical behavior

[K.J. Sun et al., PLB '17, PLB '18]



Possible to obtain a non-monotonic behavior of $O_{t,p,d}$ in ideal gas picture

Summary and outlook



- **Thermal-FIST** is a user-friendly open source package for general purpose statistical-thermal model applications, in particular nuclei.
- Multiplicity dependence of light nuclei abundances at the LHC consistent with basic canonical suppression considerations, but no simultaneous description of the p/π ratio
- Saha equation extends the statistical approach down to the kinetic freeze-out, offers possible explanation why the thermal model for point-like nuclei works so well.
- Feeddown from unstable fragments is sizable for yields of t , ${}^3\text{He}$, ${}^4\text{He}$

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Thanks for your attention!

Backup slides

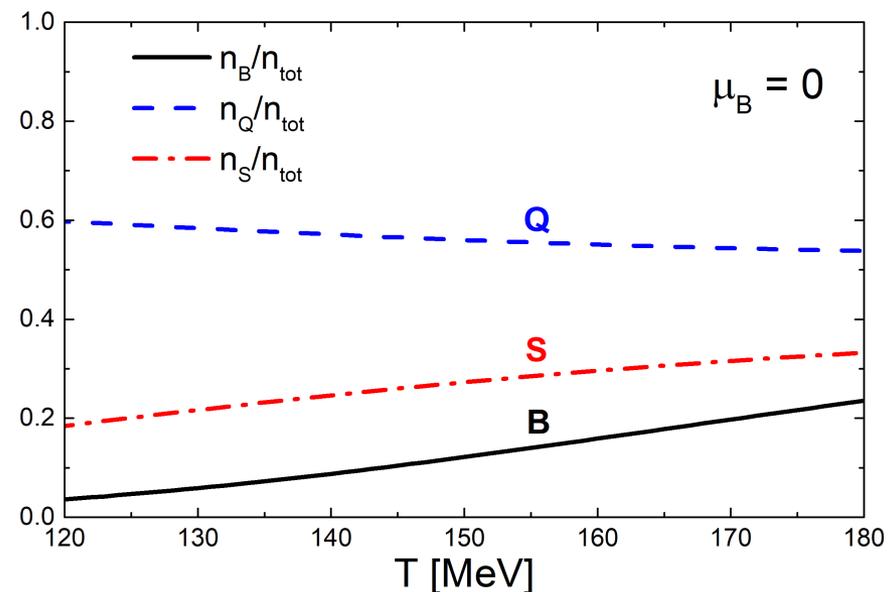
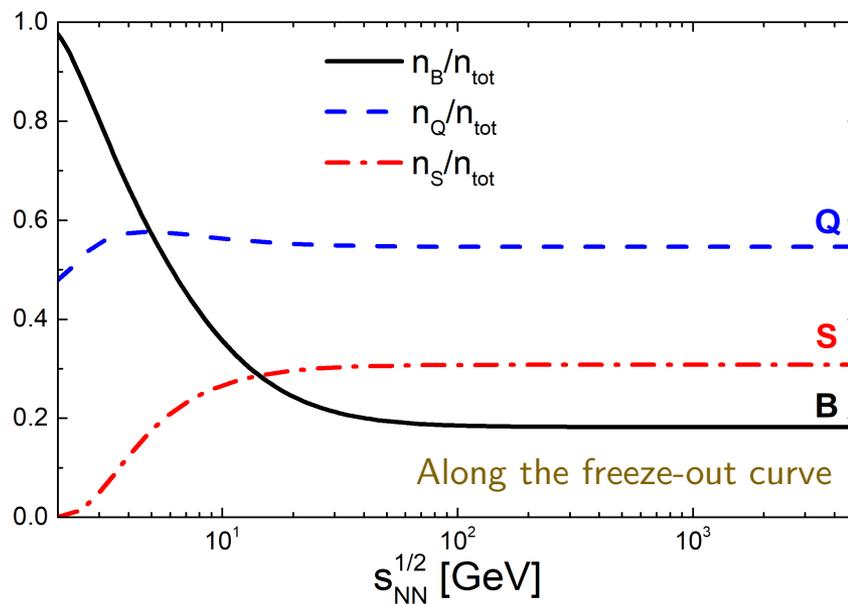
When is the canonical treatment necessary?



Normally, when the total number of particles carrying a conserved charge is **smaller or of the order of unity**

The canonical treatment is often restricted to strangeness only (**SCE**)

[STAR collaboration, 1701.07065; ALICE collaboration, 1807.11321]



- **Strangeness** conservation is most important at low energies (**HADES, CBM**)
- *Small systems at RHIC and LHC*: exact **baryon** conservation at least as important as **strangeness**

CSM at LHC



Enforce exact conservation of charges, $B = Q = S = 0$, in a *correlation volume* V_C around midrapidity

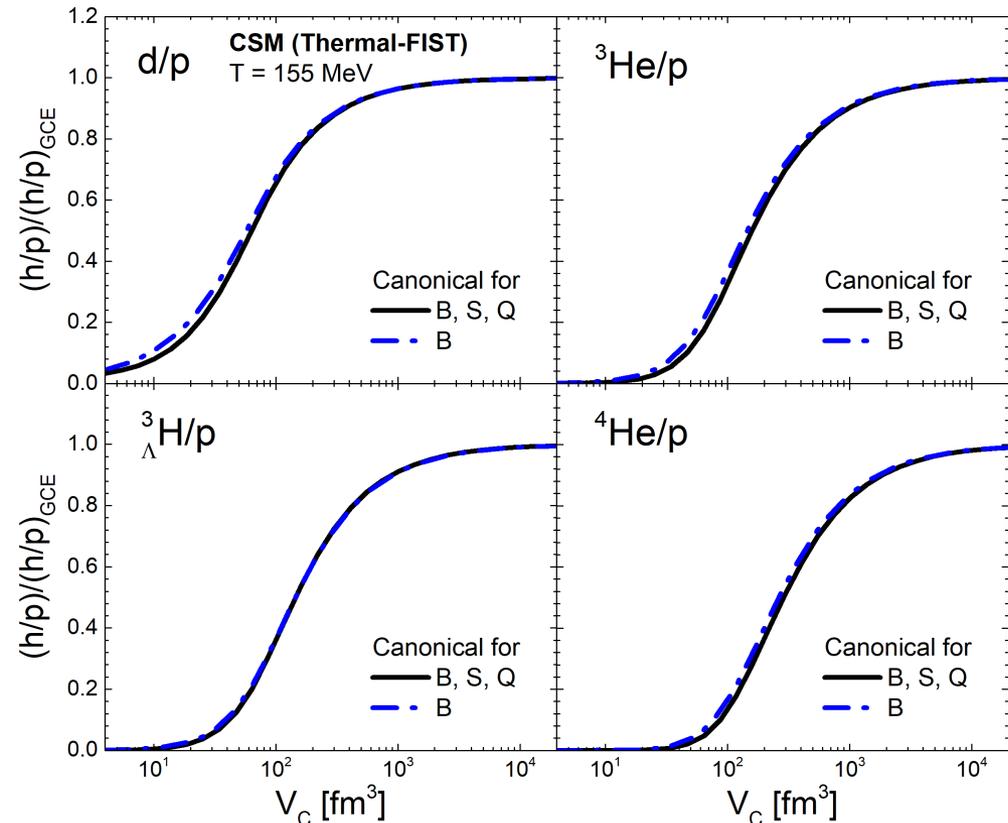
In general, $V_C \neq dV/dy$

Causality argument: exact conservation across a few units of rapidity?

[Castorina, Satz, 1310.6932]

New application: CSM for light nuclei

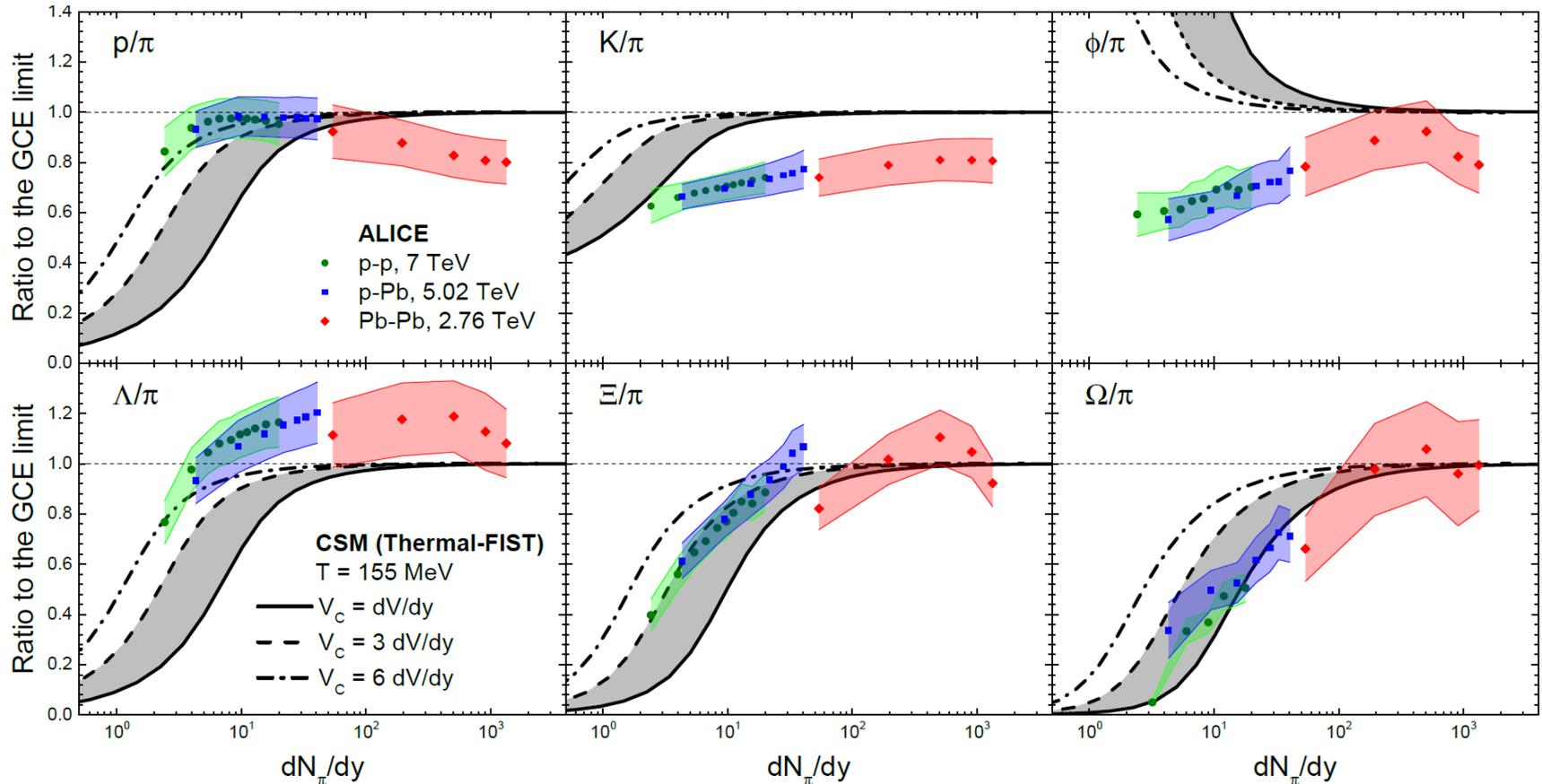
- Suppression of nuclei-to-proton ratios at low multiplicities
- For these observables sufficient to enforce exact baryon conservation only



“Vanilla” CSM at LHC



$T_{ch} = 155$ MeV, $V_C = 3dV/dy$, multiplicity dependence driven by V_C only

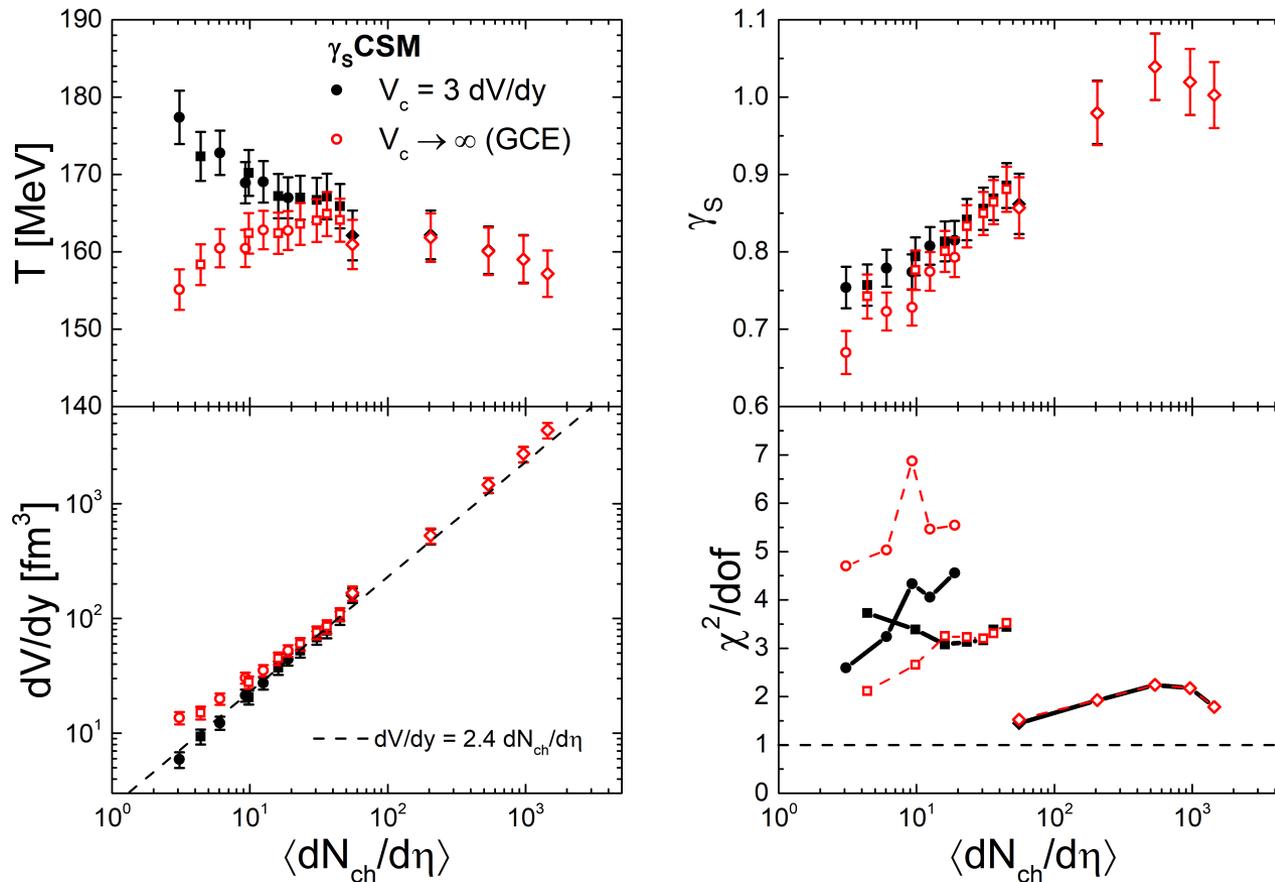


Fair for hyperons, protons and kaons worse, ϕ goes in the opposite direction

Full CSM analysis

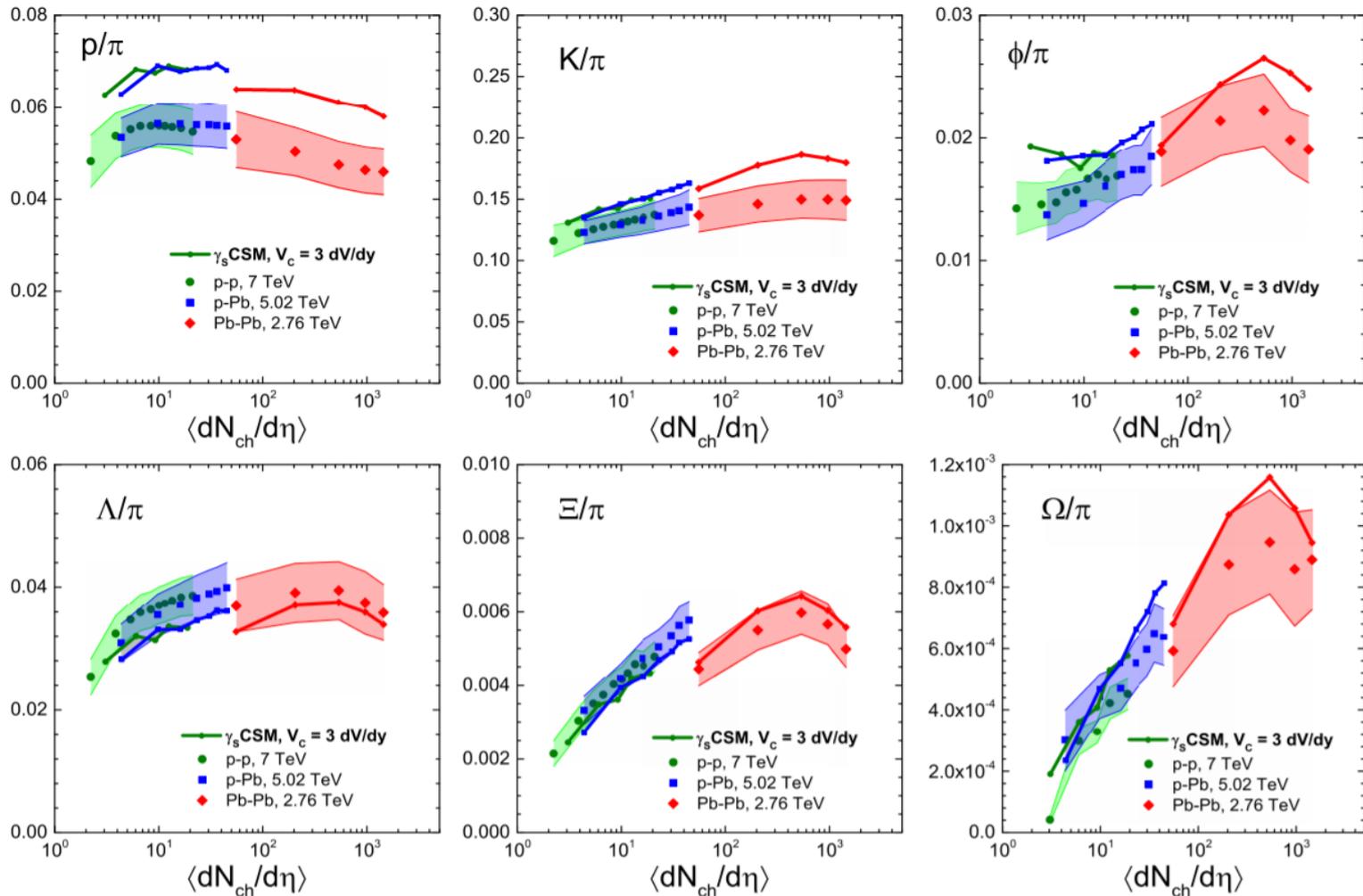


γ_S CSM: $V_C = 3dV/dy$, fit T_{ch} and γ_S at each centrality in p-p, p-Pb, Pb-Pb



Canonical suppression and strangeness saturation important below $dN_{ch}/d\eta \cong 100$

Full CSM analysis: yields

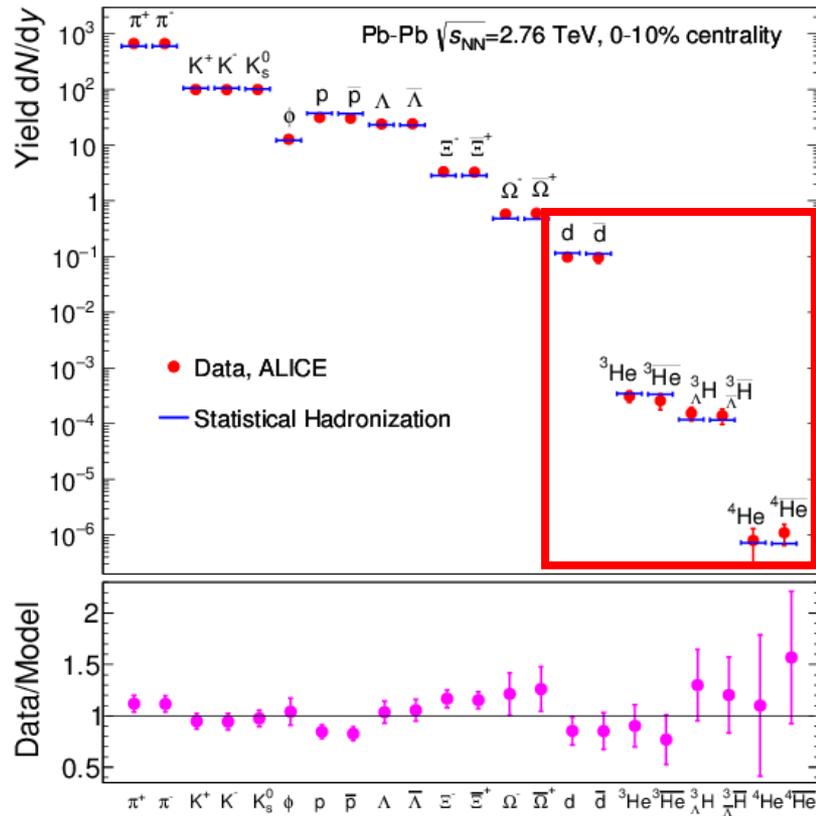


Relative accuracy of $\gamma_s \text{CSM}$ is $\sim 15\%$ across all multiplicity bins

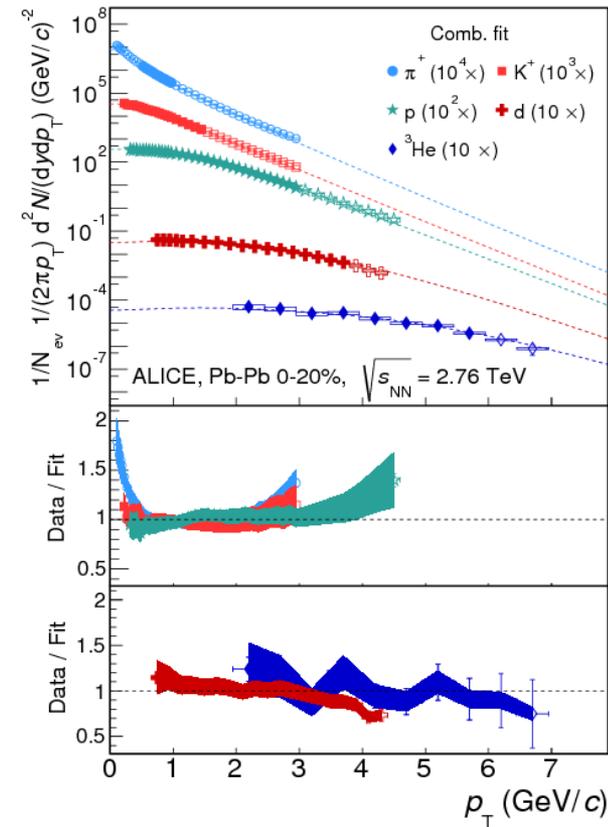
Two experimental observations at the LHC



1. Measured yields are described by thermal model at $T_{ch} \approx 155 \text{ MeV}^*$



2. Spectra described by blast-wave model at $T_{kin} \approx 100 - 120 \text{ MeV}^*$

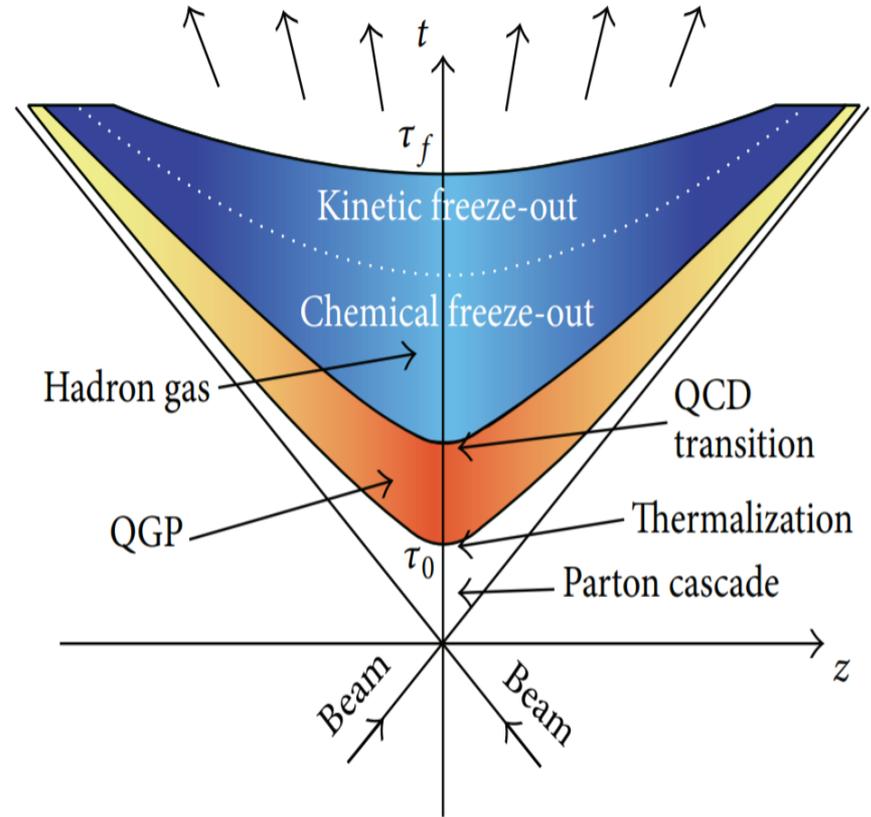
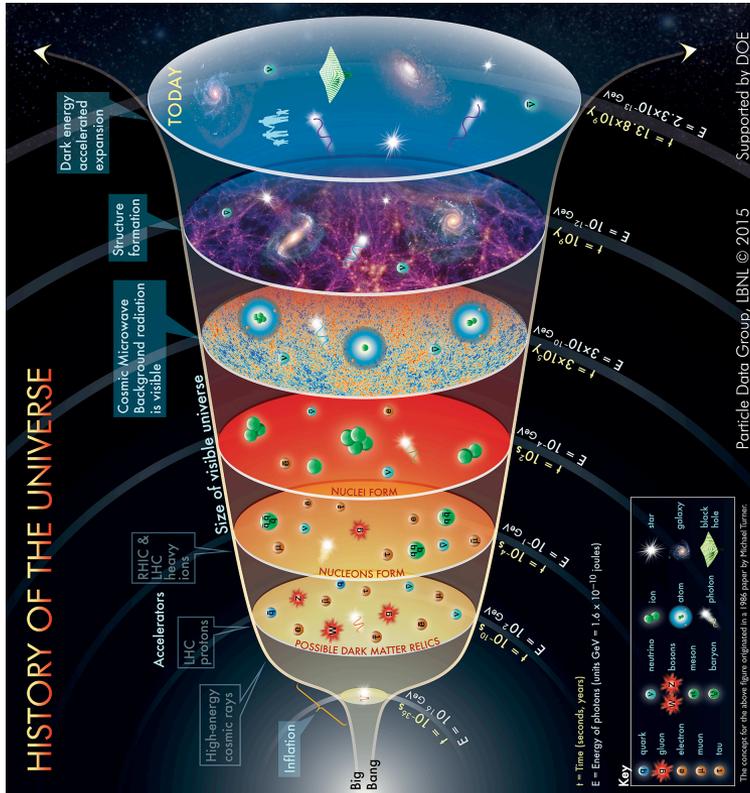


[A. Andronic et al., Nature **561**, 321 (2018)]

[ALICE collaboration, PRC **93**, 024917 (2016)]

What happens between T_{ch} and T_{kin} ?

Big Bang vs “Little Bangs”



- Hadrons (nucleons) form and “freeze-out” chemically before nuclei
- Bosons (photons or pions) catalyse nucleosynthesis

$$\text{e.g. } p + n \leftrightarrow d + \gamma \quad \text{vs} \quad p + n + \pi \leftrightarrow d + \pi$$

Big Bang vs LHC nucleosynthesis

Similarities:

- Inelastic nucleonic reactions freeze-out before nuclei formation
- Isentropic expansion of boson-dominated matter (photons in BBN vs mesons in HIC), baryon-to-boson ratio: $\eta_{BBN} \sim 10^{-10}$, $\eta_{LHC} \sim 0.05$
- Strong nuclear formation and regeneration reactions → **Saha equation**

Differences:

- Time scales: 1-100 s in BBN vs $\sim 10^{-22}$ s in HIC
- Temperatures: $T_{BBN} < 1$ MeV vs $T_{HIC} \sim 100$ MeV
- Binding energies, proton-neutron mass difference, and neutron lifetime important in BBN, less so in HICs
- $\mu_B \approx 0$ at the LHC, $\mu_B \neq 0$ in BBN
- Resonance feeddown important at LHC, irrelevant in BBN

LHC nucleosynthesis: simplified setup



- Chemical equilibrium lost at $T_{ch} = 155$ MeV, abundances of nucleons are frozen and acquire effective fugacity factors: $n_i = n_i^{eq} e^{\mu_N/T}$
- Isentropic expansion driven by effectively massless mesonic d.o.f.

$$\frac{V}{V_{ch}} = \left(\frac{T_{ch}}{T}\right)^3, \quad \mu_N \simeq \frac{3}{2} T \ln\left(\frac{T}{T_{ch}}\right) + m_N \left(1 - \frac{T}{T_{ch}}\right)$$

- Detailed balance for nuclear reactions, $X + A \leftrightarrow X + \sum_i A_i$, X is e.g. a pion

$$\frac{n_A}{\prod_i n_{A_i}} = \frac{n_A^{eq}}{\prod_i n_{A_i}^{eq}}, \quad \Leftrightarrow \quad \mu_A = \sum_i \mu_{A_i}, \quad \text{e.g. } \mu_d = \mu_p + \mu_n, \quad \mu_{3\text{He}} = 2\mu_p + \mu_n, \quad \dots$$

Saha equation



$$X_A = d_A \left[(d_M)^{A-1} \zeta(3)^{A-1} \pi^{\frac{1-A}{2}} 2^{-\frac{3+A}{2}} \right] A^{5/2} \left(\frac{T}{m_N}\right)^{\frac{3}{2}(A-1)} \eta_B^{A-1} \exp\left(\frac{B_A}{T}\right)$$

$$d_M \sim 11 - 13, \quad \eta_B \simeq 0.03$$

$$\text{BBN: } X_A = d_A \left[\zeta(3)^{A-1} \pi^{\frac{1-A}{2}} 2^{\frac{3A-5}{2}} \right] A^{\frac{5}{2}} \left(\frac{T}{m_N}\right)^{\frac{3}{2}(A-1)} \eta^{A-1} X_p^Z X_n^{A-Z} \exp\left(\frac{B_A}{T}\right)$$

(Simplified) Saha equation vs thermal model



Saha equation:
$$\frac{N_A(T)}{N_A(T_{\text{ch}})} \simeq \left(\frac{T}{T_{\text{ch}}}\right)^{\frac{3}{2}(A-1)} \exp \left[B_A \left(\frac{1}{T} - \frac{1}{T_{\text{ch}}} \right) \right]$$
 $B_A \ll T$

Thermal model:
$$\left[\frac{N_A(T)}{N_A(T_{\text{ch}})} \right]_{\text{eq.}} \simeq \left(\frac{T}{T_{\text{ch}}}\right)^{\frac{3}{2}A} \exp \left[-m_A \left(\frac{1}{T} - \frac{1}{T_{\text{ch}}} \right) \right]$$
 $m_A \gg T$

Strong exponential dependence on the temperature is eliminated in the Saha equation approach

Further, quantitative applications require numerical treatment of full spectrum of *massive* mesonic and baryonic resonances

Saha equation: Entropy production effect

