

# High-order cumulants near the critical point from molecular dynamics

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# QCD matter

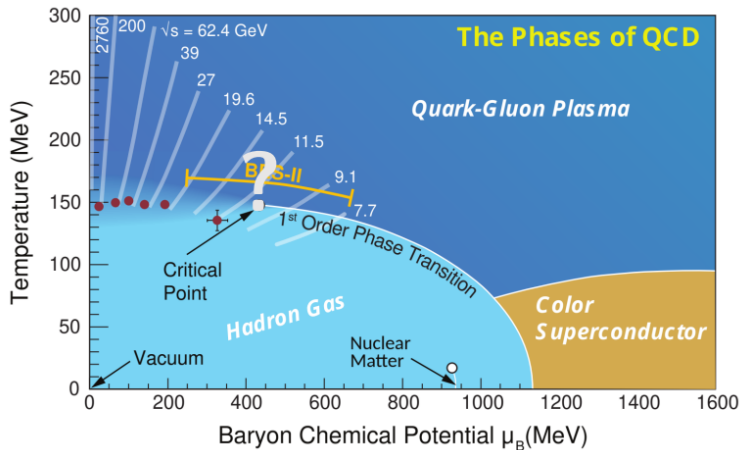


Figure from Bzdak et al., Phys. Rept. 2020 & 2015 Long Range plan

# Fluctuations as CP signature

V. A. Kuznietsov et al., Phys. Rev. C 105, 044903 (2022)

In GCE density cumulants shows singularity behavior in the critical point.

$$\ln Z^{\text{gce}}(T, V, \mu)$$

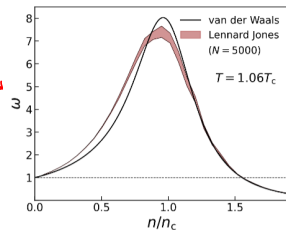
$$= \ln \left[ \sum_N e^{\mu N} Z^{\text{ce}}(T, V, N) \right],$$

$$\kappa_n \cong \frac{\partial^n (\ln Z^{\text{gce}})}{\partial (\mu_N)^n}.$$

The real expression for  $Z^{\text{gce}}$  is unknown in QCD matter.

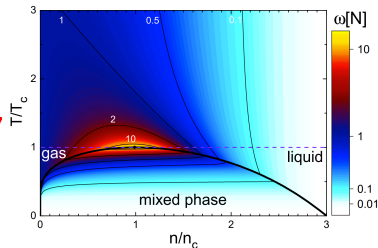
V. Vovchenko, D. V. Anchishkin, M. I. Gorenstein and R. V. Poberezhnyuk, Phys. Rev. C92 054901 (2015)

## Scaled variance



$$\omega = \kappa_2 / \kappa_1$$

$$\kappa_2 \sim \xi^2$$

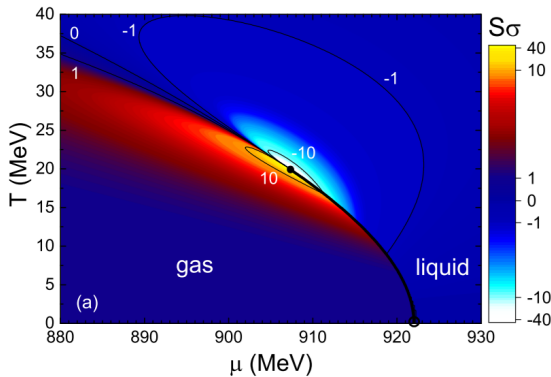


# Fluctuations as CP signature

M.A. Stephanov, PRL, 102 (2009), 032301

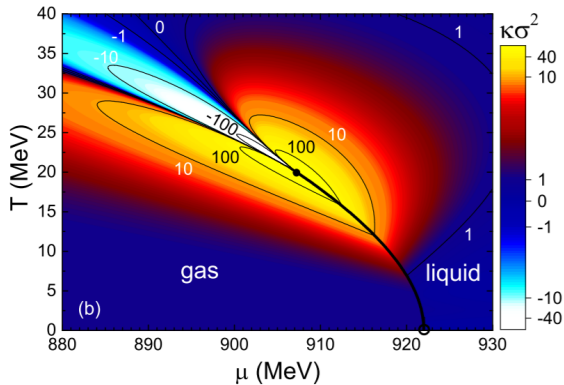
$$\kappa_3 \sim \xi^{9/2}$$

Skewness  $S\sigma = \kappa_3/\kappa_2$



$$\kappa_4 \sim \xi^7$$

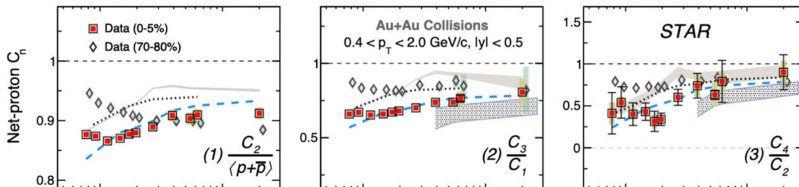
Kurtosis  $\kappa\sigma^2 = \kappa_4/\kappa_2$



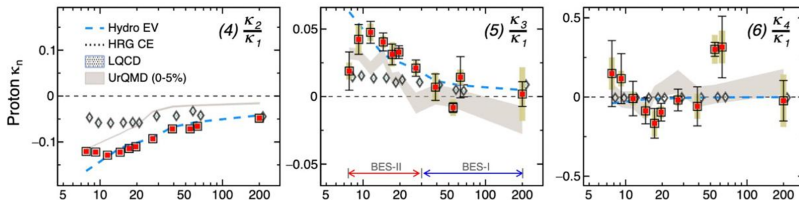
V. Vovchenko et al, PRC 92 (2015) 5, 054901

# STAR measurments

## Net-proton cumulant ratios



## Proton/antiproton factorial cumulant ratios



# Connection to the experiment

## Theory

- Coordinate and/or momentum space
- In contact with the heat bath
- Conserved charges
- Uniform
- Fixed volume
- No CP at the transport models

## Experiment

- Momentum space
- Expanding in vacuum
- Non-conserved charges
- Inhomogeneous
- Fluctuating volume
- 3D Ising CP possibility

Need microscopic description of fluctuations

# Lennard-Jones potential

The Lennard-Jones potential reads

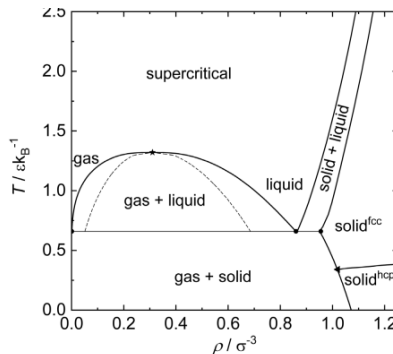
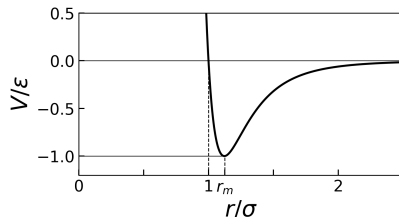
$$V_{\text{LJ}} = 4\varepsilon \left[ \left( \frac{\sigma}{r} \right)^{12} - \left( \frac{\sigma}{r} \right)^6 \right],$$

In reduced dimensionless variables it can be rewritten as

$$\tilde{V}_{\text{LJ}} = 4[\tilde{r}^{-12} - \tilde{r}^{-6}],$$

where the reduced variables are used:  $\tilde{r} = r/\sigma$   
and  $\tilde{V}_{\text{LJ}} = V_{\text{LJ}}/\varepsilon$ .

The LJ fluid contains a critical point in the 3D  
Ising universality class.



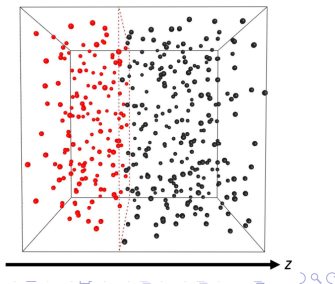
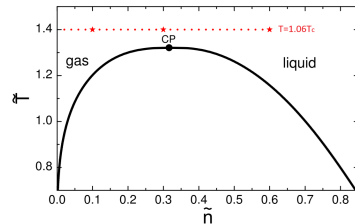
# Simulation setup

$$m \frac{d^2 \tilde{\mathbf{r}}_{ij}}{d\tilde{t}^2} = -\vec{\nabla} \tilde{V}_{LJ}(\tilde{\mathbf{r}}_{ij})$$

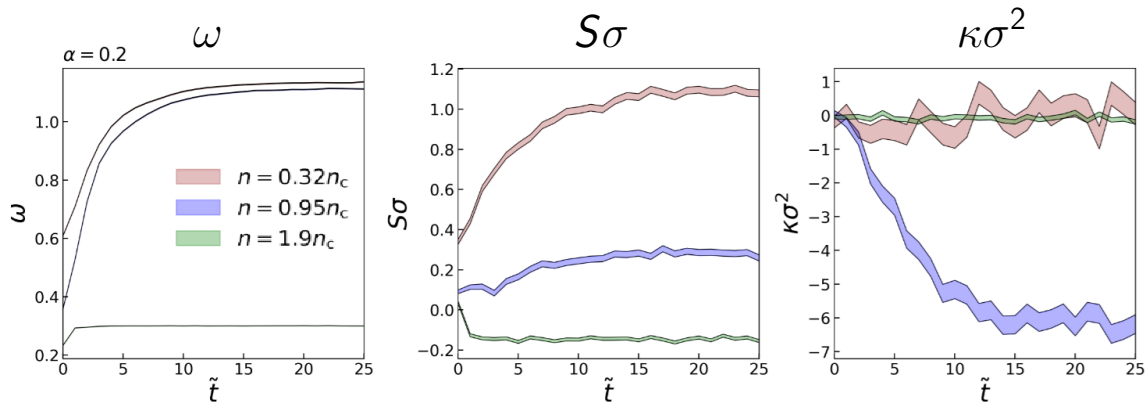
- Three points on the phase diagram,  $\tilde{n} = 0.1 \approx 0.3n_c$ ,  $\tilde{n} = 0.3 \approx 0.95n_c$ ,  $\tilde{n} = 0.6 \approx 2n_c$ , all for  $T = 1.06T_c$
- $N_{\text{ev}} \simeq 2$  million events at each density.
- Initialize each event with random initial coordinates and momenta
- Run each event for long time ( $\tilde{t} = 50$ ), write snapshots to file at regular time intervals
- Calculate observables as event-by-event (ensemble) average

The simulations are performed on UH PhysGPU cluster.

Code is available at: <https://github.com/vlvovch/lennard-jones-cuda>



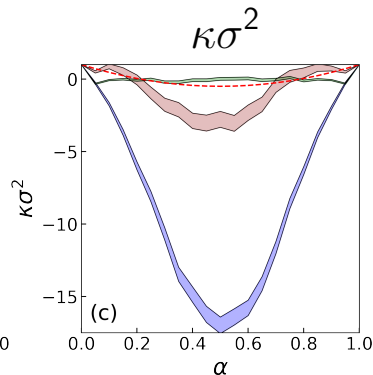
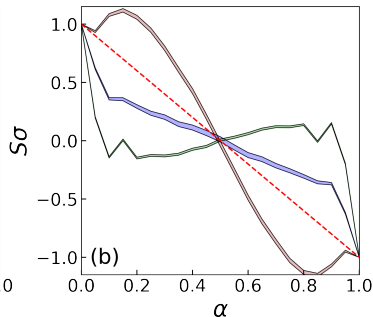
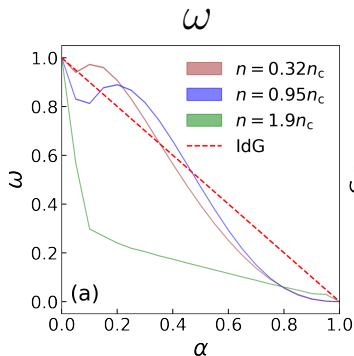
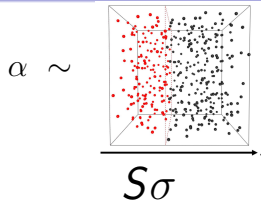
# Cumulants equilibration



All three cumulant ratios equilibrate at comparable time scales

# Acceptance dependence of cumulants

Large CP signal in the coordinate space for all cumulants.



# Subensemble method – correction for global conservation

Cumulants, corrected for the baryon conservation reads (**V Vovchenko O. Savchuk, R. Poberezhnyuk, M. Gorenstein, V. Koch, Phys. Let. B, 2020**)

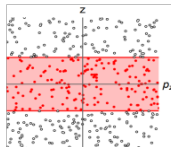
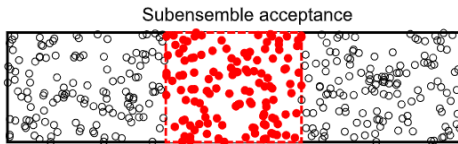
$$\tilde{\omega} \equiv \omega / (1 - \alpha),$$

$$\tilde{s}\sigma \equiv s\sigma / (1 - 2\alpha),$$

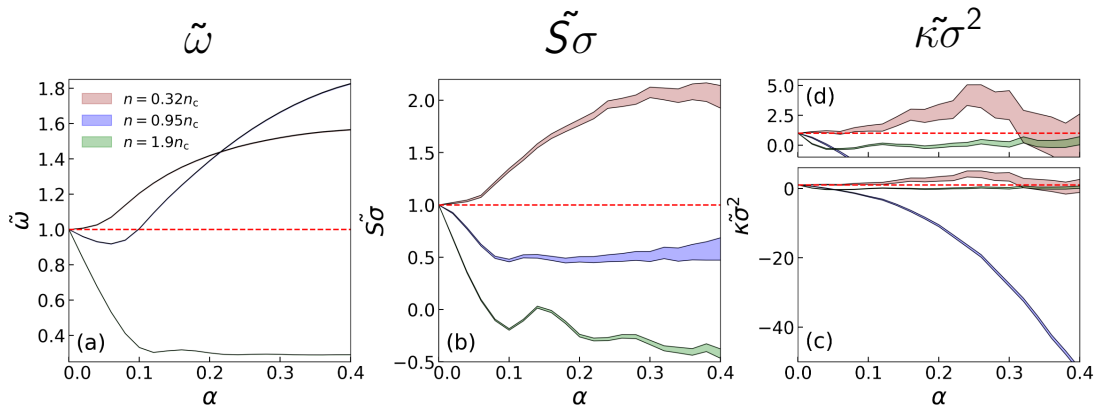
$$\tilde{\kappa}\sigma^2 \equiv \frac{1}{(1 - 3\alpha(1 - \alpha))} \kappa\sigma^2 - \frac{3\alpha(1 - \alpha)}{1 - 3\alpha(1 - \alpha)} (\tilde{s}\sigma)^2,$$

where  $\alpha = \langle N \rangle / N$  - subensemble fraction.

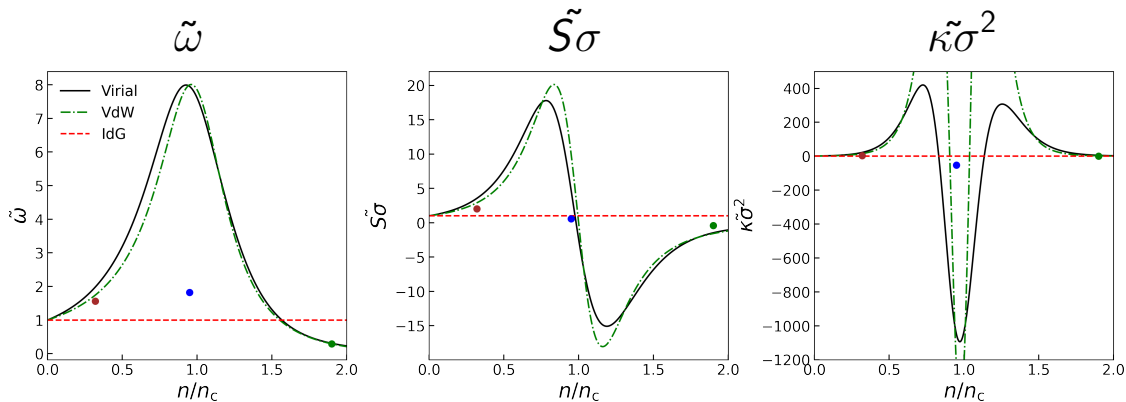
Unfortunately, higher corrected cumulants have pole at  $\alpha = 0.5$ , which leads to the artificial infinities because we compute the moments values with finite precision.



# Cumulants with global conservation corrections



# Virial estimation of GCE limit



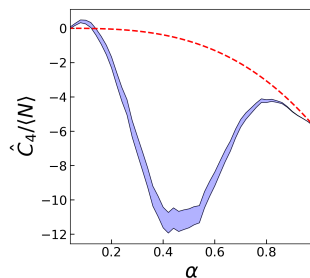
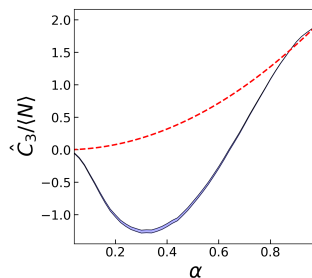
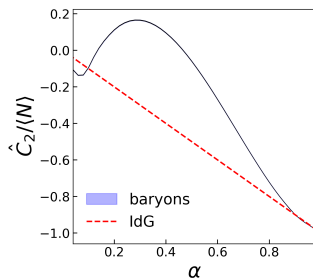
# Factorial cumulants ( $n = 0.95n_c$ )

One can express the factorial cumulants in terms of the mean density  $\langle N \rangle = \kappa_1$  and standard cumulants  $\kappa_i$

A. Bzdak, V. Koch, and N. Strodthoff, *Phys. Rev. C* **95**, 054906 (2017)

$$\hat{C}_2 = \kappa_2 - \langle N \rangle, \quad \hat{C}_3 = \kappa_3 - 3\kappa_2 + 2\langle N \rangle,$$

$$\hat{C}_4 = -6\langle N \rangle + 11\kappa_2 - 6\kappa_3 + \kappa_4.$$



# Baryons and protons

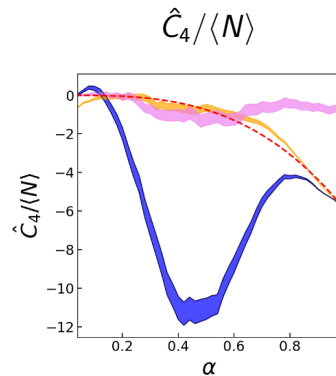
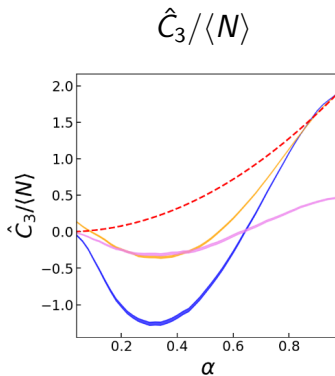
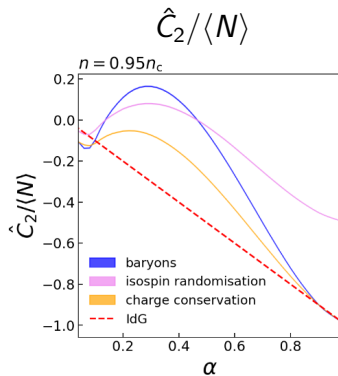
Experiments measure protons, rather than not all baryons. There are two ways to incorporate protons:

- 1 **Isospin randomisation:** Each time, randomly label a baryon as a proton with a probability of 50%. This implies isospin randomization expected to be valid for large collision energies ( $\geq 10$  GeV). In this case, factorial cumulants scale as  $F_k^{\text{baryon}} = 2^k F_k^{\text{proton}}$  for a symmetric system.
- 2 **Charge conservation (conserved proton number):** Neglect the production of pions: charge conservation implies conservation of proton number. Treat the first  $N/2$  particles as protons. Expected to be appropriate for smaller energies ( $< 10$  GeV).

M. Kitazawa and M. Asakawa, Phys. Rev. C 86, 024904 (2012)

M. Barej and A. Bzdak, Phys. Rev. C 102, 064908 (2020)

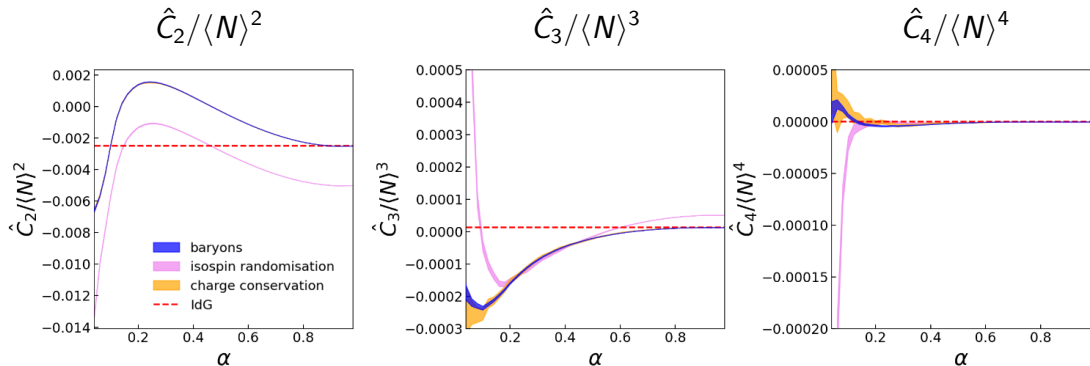
# Factorial cumulants ( $n = 0.95n_c$ )



Taking protons instead of baryons significantly reduces the signal, and no significant deviation from the ideal gas baseline seen for  $\hat{C}_4 \Rightarrow$  may be relevant for RHIC-BES-II observations?

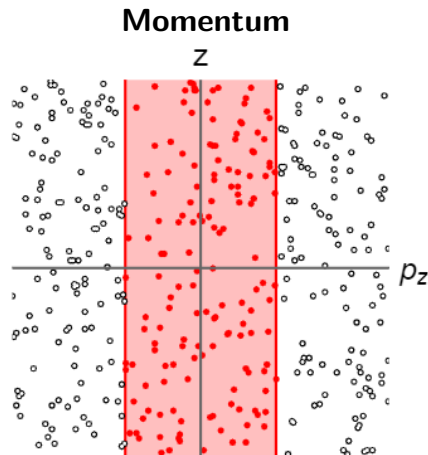
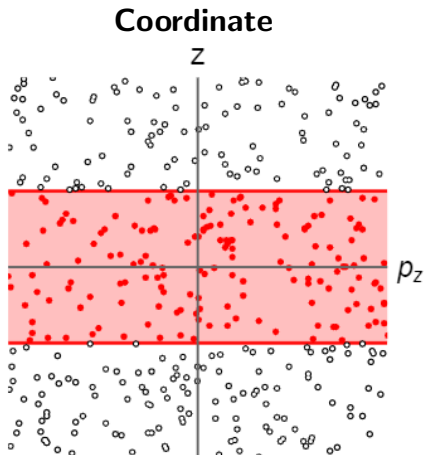
# Normalised factorial cumulants ( $n = 0.95n_c$ )

Acceptance dependence of normalized FC suggested in [A. Bzdak et al., arXiv:2503.16405](#)



Normalized factorial cumulants show non-flat acceptance dependence, indicating short-range correlations, the signal is diluted in the case of protons

# Momentum vs coordinate cuts



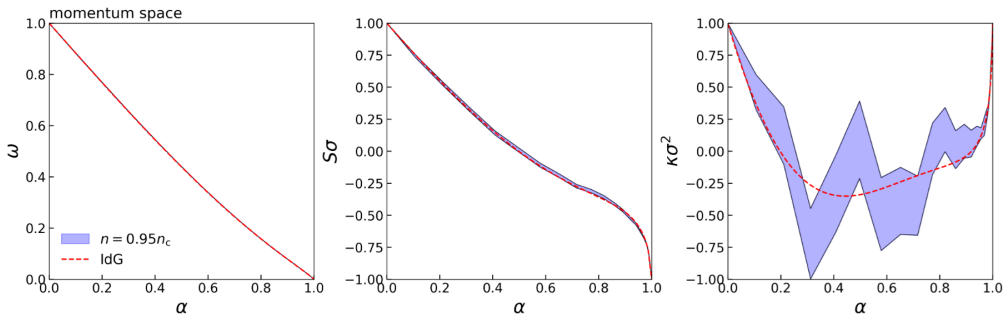
# Fluctuations in the momentum space

Red dashed line corresponds to the large  $N$  limit formulas for the Ideal gas result with the energy conservation. For scaled variance such formula could be written in a compact form

V. A. Kuznietsov et al.,  
PRC 105, 044903 (2022)

$$\omega_{\text{IdG}}^{\text{mom}} = 1 - \alpha - \frac{2e^{-2\text{erf}^{-1}(\alpha)^2}\text{erf}^{-1}(\alpha)^2}{3\pi\alpha}.$$

Similar results were found for skewness and kurtosis could, but the analytical form becomes more complicated.



# Collective flow

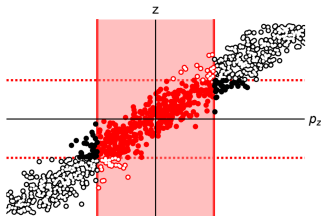
We now add collective flow to the longitudinal velocity (rapidity) as follows

$$\tilde{v}_z = \tilde{v}_z^{\text{LJ}} + \kappa \left( \frac{\tilde{z}}{\tilde{L}} - \frac{1}{2} \right)$$

thermal part

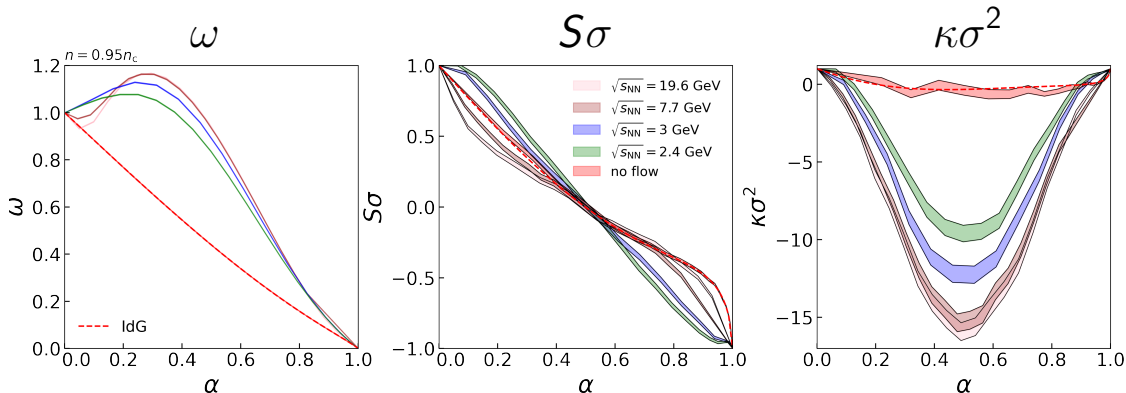
collective flow part

$$y(\tilde{z}) = y_{\text{th}} + \sqrt{\frac{T_{\text{frz}}}{m\tilde{T}}} \kappa \left( \frac{\tilde{z}}{\tilde{L}} - \frac{1}{2} \right).$$



We map  $\kappa$  to the collision energies, using the heavy-ion inspired parameters:  
 $T_{\text{frz}} = 150 \text{ MeV}$ ,  $m = 938 \text{ MeV}$ .

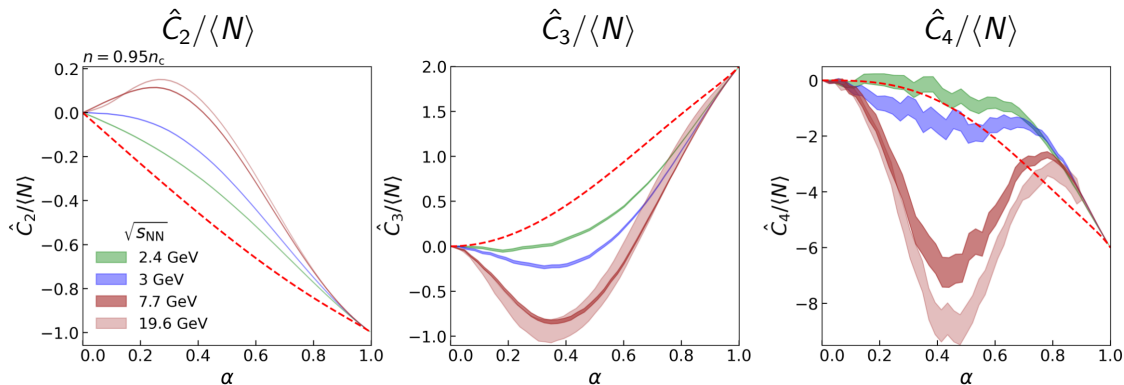
# Cumulants at fixed $\alpha$ ( $n = 0.95n_c$ )



With correlation cumulants in the momentum space becomes notably non-Gaussian.<sup>1</sup>

<sup>1</sup>NB: This is not a prediction of  $\sqrt{s_{NN}}$  dependence but an estimation of the signal for a freeze-out at critical density.

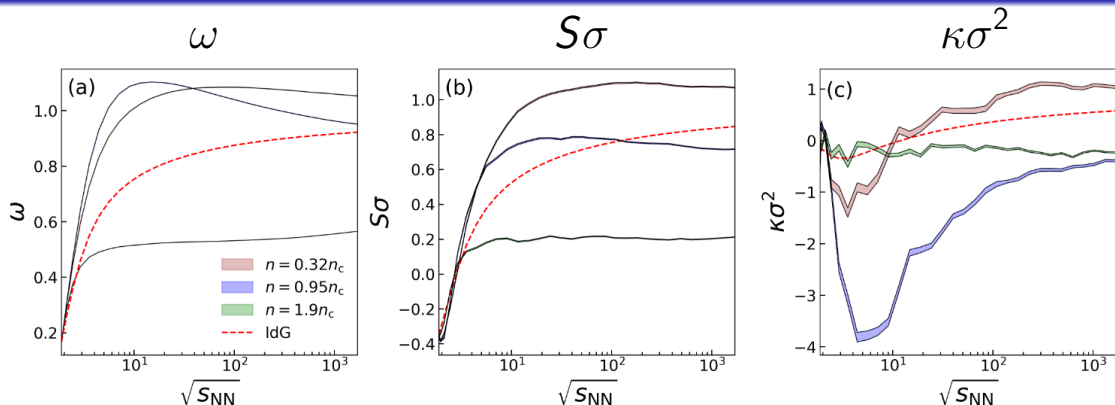
# Factorial cumulants at fixed $\alpha$ ( $n = 0.95n_c$ )



With correlation factorial cumulants in the momentum space becomes notably non-Gaussian.<sup>2</sup>

<sup>2</sup>NB: This is not a prediction of  $\sqrt{s_{NN}}$  dependence but an estimation of the signal for a freeze-out at critical density.

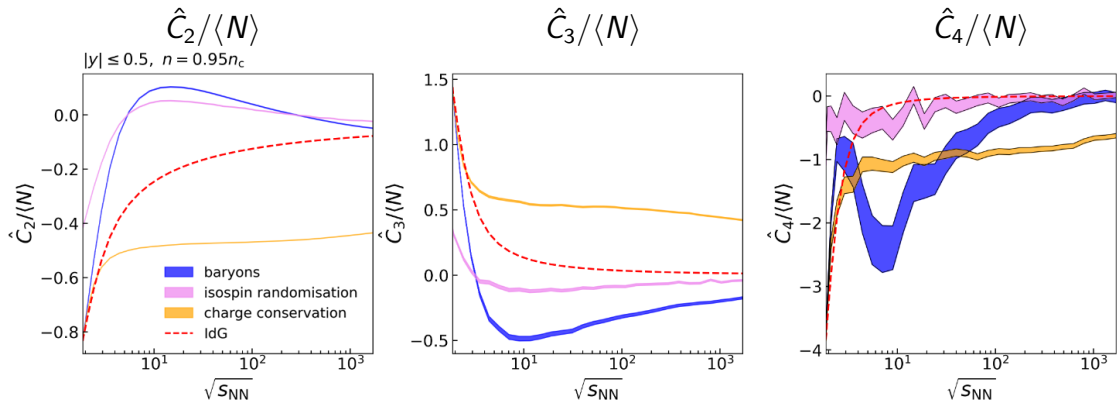
# Cumulants at fixed $|y| < 0.5$ acceptance



The interplay between the increasing flow and the decreasing of subsystem fraction  $\alpha$  gives a non-trivial dependence on  $\sqrt{s_{NN}}$ .<sup>3</sup>

<sup>3</sup>NB: Once again, this is not a prediction of actual  $\sqrt{s_{NN}}$  dependence but an estimation of the signal for a freeze-out at critical density occurring at a given  $\sqrt{s_{NN}}$ .

# Factorial cumulants at fixed $|y| < 0.5$ acceptance ( $n = 0.95n_c$ )



Proton factorial cumulants do not show as strong signal as observed for baryons.<sup>4</sup>

<sup>4</sup>NB: Once again, this is not a prediction of actual  $\sqrt{s_{NN}}$  dependence but an estimation of the signal for a freeze-out at critical density occurring at a given  $\sqrt{s_{NN}}$ .

# Summary

1. High-statistics (2 million events) simulations of the Lennard-Jones fluid have been performed to compute skewness and kurtosis of particle number for the **first time in a microscopic MD framework**.
2. Scaled variance, skewness, and kurtosis, have same equilibration rates and **show large deviations from Poisson baseline near the CP in coordinate space**. Same is true for the factorial cumulants.
3. Factorial cumulants were counted uncovering the n-particle correlations in the system. High and low energy regime for proton factorial cumulants is studied and signal suppression observed.
4. Momentum-space fluctuations coincide with the analytical expectations for the Ideal gas with the energy conservation → *no effects of the CP in momentum space even for the higher cumulants in the system without collective flow*.
5. Implementing the Bjorken-like flow in the system recovers the CP signal in the momentum space. The results confirm the sensitivity of non-Gaussian fluctuations to criticality.
  - **Future plans:** Implementation of multiple components (protons, neutrons, ...) and more realistic expansion dynamics, comparison with experimental data, expansions CP studies with LJ virial EOS.

**THANK YOU FOR ATTENTION!**

**Questions?**

# Cumulants definition

Scaled variance

$$\omega = \frac{\kappa_2}{\kappa_1} = \frac{\langle N^2 \rangle - \langle N \rangle^2}{\langle N \rangle}.$$

Skewness

$$S_\sigma = \frac{\kappa_3}{\kappa_2} = \frac{\langle N^3 \rangle - 3\langle N^2 \rangle \langle N \rangle + 2\langle N \rangle^3}{\langle N^2 \rangle - \langle N \rangle^2}.$$

(Excess) kurtosis

$$\begin{aligned} \kappa_\sigma^2 = \frac{\kappa_4}{\kappa_2} &= \frac{\langle N^4 \rangle - 4\langle N^3 \rangle \langle N \rangle + 6\langle N^2 \rangle \langle N \rangle^2 - 3\langle N \rangle^4}{\langle N^2 \rangle - \langle N \rangle^2} \\ &\quad - 3(\langle N^2 \rangle - \langle N \rangle^2). \end{aligned}$$

# Collective flow

We can identify the boundary rapidity  $y(\tilde{z} = \tilde{L}/2)$  with the center-of-mass rapidity

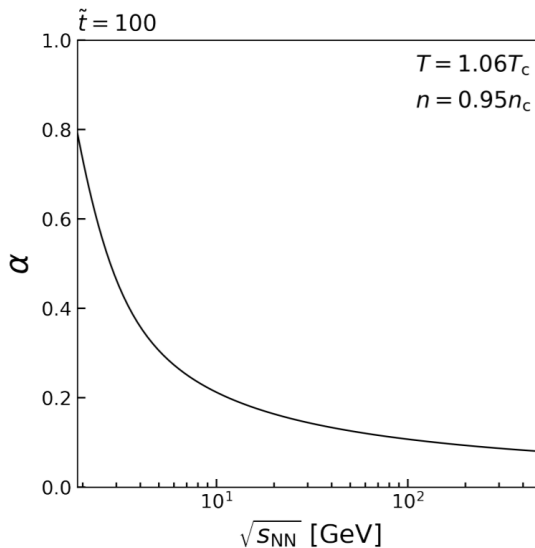
$$y_{\text{cm}} = y_{\text{beam}}/2 = \ln \left[ \frac{\sqrt{s_{\text{NN}}} + \sqrt{s_{\text{NN}} - 4m_N^2}}{2m_N} \right].$$

$$\langle y(\tilde{z} = \tilde{L}/2) \rangle = \sqrt{\frac{T_{\text{frz}}}{m\tilde{T}}} \kappa \frac{1}{2} = y_{\text{cm}},$$

therefore

$$\kappa = 2y_{\text{cm}} \sqrt{\frac{m\tilde{T}}{T_{\text{frz}}}}, \quad \sqrt{s_{\text{NN}}} = 2m \operatorname{ch} \left( \frac{\kappa}{2} \sqrt{\frac{T_{\text{fr}}}{m\tilde{T}}} \right).$$

# $\alpha$ as a function of $\sqrt{s_{\text{NN}}}$ at $y_{\text{cut}} = 0.5$



# Flow at fixed $\alpha = 0.2$

