

Phase Structure of Strongly Interacting Matter under Extreme Conditions

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Bogolyubov Readings 2025



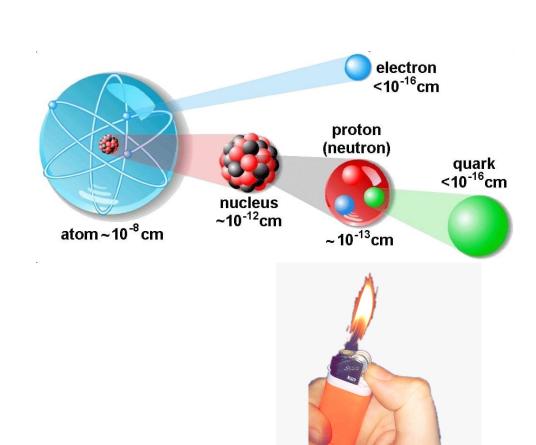
Nov 25, 2025

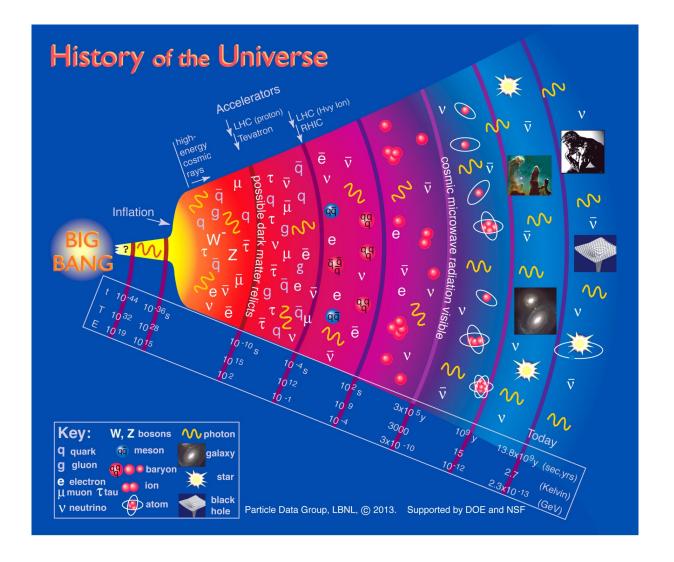




Structure of matter







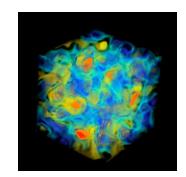
Strongly interacting matter



Theory of strong interactions: Quantum Chromodynamics (QCD)

$$\mathcal{L} = \sum_{q=u,d,s,...} ar{q} \left[i \gamma^{\mu} (\partial_{\mu} - i g A_{\mu}^{\mathsf{a}} \lambda_{\mathsf{a}}) - m_{q}
ight] q - rac{1}{4} \, G_{\mu
u}^{\mathsf{a}} G_{\mathsf{a}}^{\mu
u}$$

- Basic degrees of freedom: quarks and gluons that carry color charge
- At smaller energies confined into baryons (qqq) and mesons $(q\bar{q})$

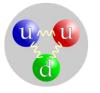


Scales

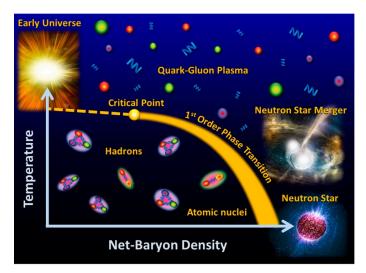
- Length: 1 femtometer = 10^{-15} m
- Temperature: $100 \text{ MeV}/k_B = 10^{12} \text{ K}$



- Early Universe
- Astrophysics: Neutron star (mergers)







QCD features and emergent phenomena

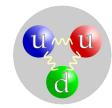


- Asymptotic freedom Gross, Politzer, Wilczek (1973)
 - Interaction becomes weaker at high energies/small distances
 - Theory is in perturbative regime at small distances

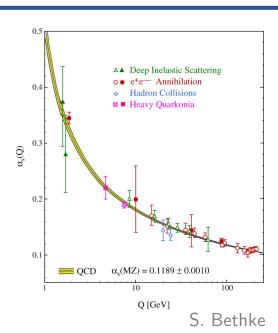


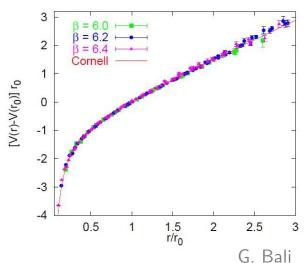
2004

- Hadrons (confinement)
 - No free quarks or gluons ever observed
 - They must form composite, color-neutral objects the hadrons
 - Proton (uud) and neutron (udd)
 - No small parameter makes the theory virtually untractable ☺



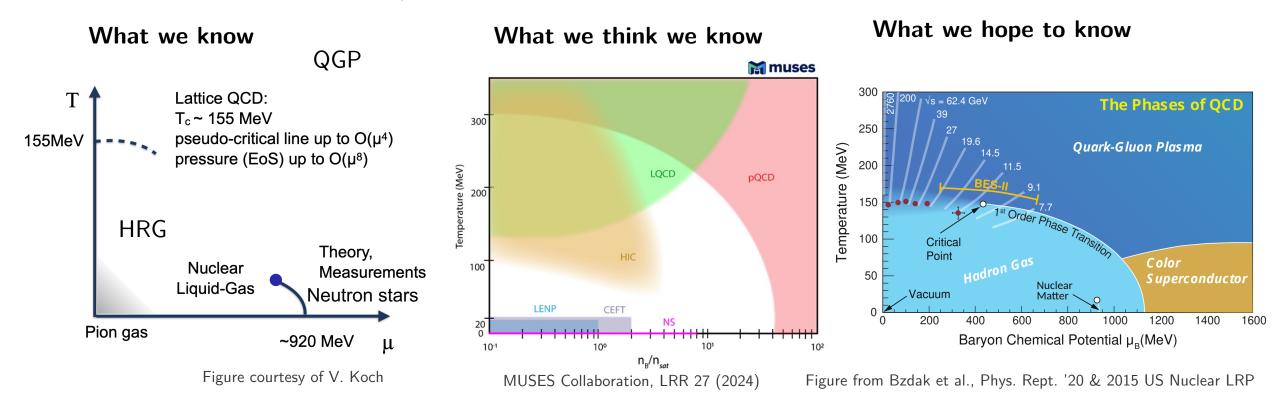
- Dynamical mass generation
 - Proton (uud) mass is $m_p = 938 \text{ MeV/c}^2$ but $m_u + m_u + m_d \sim 15 \text{ MeV/c}^2$
 - >95% of proton's mass from QCD, only <5% is from Higgs





QCD under extreme conditions

$$\mathcal{L} = \sum_{q=\mu,d,s,...} ar{q} \left[i \gamma^{\mu} (\partial_{\mu} - i g A_{\mu}^{\mathsf{a}} \lambda_{\mathsf{a}}) - m_{q}
ight] q - rac{1}{4} \, G_{\mu
u}^{\mathsf{a}} G_{\mathsf{a}}^{\mu
u}$$



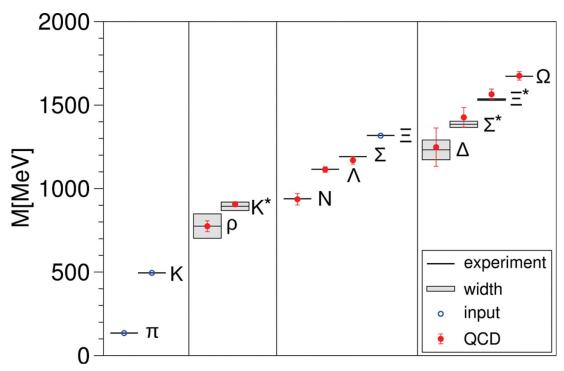
- Dilute hadron gas at low T & $\mu_{\rm B}$ due to confinement, quark-gluon plasma high T & $\mu_{\rm B}$
- Nuclear liquid-gas transition in cold and dense matter, lots of other phases conjectured

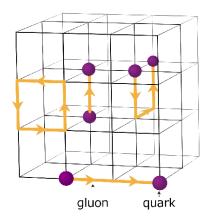
QCD Phase Diagram: From zero to non-zero density

Non-perturbative methods

First-principle tool: Lattice QCD

Ab-initio calculation of hadron masses





BMW Collaboration, Science 322, 1224 (2008)

Remarkable agreement of QCD with the experiment

QCD transition from lattice QCD



$$Z = \operatorname{Tr}(e^{-(\hat{H} - \mu \hat{N})/T}) = \int DU \det M[U, \mu] e^{-S_{YM}}$$

$$P = P(T, \mu)$$
equation of state
lattice QCD
$$P = P(T, \mu)$$

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equation of state
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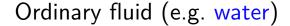
$$P = P$$

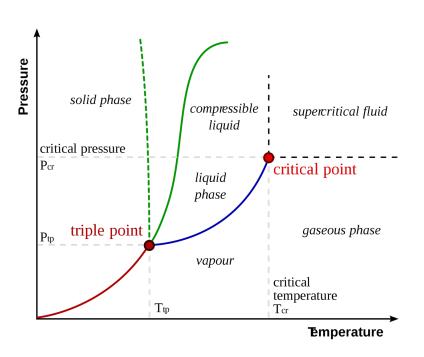
- Analytic crossover at vanishing net baryon density at $T_{pc} \approx 155$ MeV a first-principle result [Y. Aoki et al., Nature 443, 675 (2006)]
- Finite density: $\mu_B > 0$ (excess of baryons over antibaryons) encounters the sign problem $\det M[U, \mu] = |\det M[U, \mu]| e^{i\theta}$

The challenge of discovering the QCD critical point

QCD critical point







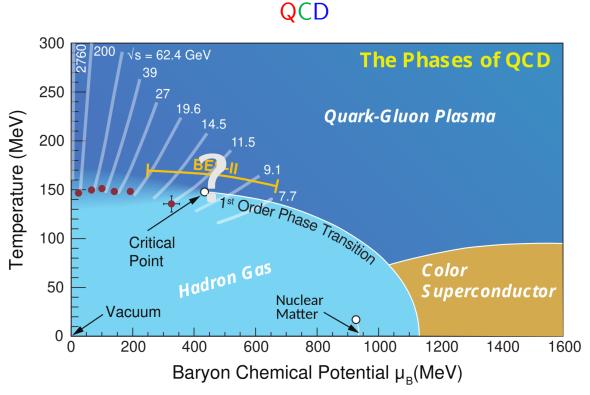


Figure from Bzdak et al., Phys. Rept. '20 and 2015 Nuclear Long Range Plan

What is the nature of the quark-hadron transition at finite baryon density?

Is there a QCD phase transition and critical point? Where?

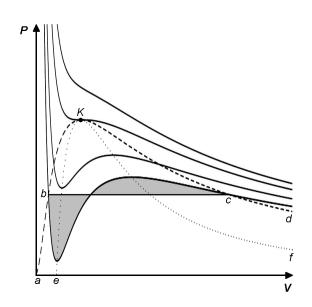
Lattice QCD: sign problem prevents simulations at non-zero baryon density

Heavy-ion collisions: access to finite density but might be too short-lived to observe a signal

Extrapolating critical point from lattice QCD

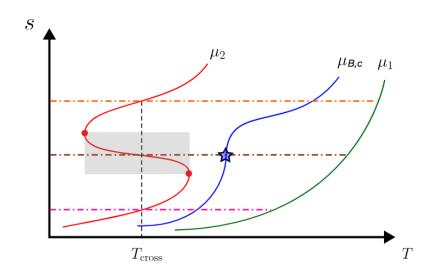


van der Waals (1873)





change the variables



Critical Point:

$$\left(\frac{\partial P}{\partial \rho_B}\right)_T = 0, \qquad \left(\frac{\partial^2 P}{\partial \rho_B^2}\right)_T = 0.$$

$$\left(\frac{\partial T}{\partial s}\right)_{\mu_B} = 0, \qquad \left(\frac{\partial^2 T}{\partial s^2}\right)_{\mu_B} = 0.$$

.16026 EX

Extrapolate from $\mu_B = 0!$

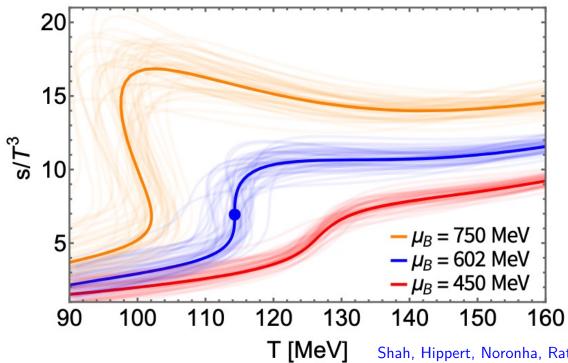
Looking for entropy crossings



Critical point ruled out (2 σ level) at $\mu_B < 400$ MeV

Borsanyi et al., arXiv:2502.10267

Try going further



Expansion around $\mu_B = 0$

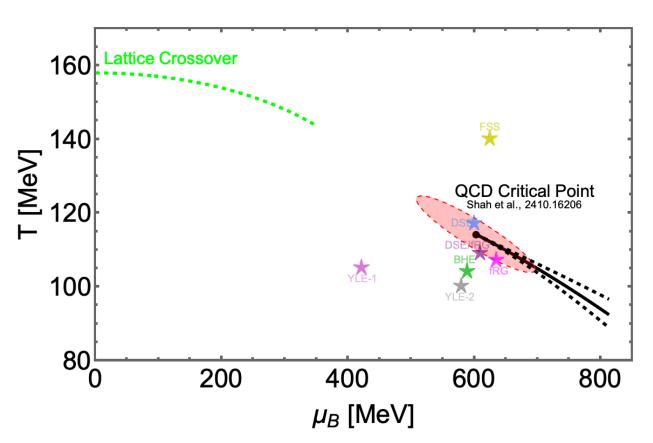
$$T_s(\mu_B; T_0) = T_0 + \alpha_2(T_0) \frac{\mu_B^2}{2}$$

Shah, Hippert, Noronha, Ratti, VV, arXiv:2410.16026

First-order phase transition emerges at $\mu_B > 600 \text{ MeV}$

QCD critical point estimates





Critical point estimate at $O(\mu_B^2)$:

 $T_c = 114 \pm 7 \text{ MeV}, \quad \mu_B = 602 \pm 62 \text{ MeV}$

Estimates from recent literature:

YLE-1: D.A. Clarke et al. (Bielefeld-Parma), arXiv:2405.10196

YLE-2: G. Basar, PRC 110, 015203 (2024)

BHE: M. Hippert et al., arXiv:2309.00579

fRG: W-J. Fu et al., PRD 101, 054032 (2020)

DSE/fRG: Gao, Pawlowski., PLB 820, 136584 (2021)

DSE: P.J. Gunkel et al., PRD 104, 052022 (2021)

FSS: A. Sorensen et al., arXiv:2405.10278

Optimist's view: Different estimates converge onto the same region because QCD CP is likely there

Pessimist's view: Different estimates converge onto the same region because it's the closest not yet ruled out by LQCD

Can be tested in laboratory with heavy-ion collisions

Critical point and heavy-ion collisions

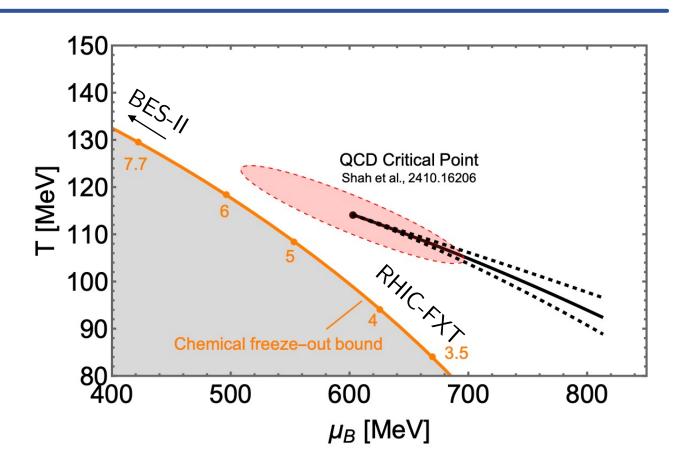


Control parameters

- Collision energy $\sqrt{s_{NN}} = 2.4 5020 \text{ GeV}$
 - Scan the QCD phase diagram
- Size of the collision region
 - Expect stronger signal in larger systems

Measurements

Final hadron abundances and momentum distributions event-by-event



Chemical freeze-out curve and CP

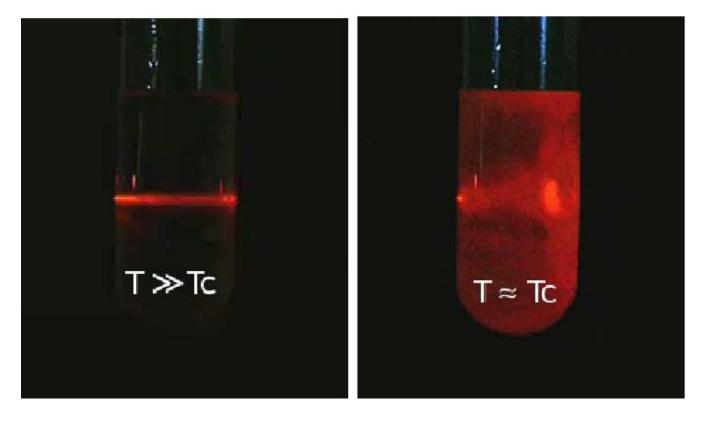
- Sets the **lower bound** on the temperature of the CP [Lysenko, Poberezhnyuk, Gorenstein, VV, arXiv:2408.06473]
- Caveats: strangeness neutrality ($\mu_S \neq 0$), uncertainty in the freeze-out curve
- CP may be close to freeze-out at $\sqrt{s_{NN}} \sim 3.5 5$ GeV

Critical point and fluctuations



Density fluctuations at macroscopic length scales

Critical opalescence



Unfortunately, we cannot do this in heavy-ion collisions

Event-by-event fluctuations and statistical mechanics



Consider a fluctuating number N

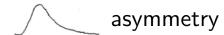
Cumulants:
$$G_N(t) = \ln \langle e^{tN} \rangle = \sum_{n=1}^{\infty} \kappa_n \frac{t^n}{n!}$$

variance
$$\kappa_2 = \langle (\Delta N)^2 \rangle = \sigma^2$$

skewness
$$\kappa_3 = \langle (\Delta N)^3 \rangle$$

kurtosis
$$\kappa_4 = \langle (\Delta N)^4 \rangle - 3 \langle (\Delta N^2) \rangle^2$$

width



Experiment:

$$P(N) \sim rac{N_{events}(N)}{N_{events}^{total}}$$

Statistical mechanics:

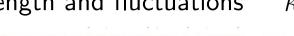
Grand partition function In
$$Z^{\rm gce}(T,V,\mu) = \ln \left[\sum_N e^{\mu N} Z^{\rm ce}(T,V,N) \right], \qquad \kappa_n \propto \frac{\partial^n (\ln Z^{\rm gce})}{\partial (\mu_N)^n}$$

$$\kappa_n \propto rac{\partial^n (\ln\,Z^{
m gce})}{\partial (\mu_N)^n}$$

Example: (Nuclear) Liquid-gas transition



(QCD) critical point: large correlation length and fluctuations





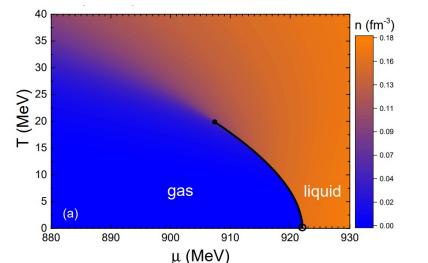
 $\kappa\sigma^2$

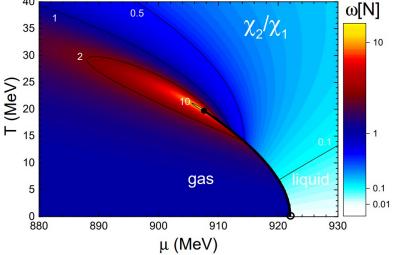
liquid

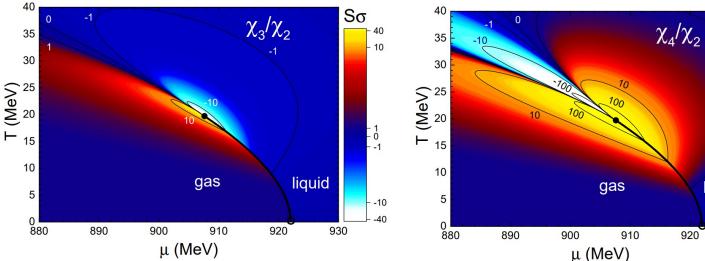
930

$$\kappa_2 \sim \xi^2, \quad \kappa_3 \sim \xi^{4.5}, \quad \kappa_4 \sim \xi^7$$
 w[N] $\qquad \qquad \xi \to \infty$

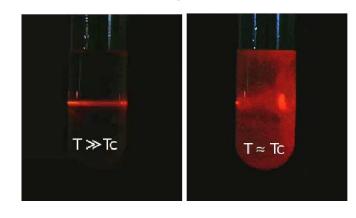
M. Stephanov, PRL '09, '11







Critical opalescence



$$\langle N^2 \rangle - \langle N \rangle^2 \sim \langle N \rangle \sim 10^{23}$$

in equilibrium

VV, Anchishkin, Gorenstein, Poberezhnyuk, PRC 92, 054901 (2015)

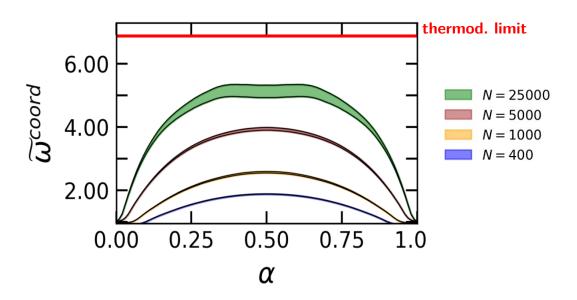
Example: Critical fluctuations in microscopic simulation

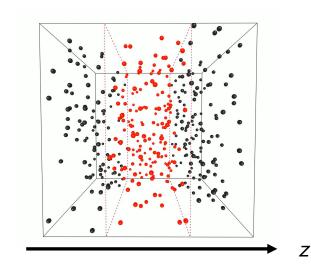


V. Kuznietsov (grad student @UH) et al., Phys. Rev. C 105, 044903 (2022)

Instead of observing system macroscopically, track each single particle

Classical molecular dynamics simulations of the **Lennard-Jones fluid** near critical point $(T \approx 1.06T_c, n \approx n_c)$ of the liquid-gas transition

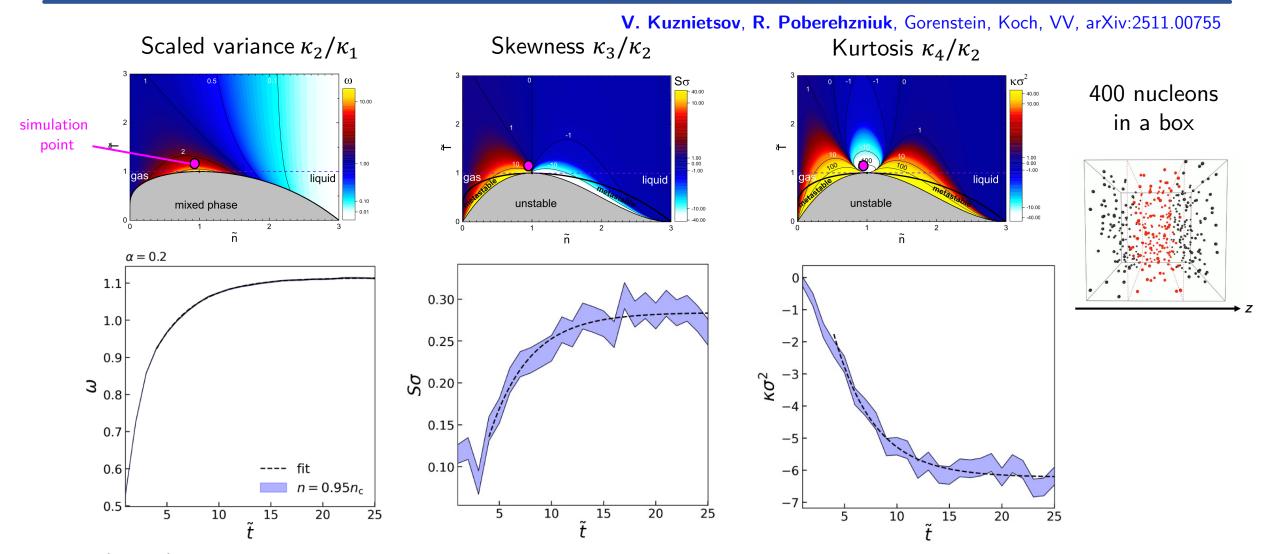




Large fluctuations survive despite strong finite-size effects and are large as advertised near the critical point

Non-Gaussian fluctuations from molecular dynamics



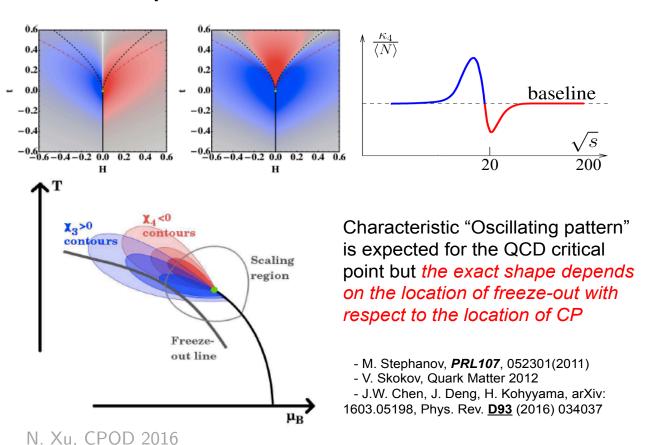


• (Non-)Gaussian cumulants equilibrate on comparable time scales

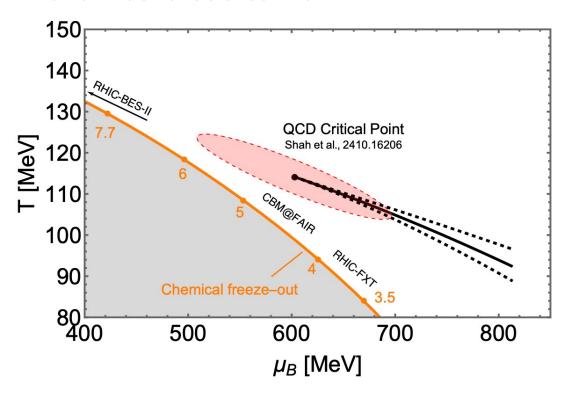
Equilibrium Expectations and Beam Energy Scan



Expectation from Calculations



Compare recent CP estimates and the freeze-out curve



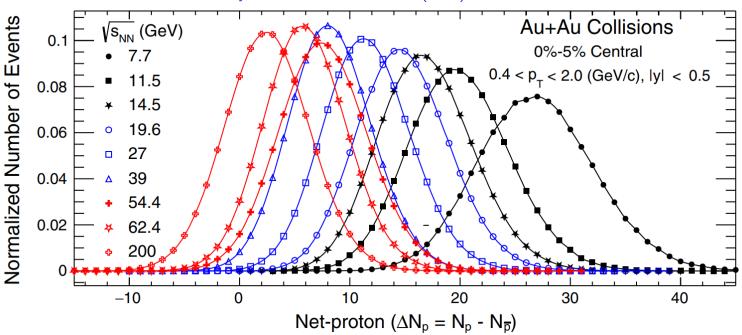
One of primary motivations for beam energy scan (BES) programs at RHIC BES-I (7.7-200 GeV) and BES-II (3-4.5 & 7.7-39 GeV)

Measuring cumulants in heavy-ion collisions



Count the number of events with given number of e.g. (net) protons

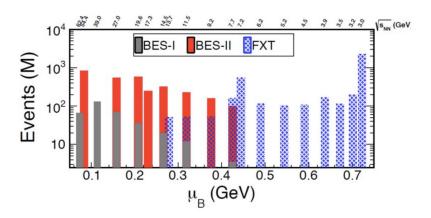
STAR Collaboration, Phys. Rev. Lett. 126, 092301 (2021)



Cumulants are extensive, $\kappa_n \sim V$, use ratios to cancel out the volume

$$\frac{\kappa_2}{\langle N \rangle}$$
, $\frac{\kappa_3}{\kappa_2}$, $\frac{\kappa_4}{\kappa_2}$

$$P(\Delta N_p) \sim rac{N_{
m events}(\Delta N_p)}{N_{events}^{total}}$$



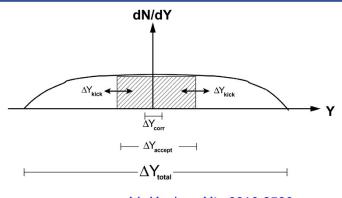
Statistics-hungry observables

Look for subtle critical point signals

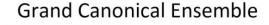
Statistical ensemble in HICs is neither CE or GCE

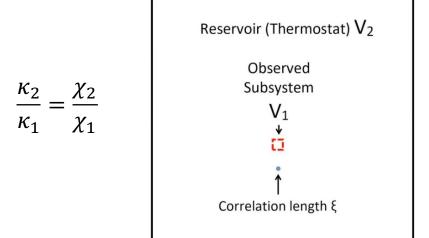


- Experimental measurements apply momentum cuts
 - The idea is to mimick GCE conditions
- However, in reality, the measured subsystem is of comparable size to total system where baryon number does not fluctuate and the canonical ensemble (CE) applies

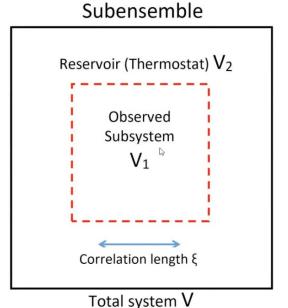


V. Koch, arXiv:0810.2520





VS $\alpha \equiv V_1/V \\ (0 \ll \alpha \ll 1)$



$$\frac{\kappa_2}{\kappa_1} = (1 - \alpha) \frac{\chi_2}{\chi_1}$$

• Statistical ensemble relevant for heavy-ion collisions is something "in-between" GCE and CE

Theory vs experiment



guidance from theory (e.g. lattice) experiment (the real thing) m_s/m_l=20 (open) 27 (filled) χ_4^B/χ_2^B 240 260 T [MeV] 3) (3) C₄/C₂ canonical effects efficiency correction kinematic cuts աստատա 0.5 CBWC purity ····· Hydro (cons) 10 20 100 200

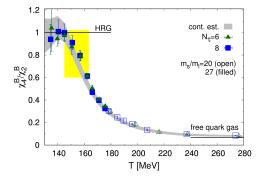
Collision Energy $\sqrt{s_{NN}}$ (GeV)

Theory vs experiment



guidance from theory (e.g. lattice)

experiment (the real thing)

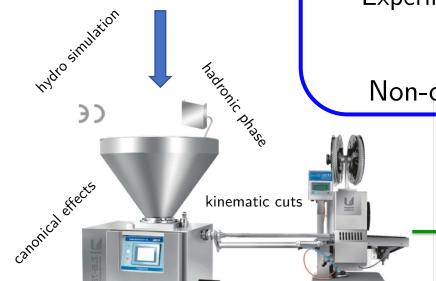


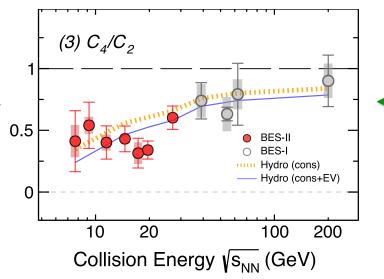
This was done in [VV, V. Koch, C. Shen, Phys. Rev. C 105, 014904 (2022)]

- Full hydro simulation
- Lattice QCD-like baryon susceptibilities (interacting HRG)
- Global baryon conservation (SAM)
- Experimental kinematic cuts



Non-critical baseline (hydro EV) prediction







Hints from RHIC-BES-I

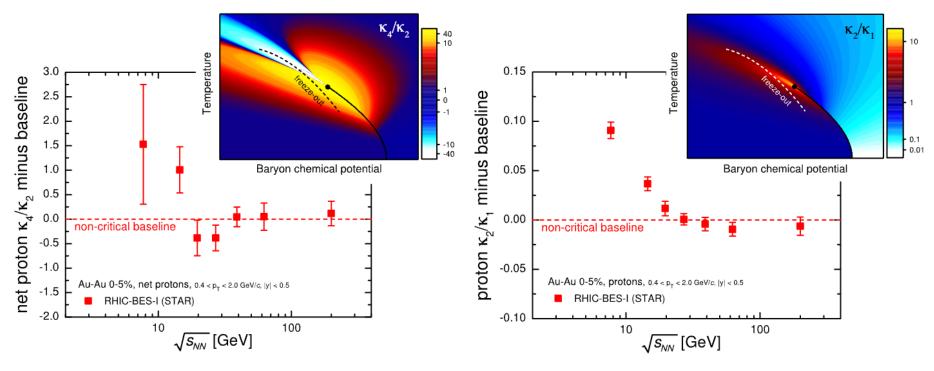


Quantitative calculations of critical fluctuations are still not available

State-of-the-art non-critical baseline computed using hydrodynamics

VV, V. Koch, C. Shen, Phys. Rev. C 105, 014904 (2022)

Subtract it from the data and look for a possible signal of CP



Analysis of RHIC-BES-II data in progress

First data from RHIC-BES-II



PHYSICAL REVIEW LETTERS 135, 142301 (2025)

Editors' Suggestion

Featured in Physics

Precision Measurement of Net-Proton-Number Fluctuations in Au + Au Collisions at RHIC

(STAR Collaboration)

(Received 26 March 2025; accepted 18 July 2025; published 29 September 2025)

We report precision measurements on cumulants (C_n) and factorial cumulants (κ_n) of (net) proton number distributions up to fourth order in Au + Au collisions over center-of-mass energies $\sqrt{s_{NN}} = 7.7-27$ GeV from phase II of the Beam Energy Scan program at RHIC. (Anti)protons are selected

(7) SEPTEMBER 29, 2025

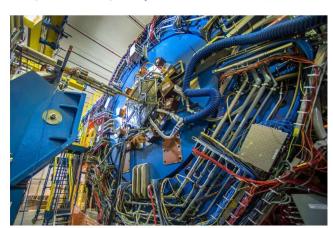


Editors' notes

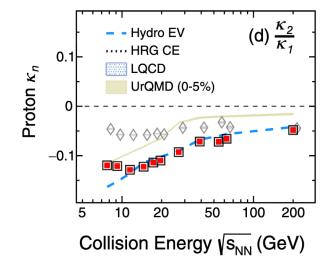
High-order analysis reveals more signs of phase-change 'turbulence' in nuclear matter

by Lawrence Berkeley National Laboratory

edited by Lisa Lock, reviewed by Robert Egan



The STAR detector at the U.S. Department of Energy's Brookhaven National C...



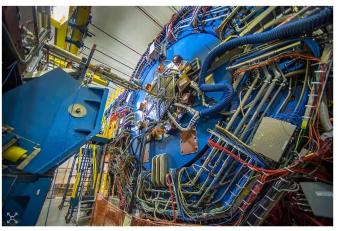
IOP Publishing

physicsworld

PARTICLE AND NUCLEAR | RESEARCH UPDATE

Hints of a boundary between phases of nuclear matter found at RHIC

09 Oct 2025



STAR at RHIC The experiment at Brookhaven National Laboratory has revealed hints of a critical point in the phase transition of nuclear matter. (Courtesy: BNL)

In a major advance for nuclear physics, scientists on the <u>STAR Detector</u> at the Relativistic Heavy Ion Collider (RHIC) in the US have spotted subtle but striking fluctuations in the number

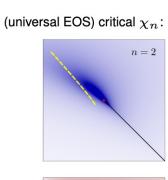
Meanwhile, theorists are racing to catch up. "The ball now moves largely to theory's court," Vovchenko says. He emphasizes the need for "quantitative predictions across energies and cumulants of various order that are appropriate for apples-to-apples comparisons with these data."

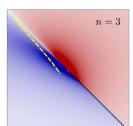
Factorial cumulants from RHIC-BES-II

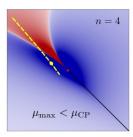


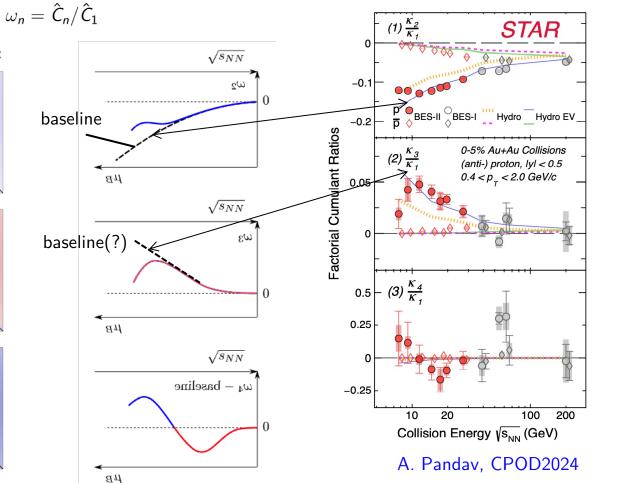
From M. Stephanov, SQM2024 & arXiv:2410.02861

STAR data:









baseline (hydro EV):

VV, V. Koch, C. Shen, PRC 105, 014904 (2022)

- describes right side of the peak in \hat{C}_3
- signal relative to baseline:
 - positive $\hat{C}_2 > 0$
 - negative $\hat{C}_3 < 0$

Conclusion:

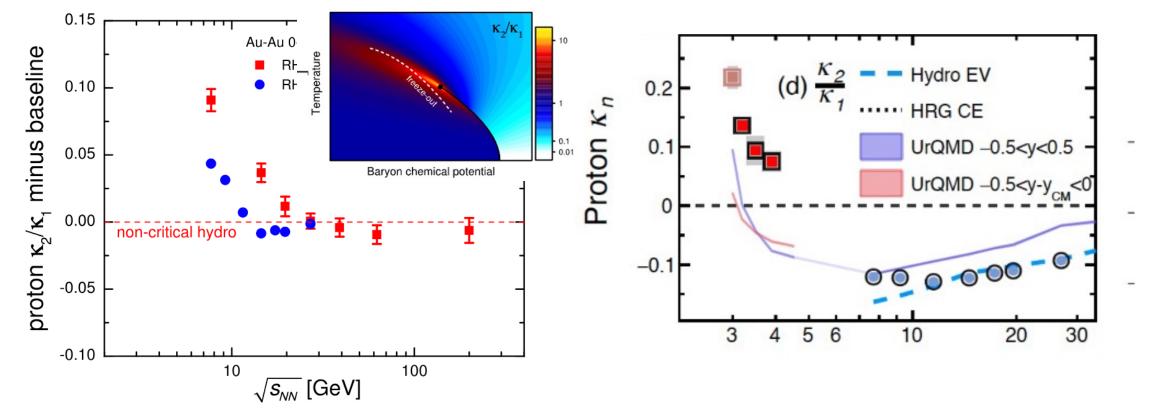
Controlling the non-critical baseline is essential

Bzdak et al review 1906.00936

Expected signatures: bump in ω_2 and ω_3 , dip then bump in ω_4 for CP at $\mu_B > 420 \text{ MeV}$

Subtracting the baseline

If $\kappa_2^{tot} \approx \kappa_2^{crit} + \kappa_2^{reg}$ try to isolate the critical part* by subtracting the baseline (here hydro EV)



Enhancement relative to the baseline at lower $\sqrt{s_{NN}}$ which continues at fixed target energies

^{*}May be a useful quantity for finite-size scaling analysis compared to the bare κ_2/κ_1

Recently, attractive and repulsive interactions implemented through a potential in rapidity

$$E_r(y_1,y_2) = \alpha_r e^{-|y_1-y_2|/\rho_r} \qquad P(y_1,y_1) = \frac{e^{-E(y_1,y_2)}}{Z} \qquad E_a(y_1,y_2) = \alpha_a |y_1-y_2|^{\beta_a}$$

$$\frac{repulsion}{\sum_{k=0}^{N} e^{-k\theta_k} \text{ one-pulsive pulsive line.}}{\sum_{k=0}^{N} e^{-k\theta_k} \text{ one-pulsive line.}} \qquad \frac{e^{-E(y_1,y_2)}}{Z} \qquad E_a(y_1,y_2) = \alpha_a |y_1-y_2|^{\beta_a}$$

$$\frac{repulsion}{\sum_{k=0}^{N} e^{-k\theta_k} \text{ one-pulsive line.}} \qquad \frac{e^{-E(y_1,y_2)}}{Z} \qquad E_a(y_1,y_2) = \alpha_a |y_1-y_2|^{\beta_a}$$

$$\frac{e^{-E(y_1,y_2)}}{Z} \qquad E_a(y_1,y_2) = \alpha_a |y_1-y_2|^{\beta_a}$$

$$\frac{e^{-E(y_1,y_2)}}{Z} \qquad \frac{e^{-E(y_1,y_2)}}{Z} \qquad \frac{e^{-E(y_1,y_$$

Interplay of repulsive (high $\sqrt{s_{NN}}$) and attractive (low $\sqrt{s_{NN}}$) interactions?

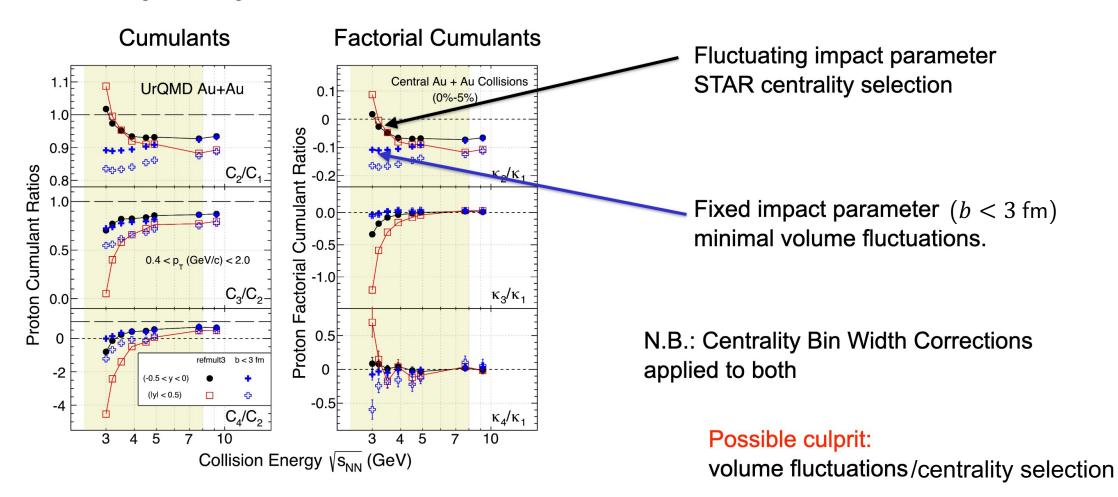
Time to celebrate!



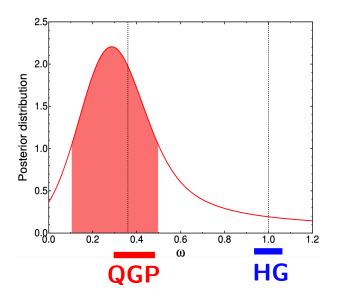
Adapted from V. Koch, ERICE2025

The (possible) culprit

X. Zhang, Y. Zhang, X. Luo, N. Xu, arXiv: 2506.18832



Other challenges: Antiproton puzzle (not described by hydro)



Charge fluctuations as a signature of QGP

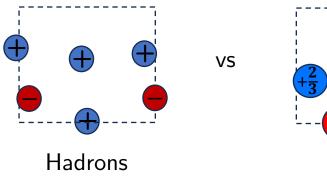
• J. Parra, R. Poberezhniuk, V. Koch, C. Ratti, VV, arXiv:2504.02085 (Phys. Rev. Lett., in print)

Charge fluctuations

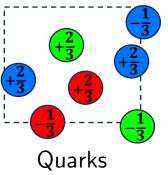


An old idea: Hadrons carry integer electric charges, quarks carry fractional electric charges.

Koch, Jeon, PRL (2000); Asakawa, Muller, Heinz, PRL (2000)



HG: large fluctuations



QGP: small fluctuations

Fluctuations depend on the square of the charges and are smaller in the QGP

Quantified by:

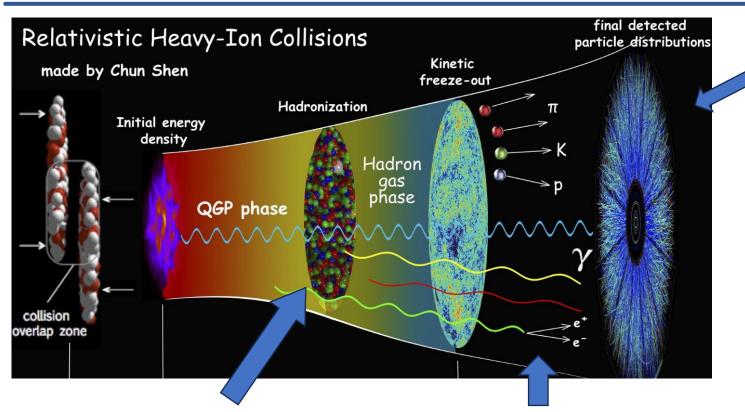
$$D = 4 \frac{\kappa_2 [N_+ - N_-]}{\langle N_{\rm ch} \rangle} = 4 \frac{\kappa_2 [Q]}{\langle Q^+ + Q^- \rangle}$$

Naïve grand canonical ensemble (GCE) expectations:

- $D_{HG} \approx 2.8 4$
- $D_{OGP} \approx 1 1.5$

Charge fluctuations: stages

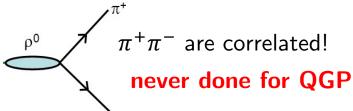




1. Fluctuations at hadronization (primordial charges)

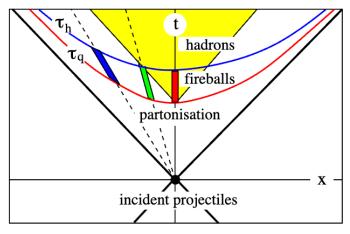
 ω — Distinguishes hadron gas ($\omega \approx 1$) from QGP ($\omega \approx 0.25 - 0.40$)

2. Resonance decays



4. Kinematical cuts never done for QGP

3. (Local) charge conservation



Castorina, Satz, IJMPE '14

never done both for HG and QGP

Charge fluctuations: hadronization



Define ω as a measure of charge fluctuations at hadronization

$$\omega = rac{\kappa_2[Q]}{\langle N_{
m ch}^{
m prim}
angle}$$
 charged multiplicity

Hadron gas: $\omega_{HG} \approx 1$ (Poisson statistics)

Free QGP: $\omega_{QGP} \approx 0.36$ (Stefan-Boltzmann limit)

More generally:

$$\omega = \frac{\kappa_2[Q]}{\langle N_{\rm ch}^{\rm prim} \rangle} = \frac{V T^3 \chi_2^Q}{S} \frac{S}{\langle N_{\rm ch}^{\rm prim} \rangle}$$

$$= egin{array}{c} \chi_2^Q \ \hline s/T^3 \end{array} egin{array}{c} S \ \hline \langle N_{
m ch}
angle \end{array} egin{array}{c} \langle N_{
m ch}
angle \end{array}$$

 $\gamma_Q \approx 1.67$ (decays)

from thermal model

The EoS
e.g. lattice QCD



ω from lattice QCD 1.2 1.0 hadron gas 0.8 0.4 0.2 0.0 150 200 250 300 350 400 T[MeV]

P. Hanus, A. Mazeliauskas, K. Reygers, PRC (2019)

Charge fluctuations: other stages



2. Hadronic phase (resonance decays)

- Decays are local and conserve charge but increase charged multiplicity, $\langle N_{ch} \rangle = \langle N_{ch}^{prim} \rangle \gamma_Q$, where $\gamma_Q \approx 1.67$
- Total charge susceptibility does not change, $\chi_2^{Q,final} = \chi_2^{Q,prim}$, but balance between self-correlation and two-particle correlations does

3. Local charge conservation [VV, PRC 110, L061902 (2024)]

- 2-point charge density correlator with a balancing term
- Local charge conservation introduced through modulation of the balancing term

$$\mathcal{C}_2^Q(\mathbf{r}_1,\mathbf{r}_2) \equiv \langle \delta
ho_Q(\mathbf{r}_1) \delta
ho_Q(\mathbf{r}_2)
angle$$

$$\mathcal{C}_2^Q(\mathbf{r}_1,\mathbf{r}_2) = \chi_2^Q \left[\delta(\mathbf{r}_1 - \mathbf{r}_2) - \frac{\varkappa(\mathbf{r}_1,\mathbf{r}_2)}{V_{\rm tot}} \right] \qquad \mathbf{r} = \eta \qquad \qquad \varkappa(\eta_1,\eta_2) \propto \exp\left[-\frac{(\eta_1 - \eta_2)^2}{2\sigma_\eta^2} \right]$$
local correlation balancing contribution local charge conservation

4. Kinematical cuts

$$\kappa_2[Q]|_{|\eta|<\eta_{\text{cut}}} \propto \int_{-\eta_{\text{cut}}/2}^{\eta_{\text{cut}}/2} d\eta_1 \int_{-\eta_{\text{cut}}/2}^{\eta_{\text{cut}}/2} d\eta_2 \mathcal{C}_2^Q(\eta_1,\eta_2) p(\eta_1) p(\eta_2)$$

Acceptance probability $p(\eta)$ from blast-wave model

Putting everything together



$$D = 4 \left\{ 1 - \left(1 - \frac{\omega}{\gamma_Q} \right) \frac{\langle p^2(\eta) \rangle}{\langle p(\eta) \rangle} - \frac{\omega}{\gamma_Q} \frac{\langle p(\eta_1) p(\eta_2) \rangle_{\varkappa}}{\langle p(\eta) \rangle} \right\}$$

$$= 4 \left\{ 1 - \left(1 - \frac{\omega}{\gamma_Q} \right) \frac{\langle p(\eta) \rangle}{\langle p(\eta) \rangle} - \frac{\omega}{\gamma_Q} \frac{\langle p(\eta_1) p(\eta_2) \rangle_{\varkappa}}{\langle p(\eta) \rangle} \right\}$$

$$= 4 \left\{ 1 - \left(1 - \frac{\omega}{\gamma_Q} \right) \frac{\langle p(\eta) \rangle}{\langle p(\eta) \rangle} - \frac{\omega}{\gamma_Q} \frac{\langle p(\eta_1) p(\eta_2) \rangle_{\varkappa}}{\langle p(\eta) \rangle} \right\}$$

$$= 4 \left\{ 1 - \left(1 - \frac{\omega}{\gamma_Q} \right) \frac{\langle p(\eta) \rangle}{\langle p(\eta) \rangle} - \frac{\omega}{\gamma_Q} \frac{\langle p(\eta_1) p(\eta_2) \rangle_{\varkappa}}{\langle p(\eta) \rangle} \right\}$$

$$= 4 \left\{ 1 - \left(1 - \frac{\omega}{\gamma_Q} \right) \frac{\langle p(\eta) \rangle}{\langle p(\eta) \rangle} - \frac{\omega}{\gamma_Q} \frac{\langle p(\eta_1) p(\eta_2) \rangle_{\varkappa}}{\langle p(\eta) \rangle} \right\}$$

$$= 4 \left\{ 1 - \left(1 - \frac{\omega}{\gamma_Q} \right) \frac{\langle p(\eta) \rangle}{\langle p(\eta) \rangle} - \frac{\omega}{\gamma_Q} \frac{\langle p(\eta) \rangle}{\langle p(\eta) \rangle} \right\}$$

$$\omega_{HG} = 1$$
 $\omega_{QGP} = 0.36$

 γ_Q - Resonance decays

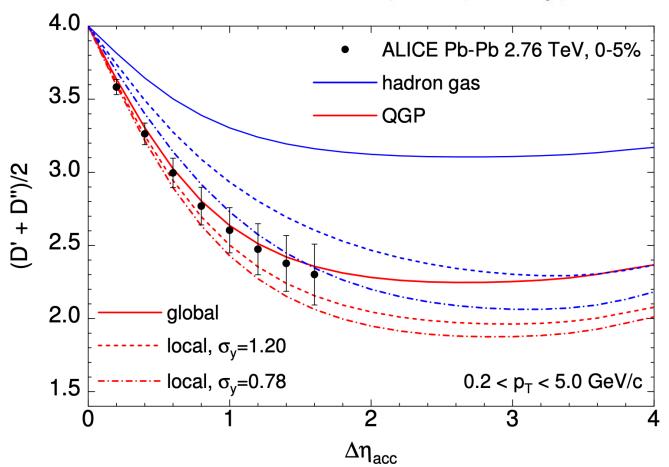
 $\langle p(\eta_1)p(\eta_2)
angle_{arkappa}$ - Pair acceptance weighted with Local Charge Conservation

$$\frac{\langle p^2(\eta) \rangle}{\langle p(\eta) \rangle}$$
 - Momentum Acceptance Cuts using p(η) from the blast-wave model

D-measure at LHC: comparison with experiment



$$D = 4 \left\{ 1 - \left(1 - \frac{\omega}{\gamma_Q} \right) \frac{\langle p^2(\eta) \rangle}{\langle p(\eta) \rangle} - \frac{\omega}{\gamma_Q} \frac{\langle p(\eta_1) p(\eta_2) \rangle_{\varkappa}}{\langle p(\eta) \rangle} \right\}$$



Parameters used:

$$\omega_{HG} = 1$$
 $\omega_{QGP} = 0.36$ $\gamma_Q = 1.67$

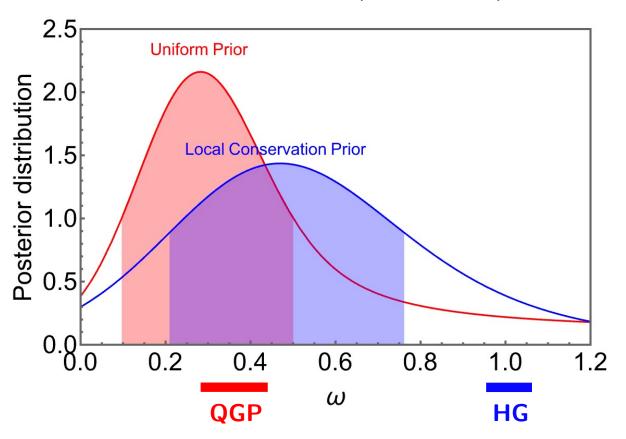
- Vary σ_y to accommodate global vs local charge conservation
 - Here values of σ_y are based on local baryon conservation estimates

VV, PRC 110, L061902 (2024)

D-measure at LHC: Bayesian analysis



Vary primordial fluctuation ω (HG vs QGP) and correlation volume V_C (local conservation) freely



Uniform prior

$$\omega \in U(0, 1.2), \quad V_c \in U(0, V_{tot}),$$

Bayes factor (QGP : HG) = 8.74

Local conservation prior

$$\omega \in U(0, 1.2), V_c \in \text{Gaussian at } (0.20 \pm 0.05) V_{tot}$$

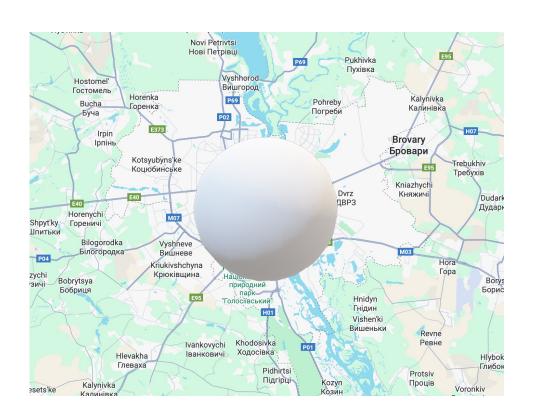
Bayes factor (QGP : HG) = 4.93

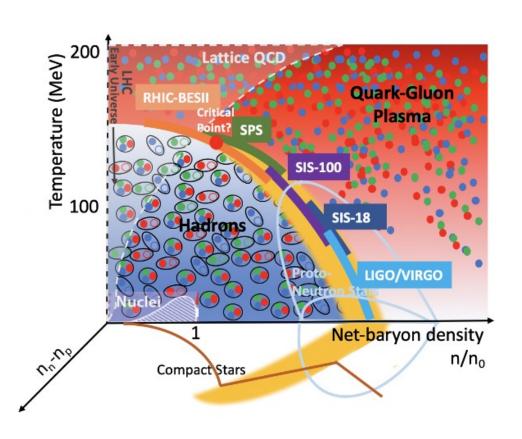
Moderate evidence for freeze-out of charge fluctuations in QGP

QCD in astrophysics



Neutron stars are extremely compact objects (1-2 solar masses confined to an 8-mile sphere)



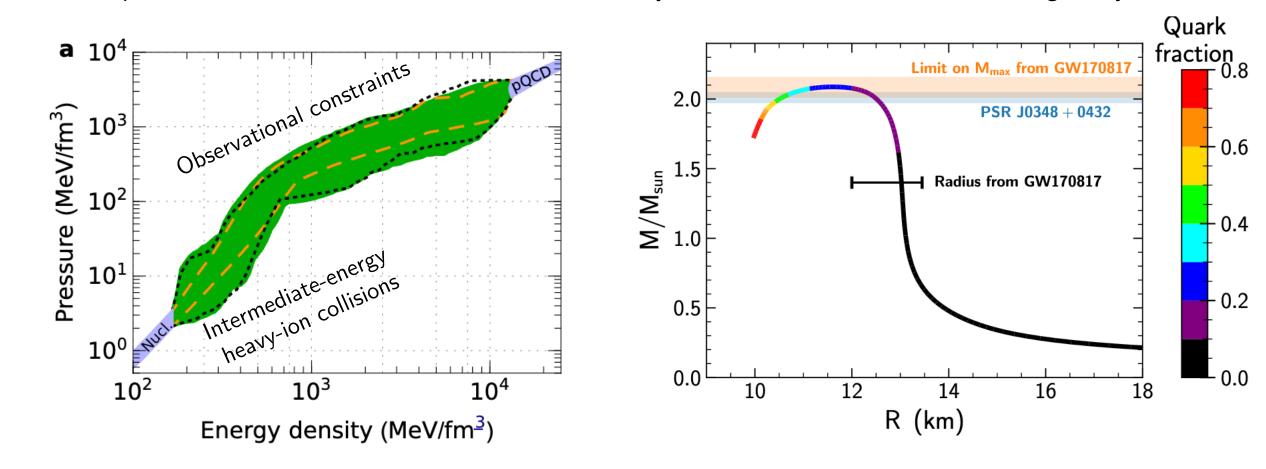


Pressure of dense nuclear matter balances the gravitational pull

QCD in astrophysics



Properties of dense nuclear matter define how heavy neutron stars can be and how large they are



Intermediate energy heavy-ion collisions probe same dense nuclear matter

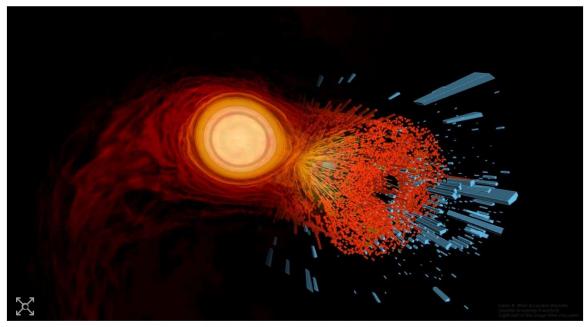
The ultimate "heavy-ion" collision



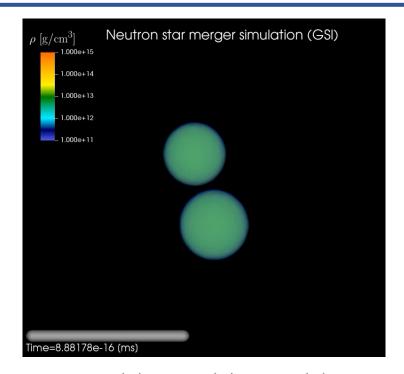
COSMOLOGY | RESEARCH UPDATE

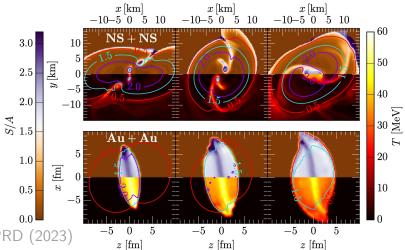
Gravitational waves from neutron-star mergers could reveal quark-gluon plasma

15 May 2020



Dark star crashes: the computer simulation of two merging neutron stars (left) blended with an image of heavy-ion collisions at CERN to highlight the connection of astrophysics with nuclear physics. Courtesy: Lukas R Weih and Luciano Rezzolla/Goethe University Frankfurt and CMS/CERN)





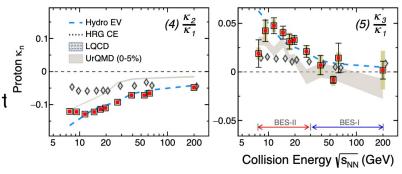
Summary

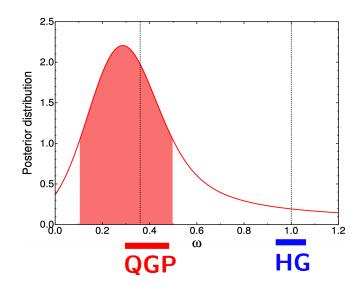


- Lattice-based constraints on the QCD critical point
 - Rule out QCD CP at $\mu_B < 450$ MeV at a 2-sigma level
 - New method based on contours of constant entropy places QCD CP at $T_c=114\pm7$ MeV, $\mu_B=602\pm62$ MeV



- Non-critical physics describe the proton data at $\sqrt{s_{NN}} \ge 20$ GeV
- $\hat{\mathcal{C}}_2 \hat{\mathcal{C}}_2^{baseline} > 0$ and $\hat{\mathcal{C}}_3 \hat{\mathcal{C}}_3^{baseline} > 0$ at $\sqrt{s_{NN}} < 10$ GeV
 - May indicate freeze-out of fluctuations on the QGP side of crossover
- Net-charge fluctuations at the LHC revisited
 - Moderate evidence for charge fluctuations freeze-out in the QGP





Outlook:

Improved description of non-critical baselines and quantitative predictions of critical fluctuations

Dynamical approaches to the QCD critical point search



- 1. Dynamical model calculations of critical fluctuations
- BEST

[X. An et al., Nucl. Phys. A 1017, 122343 (2022)]

- Fluctuating hydrodynamics (hydro+) and (non-equilibrium) evolution of fluctuations
- Equation of state with a tunable critical point [P. Parotto et al, PRC 101, 034901 (2020); J. Karthein et al., EPJ Plus 136, 621 (2021)]
- Generalized Cooper-Frye particlization [M. Pradeep, et al., PRD 106, 036017 (2022); PRL 130, 162301 (2023)]

Alternatives at high μ_B : hadronic transport/molecular dynamics with a critical point

[A. Sorensen, V. Koch, PRC 104, 034904 (2021); V. Kuznietsov et al., PRC 105, 044903 (2022)]

2. Deviations from precision calculations of non-critical fluctuations

- Non-critical baseline is not flat [Braun-Munzinger et al., NPA 1008, 122141 (2021)]
- Include essential non-critical contributions to (net-)proton number cumulants
- Exact baryon conservation + hadronic interactions (hard core repulsion)
- Based on realistic hydrodynamic simulations tuned to bulk data

[VV, C. Shen, V. Koch, Phys. Rev. C 105, 014904 (2022)]

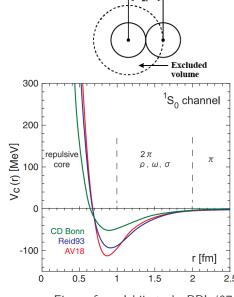


Figure from Ishii et al., PRL '07

Factorial cumulants $\hat{\mathcal{C}}_n$ vs ordinary cumulants \mathcal{C}_n



Factorial cumulants: ~irreducible n-particle correlations

$$\hat{C}_n \sim \langle N(N-1)(N-2)...\rangle_c$$

$$\hat{C}_1 = C_1$$
 $\hat{C}_2 = C_2 - C_1$
 $\hat{C}_3 = C_3 - 3C_2 + 2C_1$
 $\hat{C}_4 = C_4 - 6C_3 + 11C_2 - 6C_1$

Ordinary cumulants: mix correls. of different orders

$$egin{aligned} \mathcal{C}_n &\sim \langle \delta \mathcal{N}^n
angle_c \ &C_1 = \hat{C}_1 \ &C_2 = \hat{C}_2 + \hat{C}_1 \ &C_3 = \hat{C}_3 + 3\hat{C}_2 + \hat{C}_1 \end{aligned}$$

 $C_4 = \hat{C}_4 + 6\hat{C}_3 + 7\hat{C}_2 + \hat{C}_1$

[Bzdak, Koch, Strodthoff, PRC 95, 054906 (2017); Kitazawa, Luo, PRC 96, 024910 (2017); C. Pruneau, PRC 100, 034905 (2019)]

Factorial cumulants and different effects

- Baryon conservation
 [Bzdak, Koch, Skokov, EPJC '17]
- Excluded volume [VV et al, PLB '17]
- Volume fluctuations [Holzman et al., arXiv:2403.03598]
- Critical point
 [Ling, Stephanov, PRC '16]

 $\hat{C}_n^{\mathsf{cons}} \propto (\hat{C}_1)^n/\langle \mathcal{N}_{\mathsf{tot}}
angle^{n-1}$ small

$$\hat{C}_n^{\mathsf{EV}} \propto b^n$$
 small

• proton vs baryon $\hat{C}_n^B \sim 2^n \times \hat{C}_n^p$ same sign! [Kitazawa, Asakawa, PRC '12]

$$\hat{C}_n^{CF} \sim (\hat{C}_1)^n \kappa_n[V]$$
 depends on volume cumulants

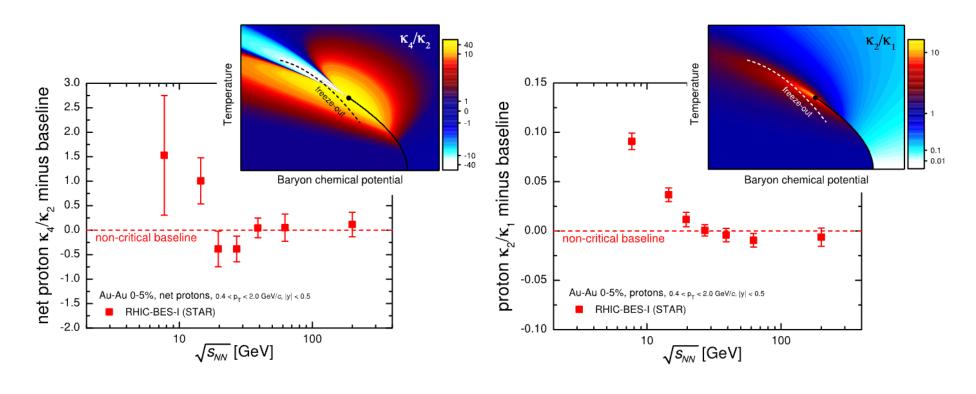
$$\hat{C}_2^{CP}\sim \xi^2$$
, $\hat{C}_3^{CP}\sim \xi^{4.5}$, $\hat{C}_4^{CP}\sim \xi^7$ large

Hints from RHIC-BES-I



VV, V. Koch, C. Shen, Phys. Rev. C 105, 014904 (2022)

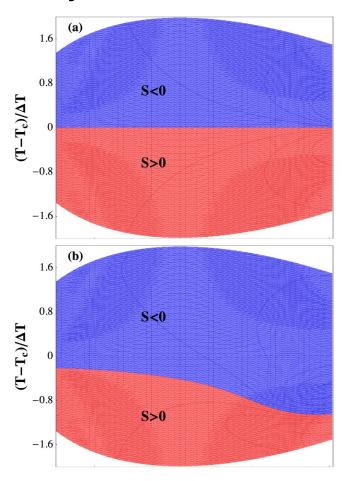
Subtracting the hydrodynamic non-critical baseline



Factorial cumulants from RHIC-BES-II and CP



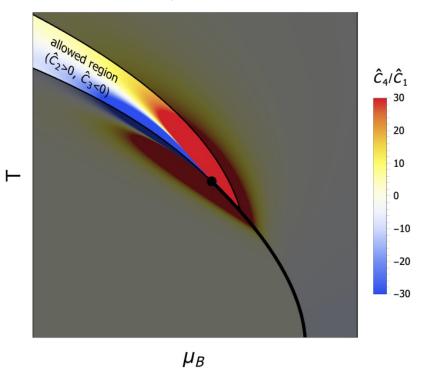
Memory effect



Mukherjee, Venugopalan, Yin, PRC 92, 034912 (2015)

Exclusion plots

Exclude $\hat{\mathcal{C}}_2 < 0 \& \hat{\mathcal{C}}_3 > 0$ regions on the phase diagram near CP

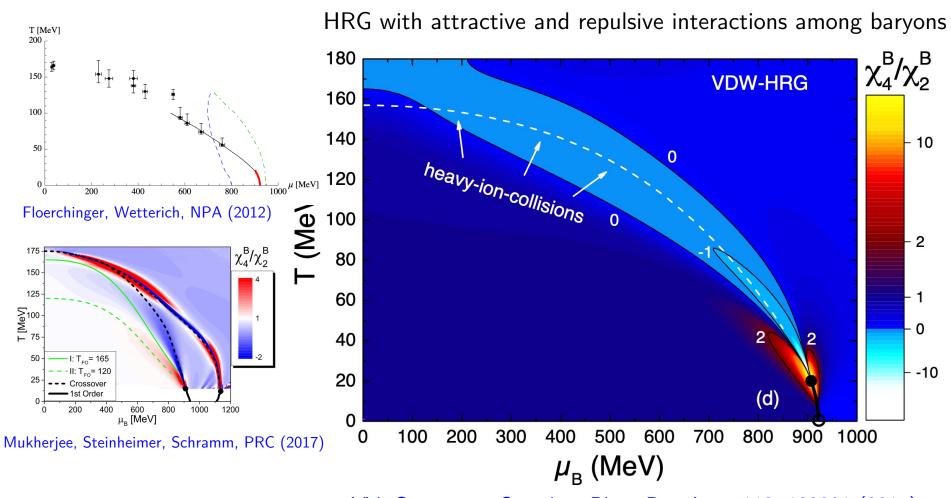


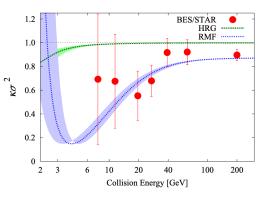
Adapted from Bzdak, Koch, Strodthoff, PRC 95, 054906 (2017) and based on the model from VV, Anchishkin, Gorenstein, Poberezhnyuk, PRC 92, 054901 (2015)

Freeze-out of fluctuations on the QGP side of the crossover?

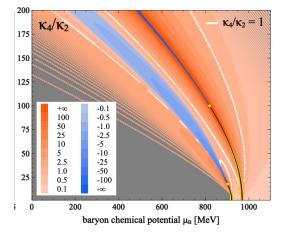
Interplay with nuclear liquid-gas transition







Fukushima, PRC (2014)

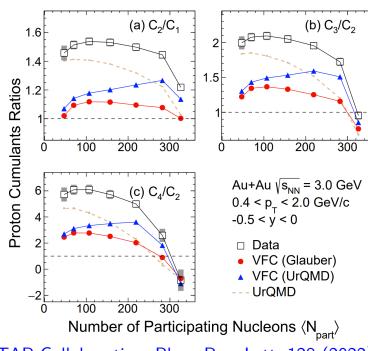


Sorensen, Koch, PRC (2020)

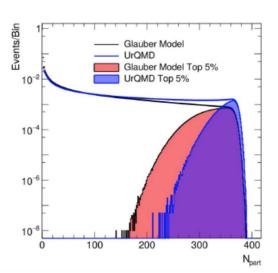
VV, Gorenstein, Stoecker, Phys. Rev. Lett. 118, 182301 (2017)

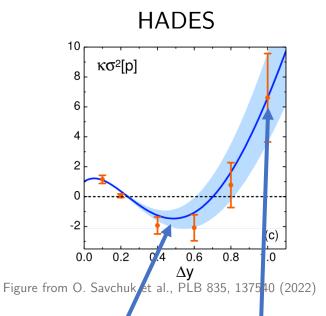
Lower energies $\sqrt{s_{NN}} \le 7.7$ GeV





STAR-FXT





STAR Collaboration, Phys. Rev. Lett. 128 (2022) 202303

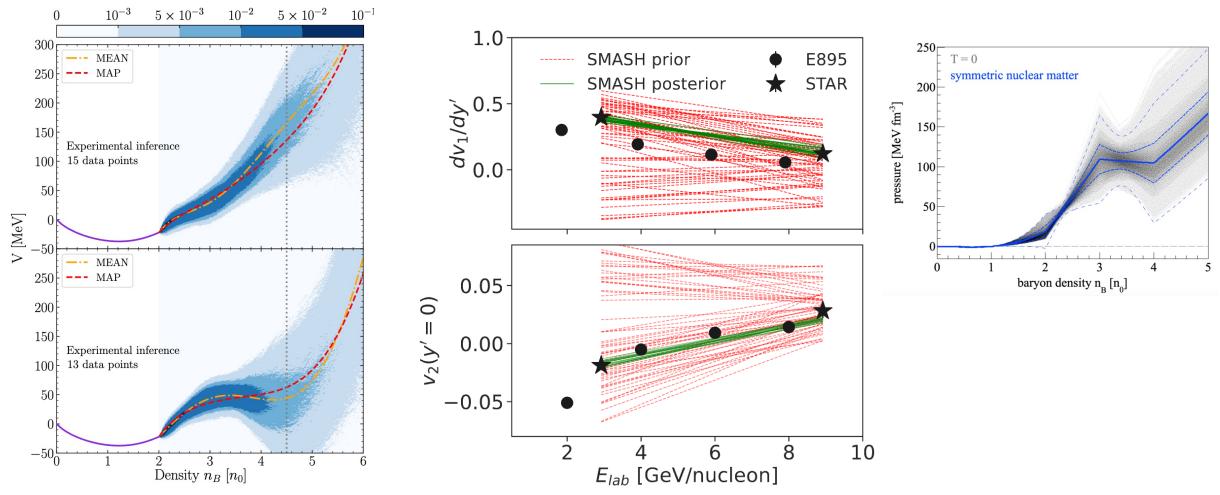
- Volume fluctuations/centrality selection appear to play an important role
 - UrQMD is useful for understanding basic systematics associated with it
- Indications for enhanced scaled variance, $\kappa_2/\kappa_1 > 1$
- κ_4/κ_2 negative and described by UrQMD (purely hadronic?), note -0.5<y<0 instead of |y|<0.5

Proper understanding of $\kappa_2/\kappa_1>1$ in both HADES and STAR-FXT is missing

Dense matter EoS from flow measurements



- Use hadronic transport (UrQMD and SMASH) with adjustable mean field to use a flexible EoS
- Extract the EoS from proton flow measurements



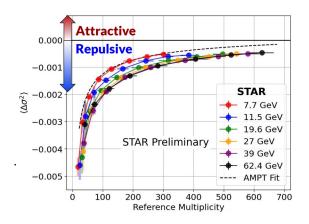
M. Kuttan, Steinheimer, Zhou, Stoecker, PRL 131, 202303 (2023)

Oliinychenko, Sorensen, Koch, McLerran, PRC 108, 034908 (2023)

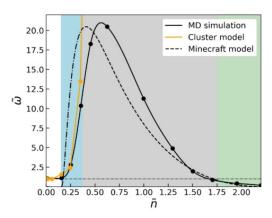
Other observables

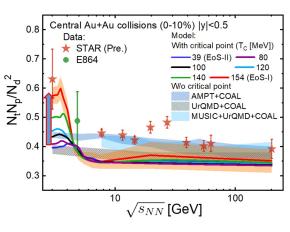


- Azimuthal correlations of protons
 - points to repulsion at RHIC-BES



- Light nuclei
 - Spinodal/critical point enhancement of density fluctuations and light nuclei production



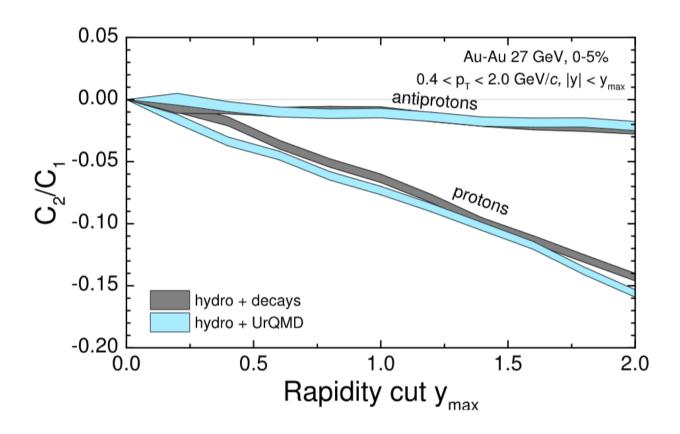


- Proton intermittency
 - No structure indicating power-law seen by NA61/SHINE
- Directed flow, speed of sound

Effect of the hadronic phase

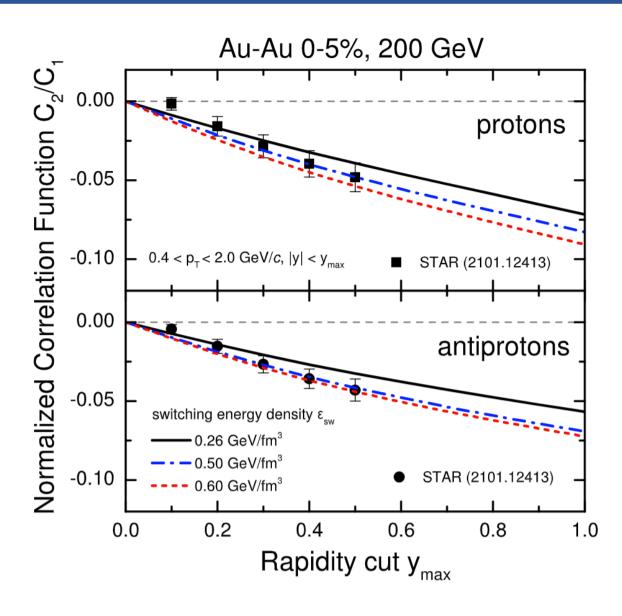


Sample ideal HRG model at particlization with exact conservation of baryon number using Thermal-FIST and run through hadronic afterburner UrQMD



Dependence on the switching energy density



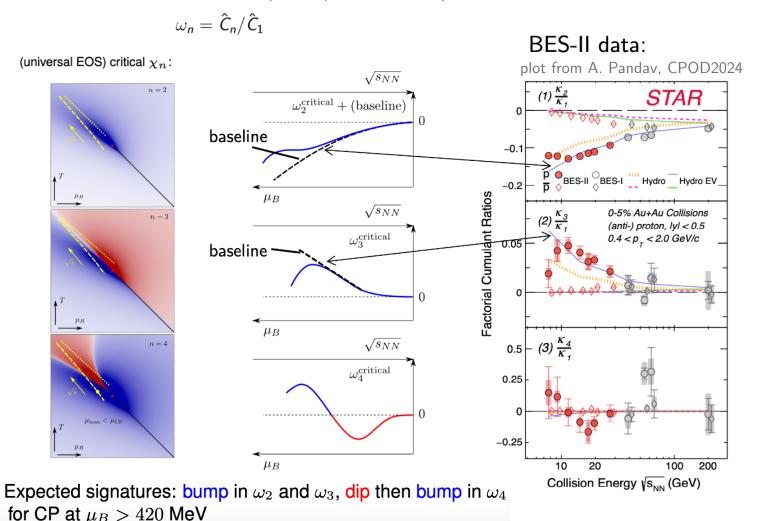


Summary slides

Understanding proton cumulants from RHIC-BES-II data



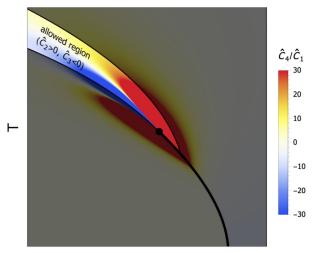
Vovchenko, Koch, arXiv:2504.01368, plot adapted from M. Stephanov, arXiv:2410.02861



- Factorial cumulants may be more instructive than ordinary cumulants for CP search
- signal relative to baseline:
 - positive $\hat{C}_2 \hat{C}_2^{baseline} > 0$
 - negative $\hat{C}_3 \hat{C}_3^{baseline} < 0$



allowed region near the CP is on QGP side of crossover



 μ_B

Scaled factorial cumulants and the antiproton puzzle



Bzdak et al. introduced reduced correlation functions — "couplings" [Bzdak, Koch, Strodthoff, PRC 95, 054906 (2017)]

$$\hat{c}_k = \frac{\hat{c}_k}{\langle N \rangle^k}$$

$$c_k = \frac{\int \rho_1(y_1) \cdots \rho_1(y_k) c_k(y_1, \dots, y_k) dy_1 \cdots dy_k}{\int \rho_1(y_1) \cdots \rho_1(y_k) dy_1 \cdots dy_k}$$

integrated correlation function in rapidity

Long-range correlations lead to acceptance-independent \hat{c}_k , including any combination of global baryon conservation and volume fluctuations (no need for CBWC)!

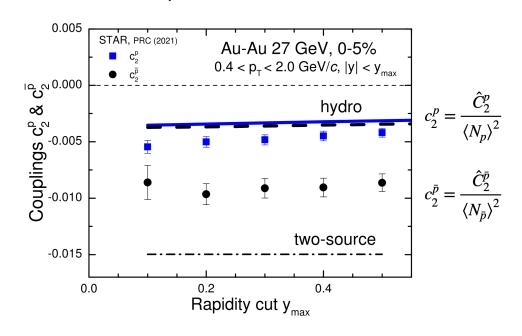
Standard hydro predicts:

$$\hat{c}_2^p \approx \hat{c}_2^{\bar{p}} = const.$$
 at a given $\sqrt{s_{NN}}$

Experiment:

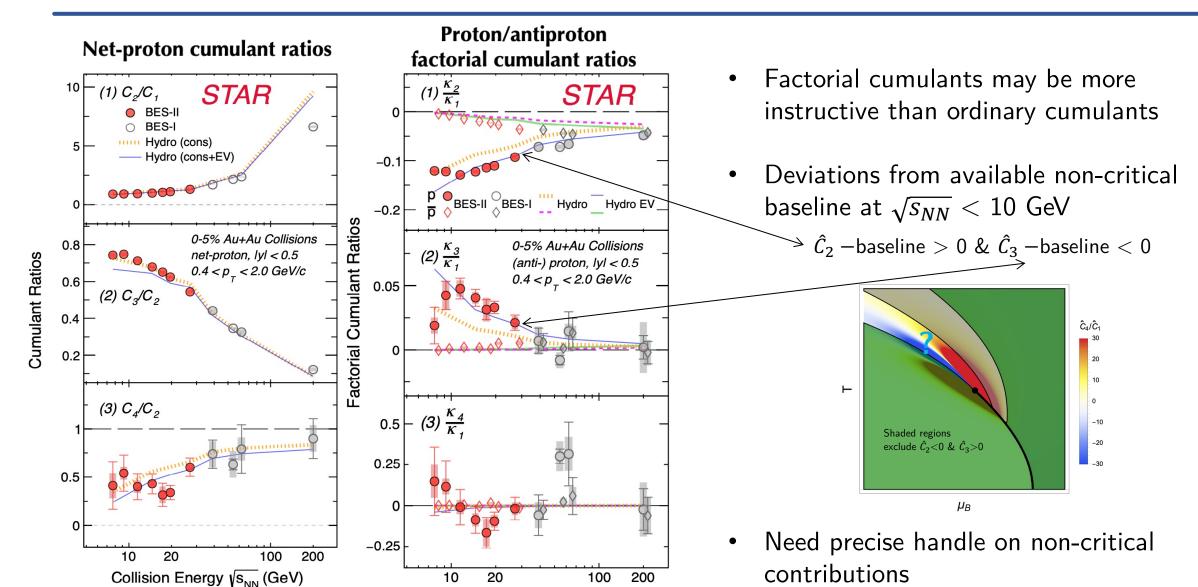
$$|\hat{c}_2^p| < |\hat{c}_2^p|$$
 the antiproton puzzle

Possible explanation: stopped and produced matter do not thermalize (no single fluid)



Understanding proton cumulants from RHIC-BES-II data

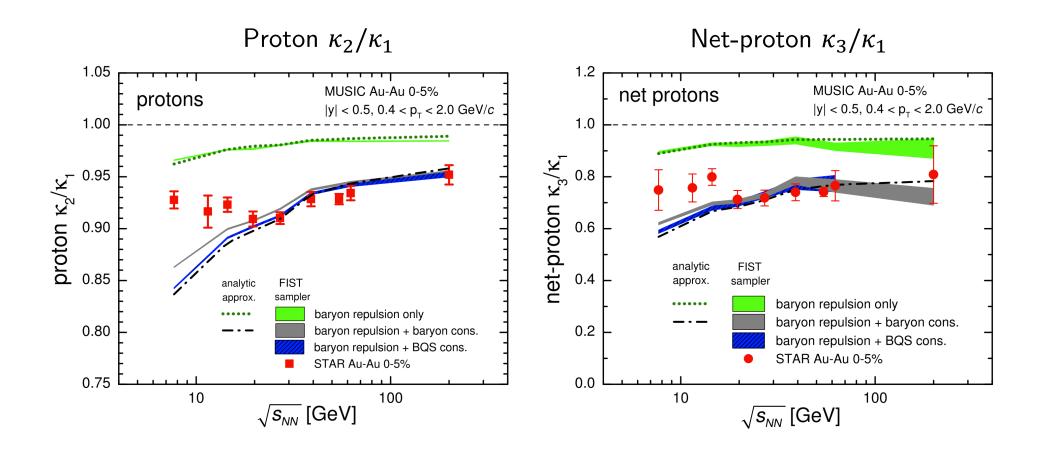




Collision Energy √s_{NN} (GeV)

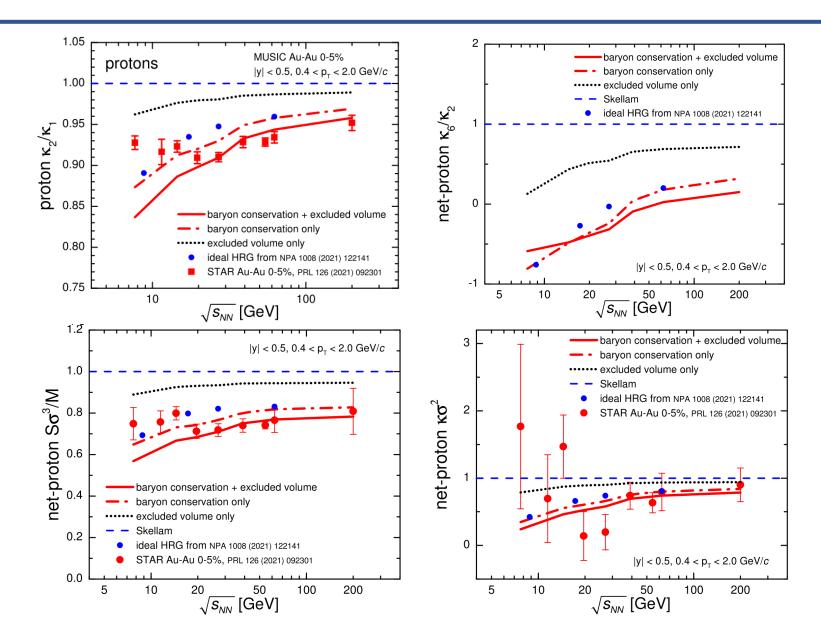
Non-critical cumulants: Analytic vs Monte Carlo





Non-critical cumulants

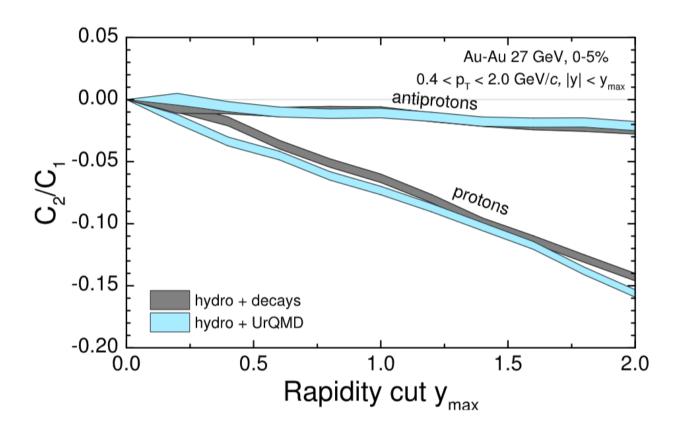




Effect of the hadronic phase



Sample ideal HRG model at particlization with exact conservation of baryon number using Thermal-FIST and run through hadronic afterburner UrQMD



Dependence on the switching energy density



