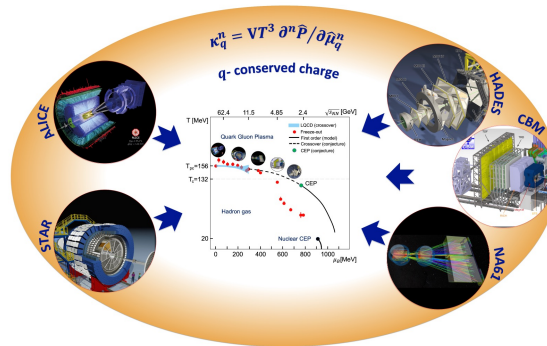


Experimental results vs theory predictions

Volodymyr Vovchenko (University of Houston)



EMMI RRTF “Fluctuations and correlations of conserved charges”

GSI, Darmstadt, Germany

November 9, 2023



Outline

- Introduction
- Some remarks on subensemble acceptance method (SAM)
 - Charge conservation for any equation of state in thermodynamic limit
 - Quantifying finite volume corrections in ideal gas
- Multiple conserved charges
- Non-critical baseline from hydrodynamics
 - Low-order vs high-order cumulants
 - Ordinary vs factorial cumulants
 - Net protons vs protons vs antiprotons
 - Protons vs baryons

Introduction

Event-by-event fluctuations and statistical mechanics

Cumulant generating function

$$K_N(t) = \ln \langle e^{tN} \rangle = \sum_{n=1}^{\infty} \kappa_n \frac{t^n}{n!}$$

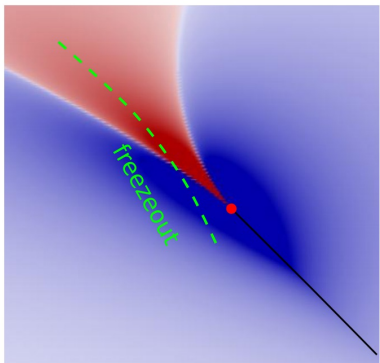
$$\kappa_n \propto \frac{\partial^n (\ln Z^{\text{gce}})}{\partial \mu^n}$$

Grand partition function

$$\ln Z^{\text{gce}}(T, V, \mu) = \ln \left[\sum_N e^{\mu N/T} Z^{\text{ce}}(T, V, N) \right]$$

Cumulants measure chemical potential derivatives of the (QCD) equation of state

- **(QCD) critical point** – large correlation length, critical fluctuations of baryon number



M. Stephanov, PRL '09, '11
Energy scans at RHIC (STAR)
and CERN-SPS (NA61/SHINE)

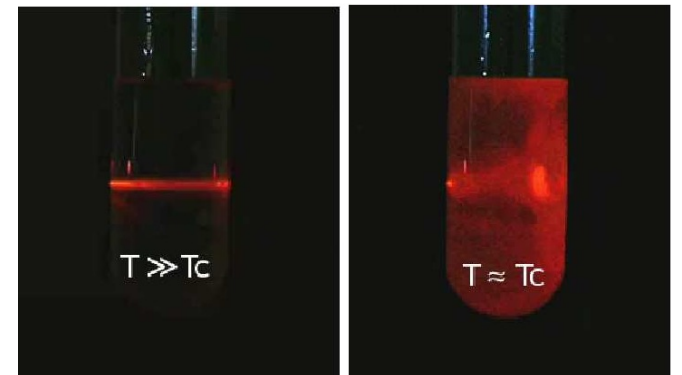
$$\kappa_2 \sim \xi^2, \quad \kappa_3 \sim \xi^{4.5}, \quad \kappa_4 \sim \xi^7$$

$$\xi \rightarrow \infty$$

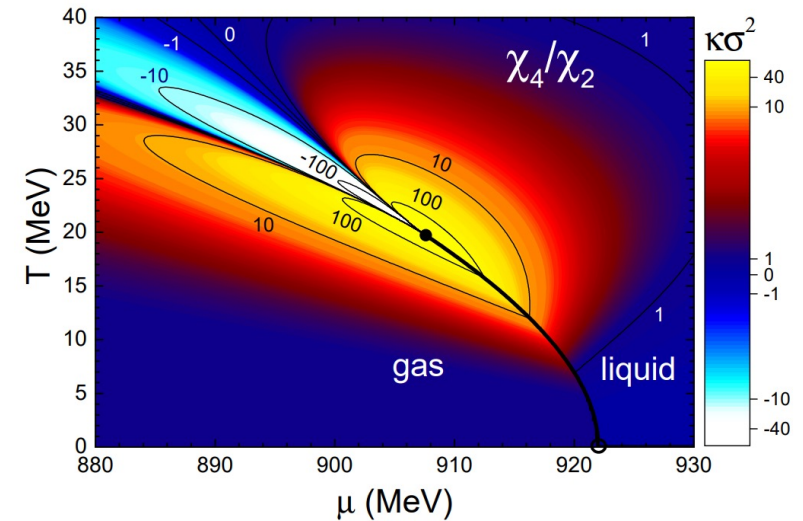
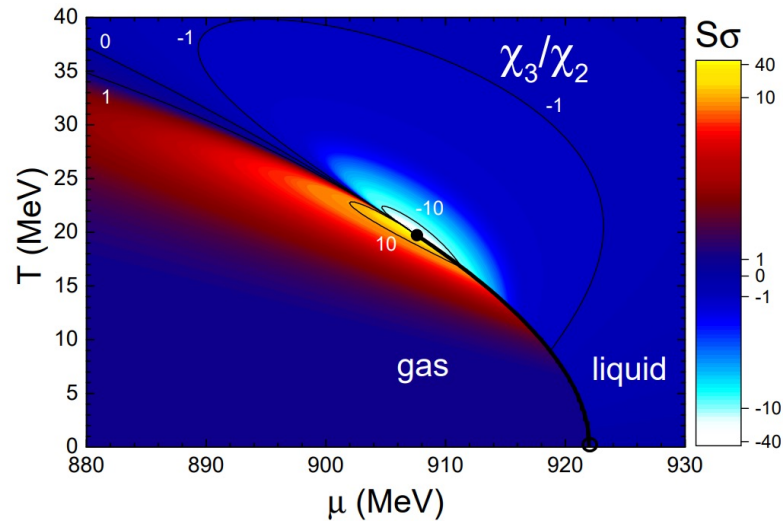
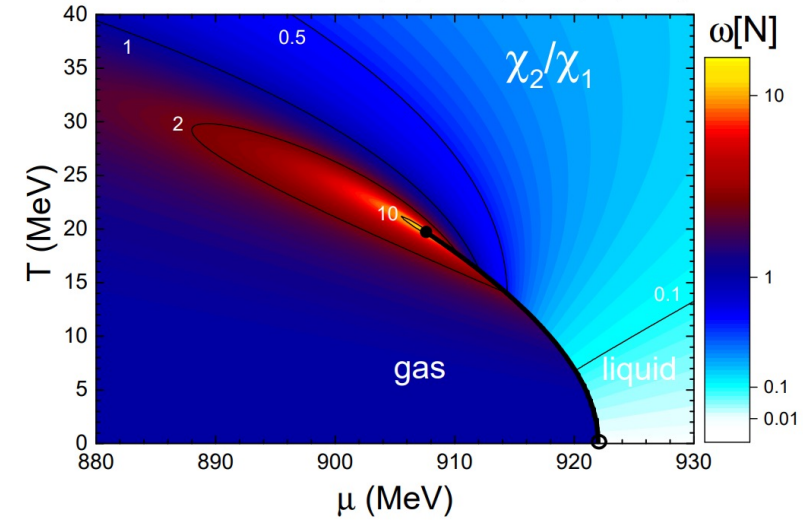
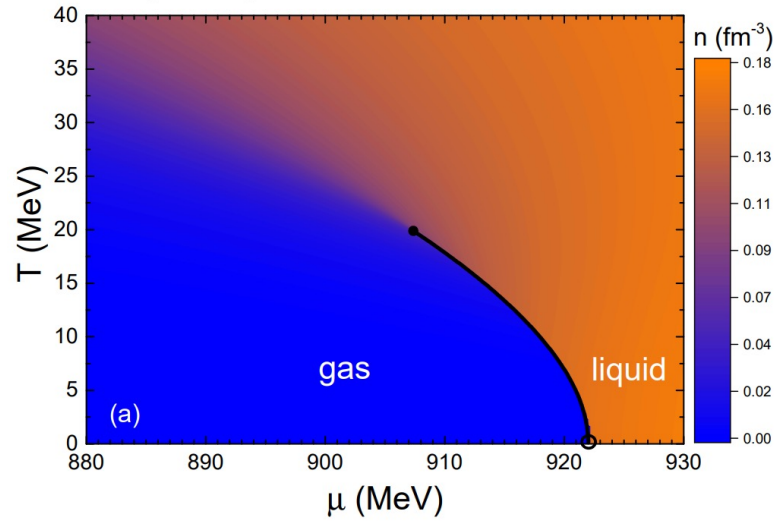
Holds both for ordinary and *factorial* cumulants

Looking for enhanced fluctuations
and non-monotonocities

Critical opalescence



Example: Nuclear liquid-gas transition



VV, Anchishkin, Gorenstein, Poberezhnyuk, PRC 92, 054901 (2015)

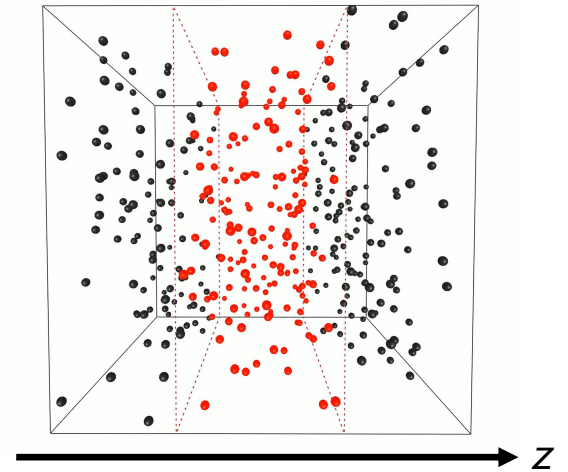
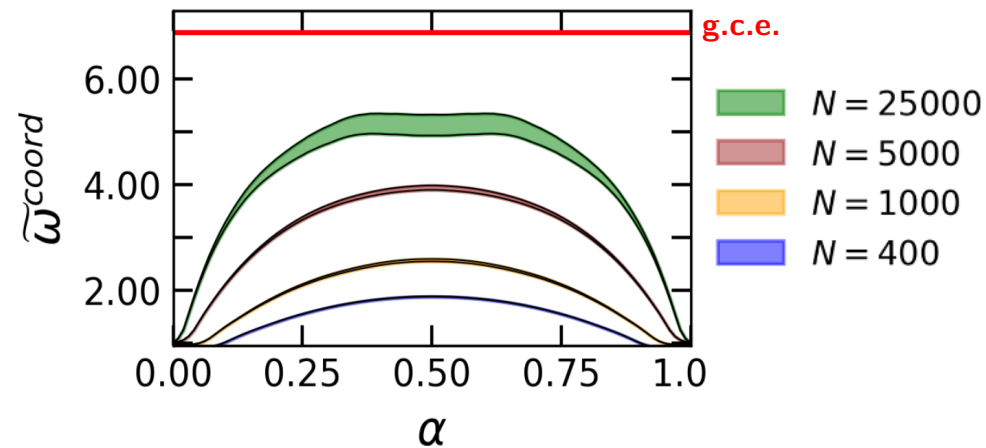
Example: Critical fluctuations in a microscopic simulation

V. Kuznietsov et al., Phys. Rev. C 105, 044903 (2022)

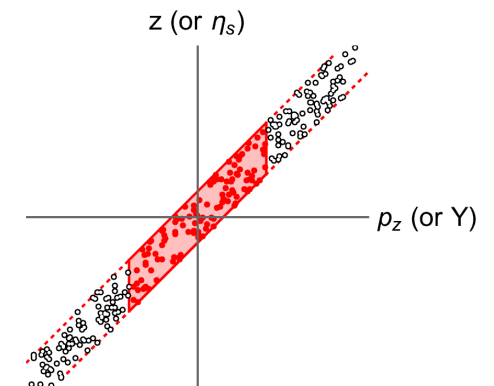
Classical molecular dynamics simulations of the **Lennard-Jones fluid** near Z(2) critical point ($T \approx 1.06T_c$, $n \approx n_c$) of the liquid-gas transition

Scaled variance in coordinate space acceptance $|z| < z^{max}$

$$\tilde{\omega}^{coord} = \frac{1}{1-\alpha} \frac{\langle N^2 \rangle - \langle N \rangle^2}{\langle N \rangle}$$



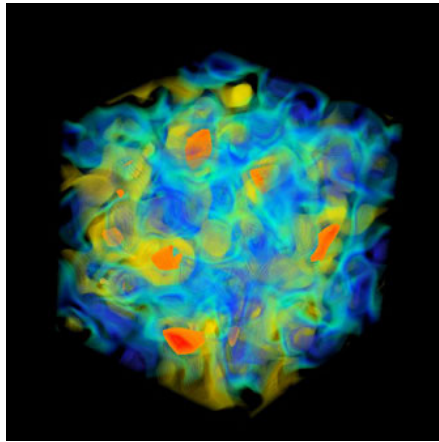
Heavy-ion collisions:
flow correlates p_z and z cuts



- Large fluctuations survive despite strong finite-size effects
- Need coordinate space cuts (collective flow helps)
- Here no finite-time effects

Theory vs experiment: Challenges for fluctuations

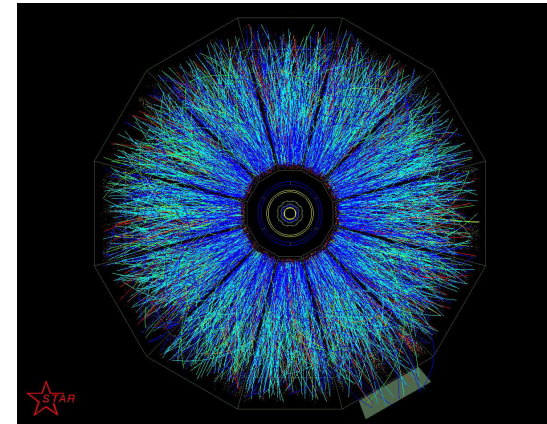
Theory (lattice QCD, fRG/DSE,...)



© Lattice QCD@BNL

- Coordinate space
- In contact with the heat bath
- Conserved charges
- Uniform
- Fixed volume

Experiment



STAR event display

- Momentum space
- Expanding in vacuum
- Non-conserved particle numbers
- Inhomogenous
- Fluctuating volume

Some remarks on subensemble acceptance method

Baryon number conservation: Subensemble acceptance method

VV, Savchuk, Poberezhnyuk, Gorenstein, Koch, PLB 811, 135868 (2020)

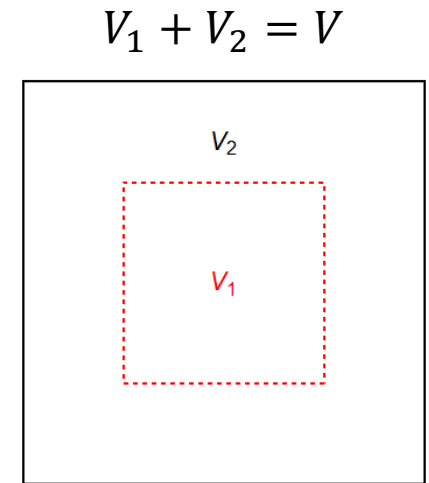
Subensemble acceptance method (SAM) – method to correct *any* EoS (e.g. *lattice QCD*, *fRG*, *ideal HRG*) for **charge conservation**

Partition a thermal system with a globally conserved charge B (*canonical ensemble*) into two subsystems which can exchange the charge

Assume **thermodynamic limit**:

$$V, V_1, V_2 \rightarrow \infty; \quad \frac{V_1}{V} = \alpha = \text{const}; \quad \frac{V_2}{V} = (1 - \alpha) = \text{const}; \quad a = 0: \text{grand-canonical ensemble}$$

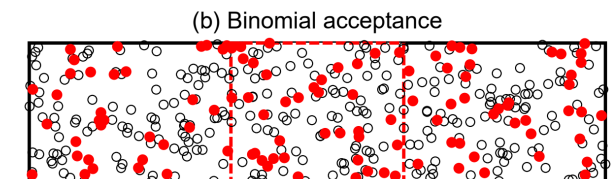
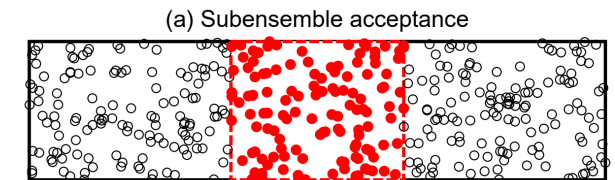
$$V_1, V_2 \gg \xi^3, \quad \xi = \text{correlation length} (= 0 \text{ in ideal gas})$$



The canonical partition function then reads:

$$Z^{\text{ce}}(T, V, B) = \text{Tr} e^{-\beta \hat{H}} \approx \sum_{B_1} Z^{\text{ce}}(T, V_1, B_1) Z^{\text{ce}}(T, V - V_1, B - B_1)$$

$$Z^{\text{ce}}(T, V, B) \stackrel{V \rightarrow \infty}{\simeq} \exp \left[-\frac{V}{T} f(T, \rho_B) \right] \quad \mu_B(T, \rho_B) = \partial f(T, \rho_B) / \partial \rho_B \quad \chi_n^B \equiv \partial^{n-1} (\rho_B / T^3) / \partial (\mu_B / T)^{n-1}$$



SAM: Computing the cumulants

VV, Savchuk, Poberezhnyuk, Gorenstein, Koch, PLB 811, 135868 (2020)

Cumulant generating function for B_1 :

$$G_{B_1}(t) \equiv \ln \langle e^{t B_1} \rangle = \ln \left\{ \sum_{B_1} e^{t B_1} \exp \left[-\frac{\alpha V}{T} f(T, \rho_{B_1}) \right] \exp \left[-\frac{\beta V}{T} f(T, \rho_{B_2}) \right] \right\} + \tilde{c} \quad \beta = 1 - \alpha$$

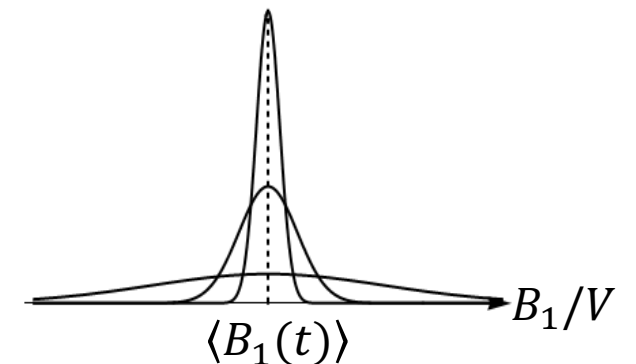
$$\tilde{\kappa}_1[B_1(t)] = \frac{\sum_{B_1} B_1 \tilde{P}(B_1; t)}{\sum_{B_1} \tilde{P}(B_1; t)} \equiv \langle B_1(t) \rangle \quad \text{with} \quad \tilde{P}(B_1; t) = \exp \left\{ t B_1 - V \frac{\alpha f(T, \rho_{B_1}) + \beta f(T, \rho_{B_2})}{T} \right\}.$$

Thermodynamic limit: $\tilde{P}(B_1; t)$ highly peaked at $\langle B_1(t) \rangle$

$\langle B_1(t) \rangle$ is a solution to equation $d\tilde{P}/dB_1 = 0$:

$$t = \hat{\mu}_B[T, \rho_{B_1}(t)] - \hat{\mu}_B[T, \rho_{B_2}(t)]$$

$$\text{where } \hat{\mu}_B \equiv \mu_B/T, \quad \mu_B(T, \rho_B) = \partial f(T, \rho_B)/\partial \rho_B$$



$t = 0$: $\rho_{B_1} = \rho_{B_2} = B/V$, $B_1 = \alpha B$, i.e. charge uniformly distributed between the subsystems

SAM: Full result up to κ_4

VV, Savchuk, Poberezhnyuk, Gorenstein, Koch, PLB 811, 135868 (2020)

$$\kappa_n[B_1] = \left. \frac{\partial^{n-1} \tilde{\kappa}_1[B_1(t)]}{\partial t^{n-1}} \right|_{t=0} \longleftrightarrow \frac{\partial^n}{\partial t^n} t = \hat{\mu}_B[T, \rho_{B_1}(t)] - \hat{\mu}_B[T, \rho_{B_2}(t)]$$

$\kappa_1[B_1] = \alpha VT^3 \chi_1^B$	$\beta = 1 - \alpha$	scaled variance	$\frac{\kappa_2[B_1]}{\kappa_1[B_1]} = (1 - \alpha) \frac{\chi_2^B}{\chi_1^B},$
$\kappa_2[B_1] = \alpha VT^3 \beta \chi_2^B$		skewness	$\frac{\kappa_3[B_1]}{\kappa_2[B_1]} = (1 - 2\alpha) \frac{\chi_3^B}{\chi_2^B},$
$\kappa_3[B_1] = \alpha VT^3 \beta (1 - 2\alpha) \chi_3^B$		kurtosis	$\frac{\kappa_4[B_1]}{\kappa_2[B_1]} = (1 - 3\alpha\beta) \frac{\chi_4^B}{\chi_2^B} - 3\alpha\beta \left(\frac{\chi_3^B}{\chi_2^B} \right)^2.$
$\kappa_4[B_1] = \alpha VT^3 \beta \left[\chi_4^B - 3\alpha\beta \frac{(\chi_3^B)^2 + \chi_2^B \chi_4^B}{\chi_2^B} \right]$			

$$\chi_n^B = \frac{\partial^n (p/T^4)}{\partial (\mu_B/T)^n} \quad \text{-- grand-canonical susceptibilities, e.g. from lattice QCD!}$$

SAM vs ideal gas

Ideal gas of baryons and antibaryons:

$$VT^3 \chi_n^B = N_B + (-1)^n N_{\bar{B}}$$

SAM:

$$\kappa_1 = \alpha B$$

$$\kappa_2 = \alpha(1 - \alpha)N$$

$$\kappa_3 = \alpha(1 - 2\alpha)B$$

$$\kappa_4 = \alpha\beta \left[(1 - 3\alpha\beta)N - 3\alpha\beta \frac{B^2}{N} \right]$$

$$B = \langle N_B \rangle - \langle N_{\bar{B}} \rangle, \quad N = \langle N_B \rangle + \langle N_{\bar{B}} \rangle$$

$$\text{Ideal gas: } g(t) = \ln \left[\left(\frac{q_+}{q_-} \right)^{B/2} \frac{I_B(2z\sqrt{q_+q_-})}{I_B(2z)} \right],$$

where $q_+ = 1 - p_B + p_B e^t$ and $q_- = 1 - p_{\bar{B}} + p_{\bar{B}} e^{-t}$.

Bzdak, Koch, Skokov, PRC 87, 014901 (2013)

Braun-Munzinger, Rustamov, Stachel, NPA 960, 114 (2017)

$$c_1 = pB,$$

$$c_2 = p(1 - p) \langle N \rangle_C,$$

$$c_3 = c_1(1 - p)(1 - 2p),$$

$$c_4 = c_2 + 3(p^2 q^2 B^2 - c_2^2) + 6pq(2z^2 pq - c_2),$$

$$z = \sqrt{\langle N_B \rangle \langle N_{\bar{B}} \rangle}. \quad \langle N_{B,\bar{B}} \rangle_C = z \frac{I_{B\mp 1}(2z)}{I_B(2z)},$$

$$c \equiv \kappa, \quad p \equiv \alpha, \quad q = 1 - \alpha$$

Why are the expressions different, especially for κ_4 ?

Ideal gas: LO and NLO in system size

$$g(t) = \ln \left[\left(\frac{q_+}{q_-} \right)^{B/2} \frac{I_B(2z\sqrt{q_+q_-})}{I_B(2z)} \right], \quad \text{Large volume limit: } N = \langle N_B \rangle + \langle N_{\bar{B}} \rangle \rightarrow \infty \text{ at fixed } r_B = \frac{B}{N}$$

Use asymptotic expansions: $I_\nu(\nu\tilde{z}) \stackrel{\nu \rightarrow \infty}{\sim} \frac{1}{\sqrt{2\pi\nu}} \frac{e^{\nu\eta}}{(1 + \tilde{z}^2)^{1/4}} \{1 + O(\nu^{-1})\},$ $I_0(z) \stackrel{z \rightarrow \infty}{\sim} \frac{e^z}{\sqrt{2\pi z}}$
for $B = 0$

$$g(t) \sim \ln \left\{ \underbrace{\exp[f(t)N]}_{\text{LO}} \underbrace{a_1(t)}_{\text{NLO}} \right\}$$

$$f(t) = \frac{r_B}{2} \ln \frac{q_+}{q_-} + \sqrt{r_B^2 + (1 - r_B^2)q_+q_-} - 1 + |r_B| \ln \left[\frac{(1 + |r_B|)\sqrt{q_+q_-}}{|r_B| + \sqrt{r_B^2 + (1 - r_B^2)q_+q_-}} \right],$$

$$a_1(t) = \frac{1}{[r_B^2 + (1 - r_B^2)q_+q_-]^{1/4}}$$

Ideal gas: LO and NLO in system size

Ideal gas:

LO

$$\kappa_1^{\text{LO}} = N \alpha r_B,$$

$$\kappa_2^{\text{LO}} = N (1 - \alpha) \alpha$$

$$\kappa_3^{\text{LO}} = N \alpha (1 - \alpha) (1 - 2\alpha) r_B$$

$$\kappa_4^{\text{LO}} = N (1 - \alpha) \alpha [1 - 3\alpha (1 - \alpha) (1 + r_B^2)]$$

NLO

$$\kappa_1^{\text{NLO}} = 0,$$

$$\kappa_2^{\text{NLO}} = -\frac{1}{2} (1 - \alpha) \alpha (1 - r_B^2),$$

$$\kappa_3^{\text{NLO}} = 0,$$

$$\kappa_4^{\text{NLO}} = -\frac{1}{2} (1 - \alpha) \alpha (1 - r_B^2) [1 - 6\alpha (1 - \alpha) (1 - r_B^2)].$$

SAM:

$$\kappa_1^{\text{SAM}} = \alpha N r_B,$$

$$\kappa_2^{\text{SAM}} = \alpha (1 - \alpha) N,$$

$$\kappa_3^{\text{SAM}} = \alpha (1 - \alpha) (1 - 2\alpha) N r_B,$$

$$\kappa_4^{\text{SAM}} = \alpha (1 - \alpha) N [1 - 3\alpha (1 - \alpha) (1 + r_B^2)].$$

SAM at NLO: [Barej, Bzdak, PRC 107, 034914 \(2023\)](#)

$$\kappa_1^{\text{SAM,NLO}} = 0,$$

$$\kappa_2^{\text{SAM,NLO}} = -\frac{1}{2} (1 - \alpha) \alpha (1 - r_B^2),$$

$$\kappa_3^{\text{SAM,NLO}} = 0,$$

$$\kappa_4^{\text{SAM,NLO}} = -\frac{1}{2} (1 - \alpha) \alpha (1 - r_B^2) [1 - 6\alpha (1 - \alpha) (1 - r_B^2)].$$

Use NLO to estimate relative errors

$$\frac{\kappa_2^{LO}}{\kappa_2} = 1 + \frac{1 - r_B^2}{N} + O(N^{-2})$$

Total number of baryons is at least 400, therefore $\delta\kappa_2 \sim 0.25\%$ or less

For comparison, LO error due to Fermi statistics is $\delta\kappa_{2,FD}^{\text{gce}} \propto \frac{K_2(2m/T) \cosh(2\mu_B/T)}{K_2(m/T) \cosh(\mu_B/T)} \sim 0.2\%$ at $\mu_B = 0$

$$\begin{aligned} \frac{\kappa_4}{\kappa_2} &= [1 - 3\alpha(1 - \alpha)(1 + r_B^2)] \\ &+ \frac{3\alpha(1 - \alpha)(1 - 4r_B^2 + 3r_B^4)}{2N} + O(N^{-2}), \end{aligned}$$

$$\begin{aligned} \frac{\kappa_6}{\kappa_2} &= 1 - 15\alpha(1 - \alpha)[1 + r_B^2 - \alpha(1 - \alpha)](3 + 6r_B^2 - r_B^4) \\ &+ \frac{15\alpha(1 - \alpha)(1 - r_B^2)[1 - 3r_B^2 - \alpha(1 - \alpha)(5 - 22r_B^2 + 9r_B^4)]}{2N} + O(N^{-2}) \end{aligned}$$

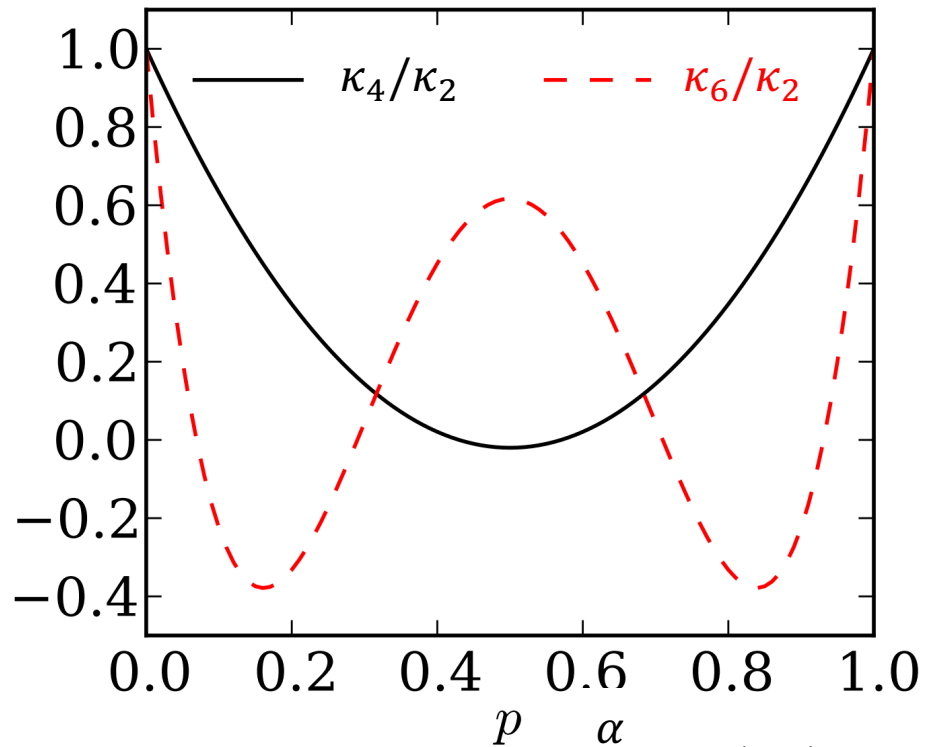
When LO corrections might be sizable?

- Strangeness conservation at low energies
- pp collisions

Comparison

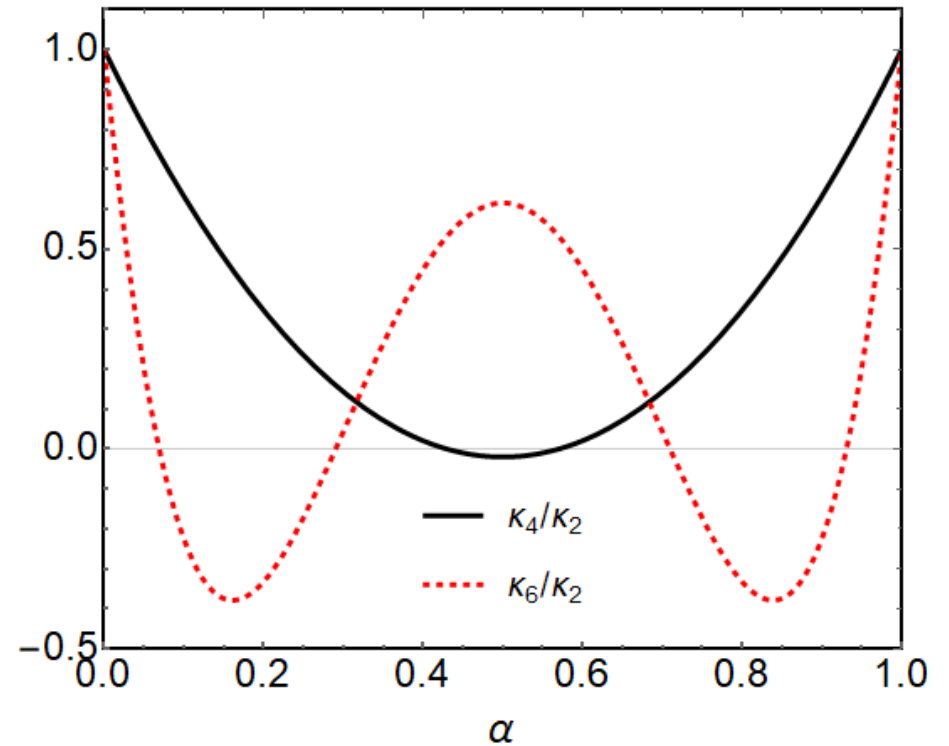
Ideal gas of baryons and antibaryons: $\chi_{2n}^B \propto \langle N_B \rangle + \langle N_{\bar{B}} \rangle$, $\chi_{2n-1}^B \propto \langle N_B \rangle - \langle N_{\bar{B}} \rangle$

Ideal gas [Bzdak, Koch, Skokov, PRC 87, 014901 (2013)]



$\langle N_B \rangle = 400$, $\langle N_{\bar{B}} \rangle = 100$

SAM/LO



SAM-2.0: apply the correction for *arbitrary* distributions inside and outside the acceptance that are peaked at the mean

- Inhomogeneous systems (e.g. RHIC and below)
- Momentum space and unequal acceptances
- Map “grand-canonical” cumulants inside and outside the acceptance to the “canonical” cumulants inside the acceptance

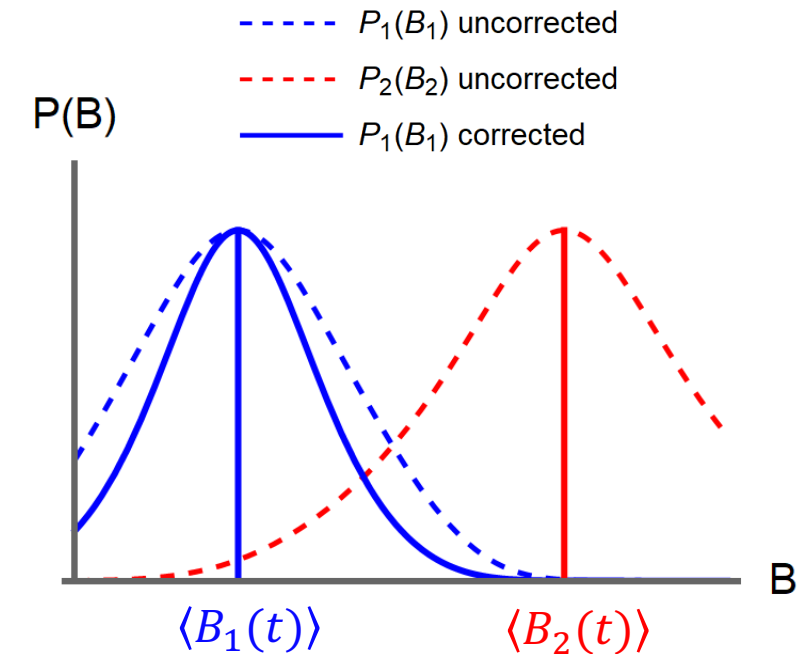
$$\kappa_{in}^{ce} = \text{SAM}[\kappa_{in}^{gce}, \kappa_{out}^{gce}]$$

$$\kappa_1^{\text{SAM2.0}} = \kappa_{1,in}^{gce},$$

$$\kappa_2^{\text{SAM2.0}} = \frac{\kappa_{2,in}^{gce} \kappa_{2,out}^{gce}}{\kappa_{2,in}^{gce} + \kappa_{2,out}^{gce}},$$

$$\kappa_3^{\text{SAM2.0}} = \frac{\kappa_{3,in}^{gce} (\kappa_{2,out}^{gce})^3 - \kappa_{3,out}^{gce} (\kappa_{2,in}^{gce})^3}{(\kappa_{2,in}^{gce} + \kappa_{2,out}^{gce})^3},$$

$$\kappa_4^{\text{SAM2.0}} = \frac{1}{(\kappa_{2,in}^{gce} + \kappa_{2,out}^{gce})^5} \times \left\{ \kappa_{4,in}^{gce} (\kappa_{2,out}^{gce})^5 + \kappa_{4,out}^{gce} (\kappa_{2,in}^{gce})^5 + (\kappa_{2,out}^{gce})^4 [\kappa_{2,in}^{gce} \kappa_{4,in}^{gce} - 3(\kappa_{3,in}^{gce})^2] \right. \\ \left. + (\kappa_{2,in}^{gce})^4 [\kappa_{2,out}^{gce} \kappa_{4,out}^{gce} - 3(\kappa_{3,out}^{gce})^2] - 6(\kappa_{2,in}^{gce})^2 (\kappa_{2,out}^{gce})^2 \kappa_{3,in}^{gce} \kappa_{3,out}^{gce} \right\}.$$



$$P_1^{ce}(B_1) \propto \sum_{B_1, B_2} P_1^{gce}(B_1) P_2^{gce}(B_2) \times \delta_{B, B_1+B_2}$$

Grand-canonical cumulants read:

$$\kappa_{n,in}^{\text{gce}} = \alpha_B \langle N_B \rangle_{GC} + (-1)^n \alpha_{\bar{B}} \langle N_{\bar{B}} \rangle_{GC}$$

$$\kappa_{n,out}^{\text{gce}} = (1 - \alpha_B) \langle N_B \rangle_{GC} + (-1)^n (1 - \alpha_{\bar{B}}) \langle N_{\bar{B}} \rangle_{GC}$$

$$\kappa_1^{\text{SAM2.0}} = N \frac{1}{2} [\alpha_B (r_B + 1) + \alpha_{\bar{B}} (r_B - 1)],$$

$$\kappa_2^{\text{SAM2.0}} = -N \frac{1}{4} (\alpha_B^2 (r_B + 1)^2 - 2\alpha_B (\alpha_{\bar{B}} (r_B^2 - 1) + r_B + 1) + \alpha_{\bar{B}} (r_B - 1) (\alpha_{\bar{B}} (r_B - 1) + 2))$$

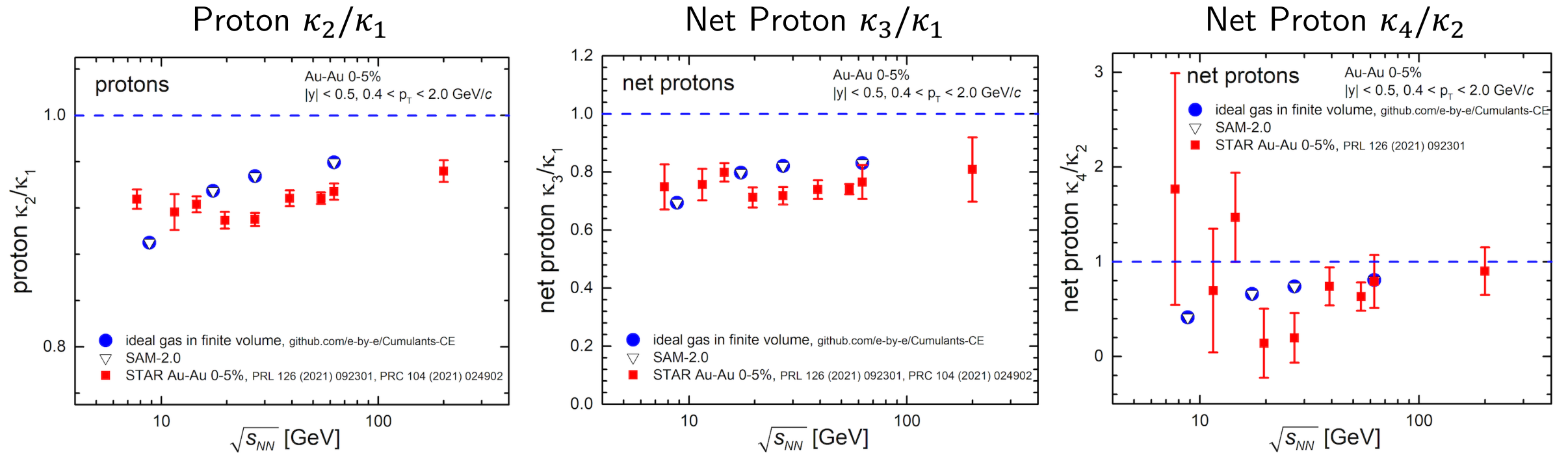
$$\begin{aligned} \kappa_3^{\text{SAM2.0}} = & -N \frac{1}{8} (\alpha_B^3 (r_B - 3)(r_B + 1)^3 - 3\alpha_B^2 (r_B + 1)^2 (\alpha_{\bar{B}} (r_B - 1)^2 - 2) \\ & + \alpha_B (3\alpha_{\bar{B}}^2 (r_B^2 - 1)^2 - 4(r_B + 1)) - \alpha_{\bar{B}} (r_B - 1) (\alpha_{\bar{B}}^2 (r_B + 3)(r_B - 1)^2 + 6\alpha_{\bar{B}} (r_B - 1) + 4)) \end{aligned}$$

$$\begin{aligned} \kappa_4^{\text{SAM2.0}} = & -N \frac{1}{16} (3\alpha_B^4 (r_B^2 - 4r_B + 5) (r_B + 1)^4 - 12\alpha_B^3 (r_B + 1)^3 (\alpha_{\bar{B}} (r_B - 1)^3 - r_B + 3) \\ & + 2\alpha_B^2 (r_B + 1)^2 (3\alpha_{\bar{B}}^2 (3r_B^2 - 1) (r_B - 1)^2 - 6\alpha_{\bar{B}} (r_B - 1)^2 + 14) \\ & - 4\alpha_B (3\alpha_{\bar{B}}^3 (r_B^2 - 1)^3 + 3\alpha_{\bar{B}}^2 (r_B^2 - 1)^2 + 2\alpha_{\bar{B}} (r_B^2 - 1) + 2(r_B + 1)) \\ & + \alpha_{\bar{B}} (r_B - 1) (3\alpha_{\bar{B}}^3 (r_B^2 + 4r_B + 5) (r_B - 1)^3 + 12\alpha_{\bar{B}}^2 (r_B + 3)(r_B - 1)^2 + 28\alpha_{\bar{B}} (r_B - 1) + 8)), \end{aligned}$$

in agreement with LO results from finite volume calculation

Comparison at RHIC-BES

Using α_B and $\alpha_{\bar{B}}$ at different energies from Braun-Munzinger et al. NPA 1008, 122141 (2021)



Ideal gas: using <https://github.com/e-by-e/Cumulants-CE>

No visible difference to SAM-2.0

Multiple conserved charges

SAM for multiple conserved charges (B,Q,S)

VV, Poberezhnyuk, Koch, JHEP 10, 089 (2020)

Key findings:

- Cumulants up to 3rd order factorize into product of binomial and grand-canonical cumulants

$$\kappa_2^{B,Q,S} = (1 - \alpha) VT^3 \chi_2^{B,Q,S} \quad \kappa_3^{B,Q,S} = (1 - \alpha)(1 - 2\alpha) VT^3 \chi_3^{B,Q,S}$$

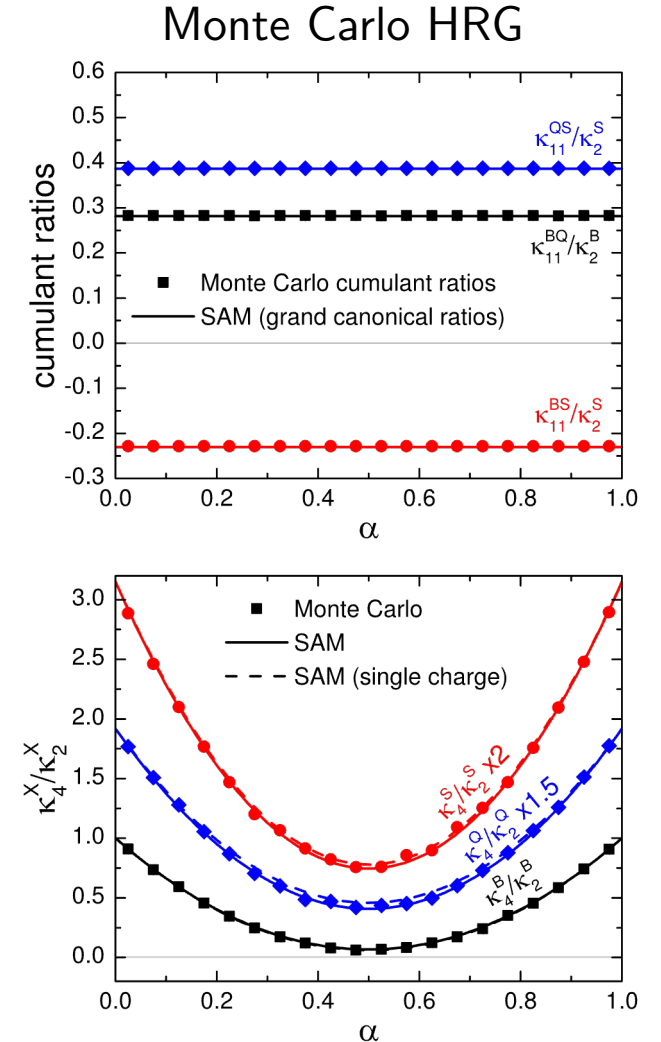
- Ratios of second and third order cumulants of **conserved charges** are NOT sensitive to charge conservation (requires same acceptance)

e.g. $\kappa_2^B / \kappa_2^Q = \chi_2^B / \chi_2^Q$

- Also true for the measurable ratios of covariances involving one non-conserved charge, such as $\kappa_{pQ} / \kappa_{kQ}$

- For order $n > 3$ charge cumulants “mix”. Effect in HRG is tiny

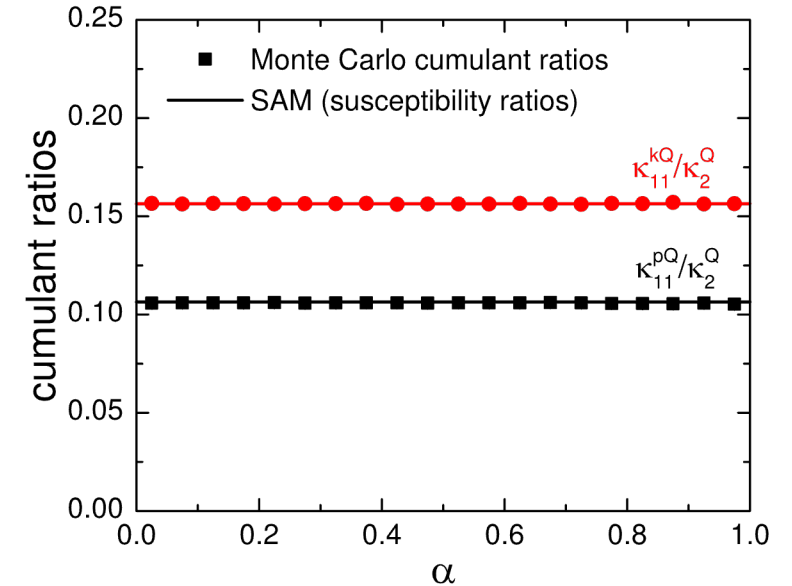
$$\kappa_4^B = \kappa_4^{B,gce} \beta \left[(1 - 3\alpha\beta) \chi_4^B - 3\alpha\beta \frac{(\chi_3^B)^2 \chi_2^Q - 2\chi_{21}^{BQ} \chi_{11}^{BQ} \chi_3^B + (\chi_{21}^{BQ})^2 \chi_2^B}{\chi_2^B \chi_2^Q - (\chi_{11}^{BQ})^2} \right]$$



SAM and non-conserved quantities

$$\kappa_{XY} = (1 - \alpha) \kappa_{XY}^{gce} + \alpha \kappa_{XY}^{ce}$$

- Mixed cumulants involving one conserved charge e.g. pQ have $\kappa_{pQ}^{ce} = 0$ thus they scale like second order charge cumulants
 - p and Q , again, must have the same a
- Cancellation does NOT occur for two non-conserved quantities, such as κ_{pK}



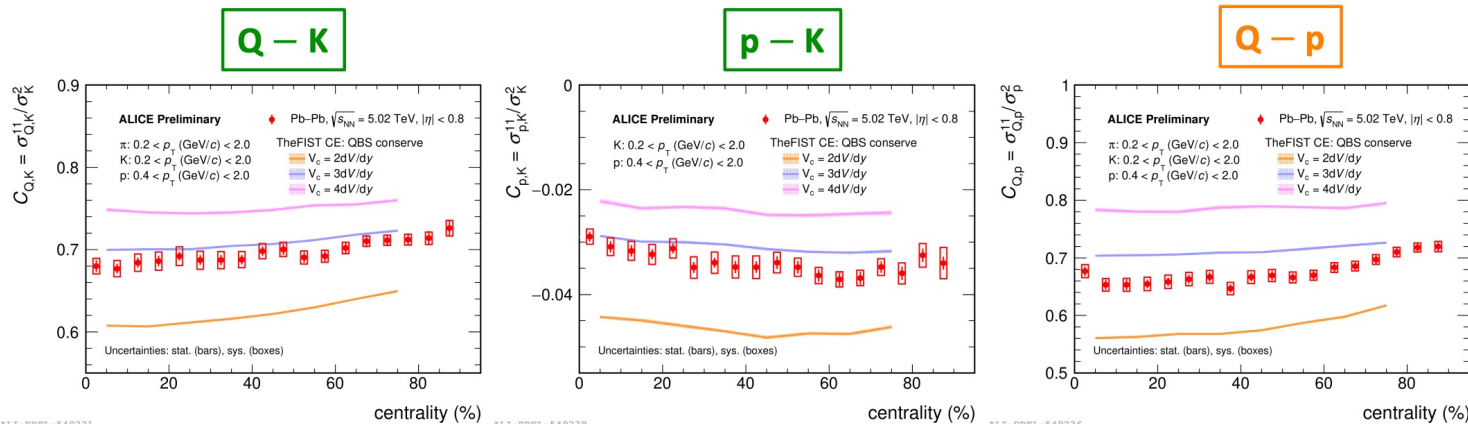
Replace $\sigma_{QK}^{11}/\sigma_K^2$ and $\sigma_{Qp}^{11}/\sigma_p^2$ by

$$\sigma_{QK}^{11}/\sigma_Q^2 \text{ and } \sigma_{Qp}^{11}/\sigma_Q^2$$



Eliminate V_c dependence?

ALICE, M. Arslandok (Monday)



Net-proton and net- Λ fluctuations

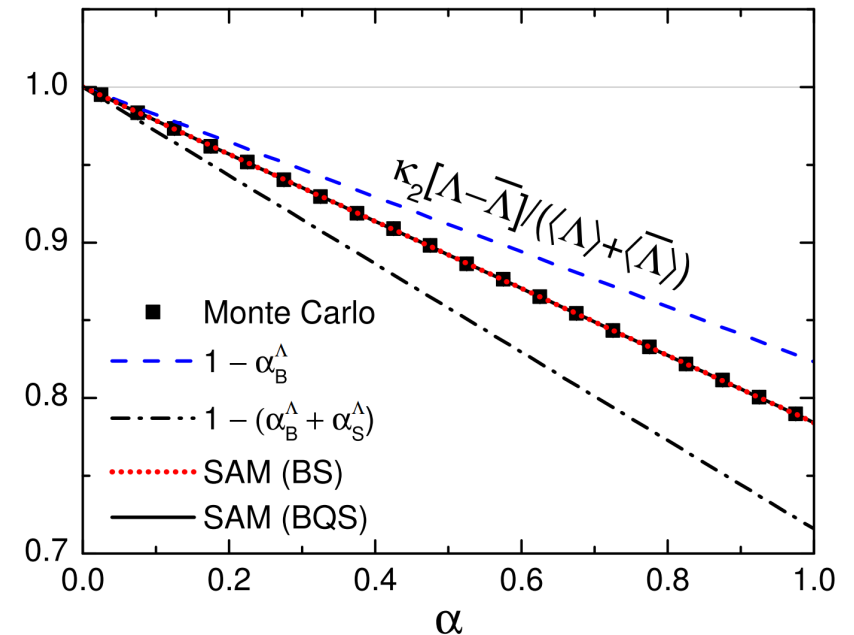
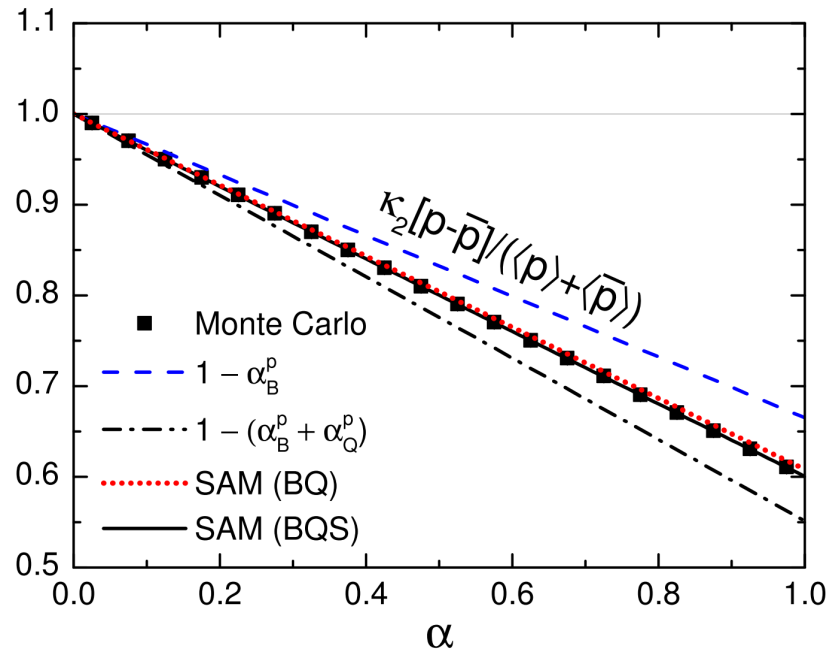
$$\kappa_{pp} = (1 - \alpha) \kappa_{pp}^{\text{gce}} + \alpha \kappa_{pp}^{\text{ce}}$$

- Allows for corrections due to electric charge (protons) or strangeness (Λ) conservation in addition to baryon number conservation.

$$\alpha = \frac{\langle p \rangle_{\text{acc}}}{\langle p \rangle_{4\pi}}$$

$$\alpha_B^p = \frac{\langle p \rangle_{\text{acc}}}{\langle B \rangle_{4\pi}}$$

$$\alpha_Q^p = \frac{\langle p \rangle_{\text{acc}}}{\langle Q \rangle_{4\pi}}$$

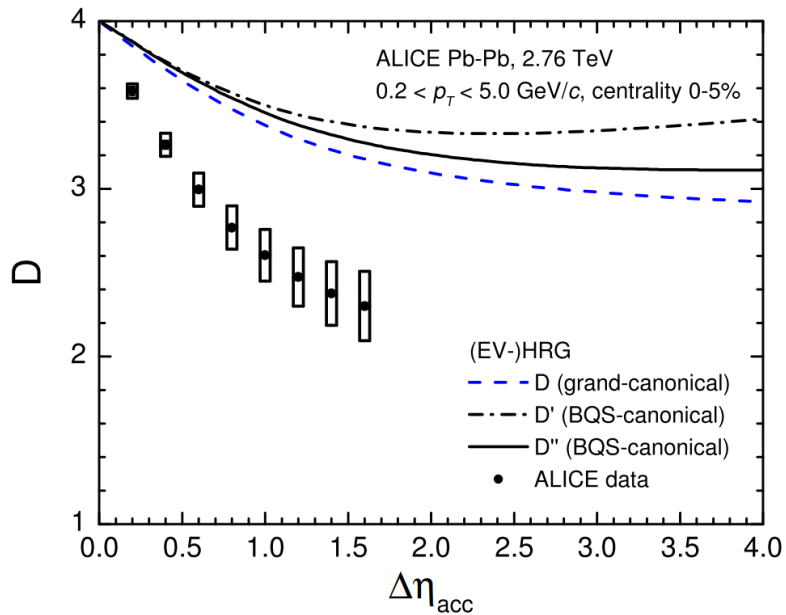


Truth lies in between the “naïve” corrections
Likely bigger effect for higher orders

Quantitative calculation with blast-wave model

- Large effect from resonance decays for pions and kaons + exact conservation of electric charge/strangeness

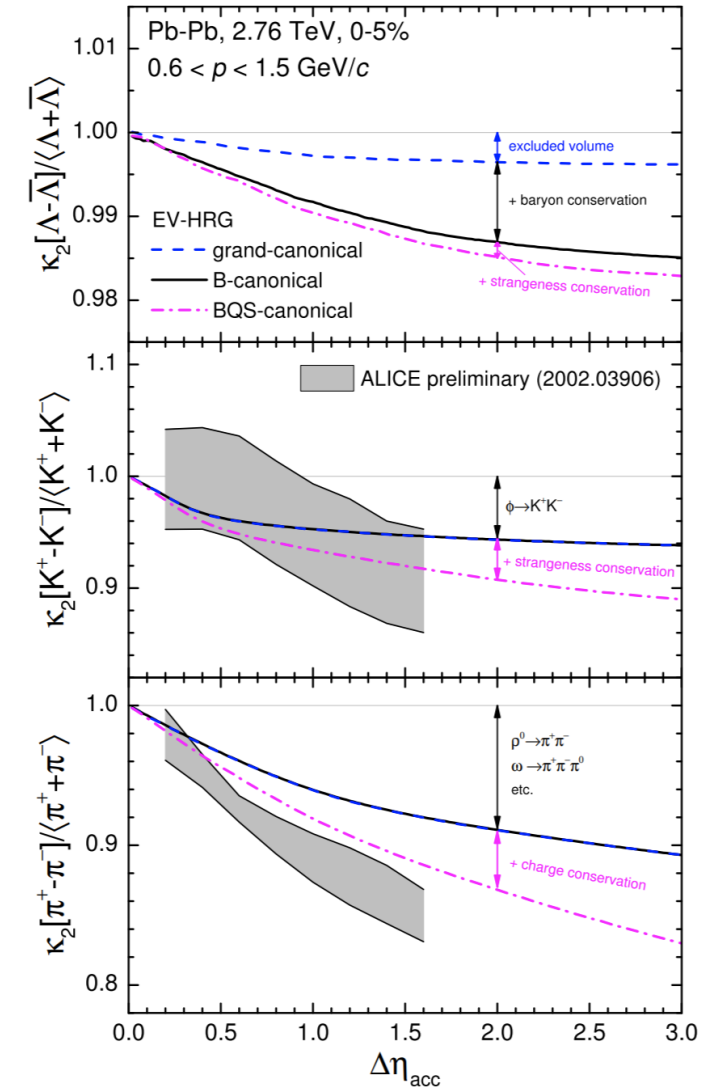
- D-measure
$$D = \frac{\langle \delta Q^2 \rangle}{\langle N_{ch} \rangle}$$



HRG: $D \sim 3 - 4$

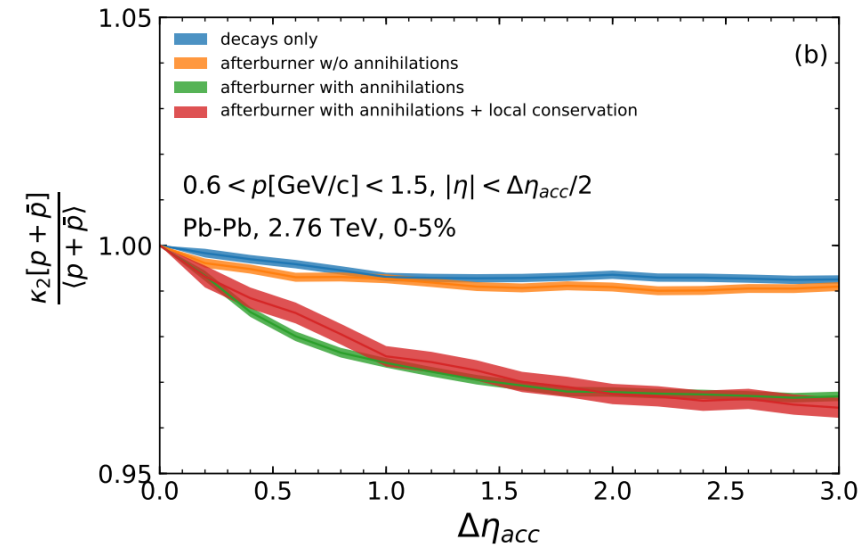
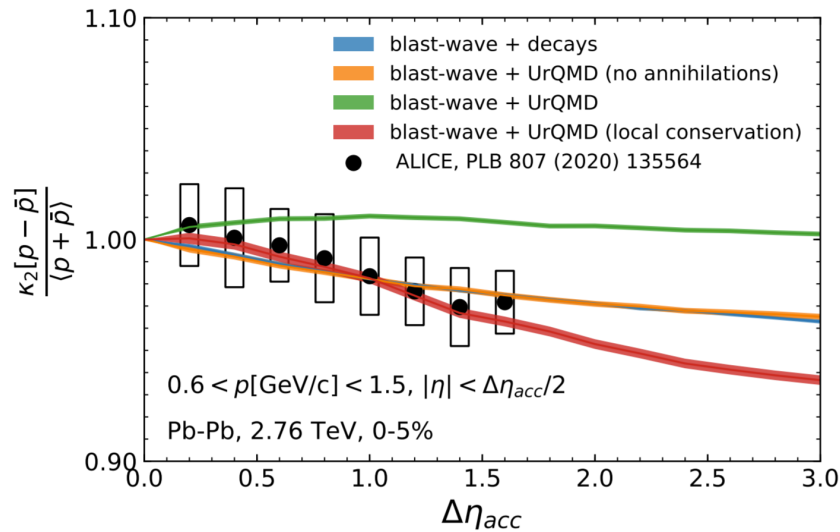
QGP?: $D \sim 1 - 1.5$

Hadronic description with global conservation challenging



Baryon annihilation

- Net protons described within errors and consistent with either
 - **global** baryon conservation without $B\bar{B}$ annihilations
 - or **local** baryon conservation with $B\bar{B}$ annihilations

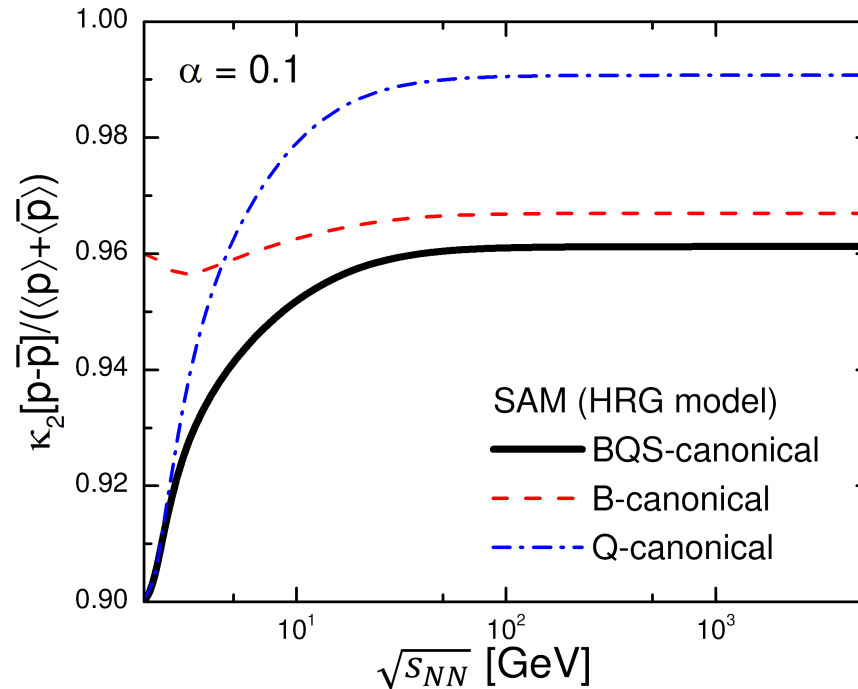


O. Savchuk et al., Phys. Lett. B 827, 136983 (2022)

- $\kappa_2[p - \bar{p}] / \langle p + \bar{p} \rangle$: Interplay of **baryon annihilation**(↗) and **local conservation**(↘)
 - Additional measurement of $\kappa_2[p + \bar{p}]$ can resolve it
- $\kappa_2[p + \bar{p}] / \langle p + \bar{p} \rangle$: Insensitive to baryon conservation at LHC, $cov(p + \bar{p}, B - \bar{B})=0$
 - Good measure for volume fluctuations?

Multiple conserved charges as fct. of collision energy

Schematic calculation in HRG model along the chemical freeze-out line, $\alpha = 0.1$

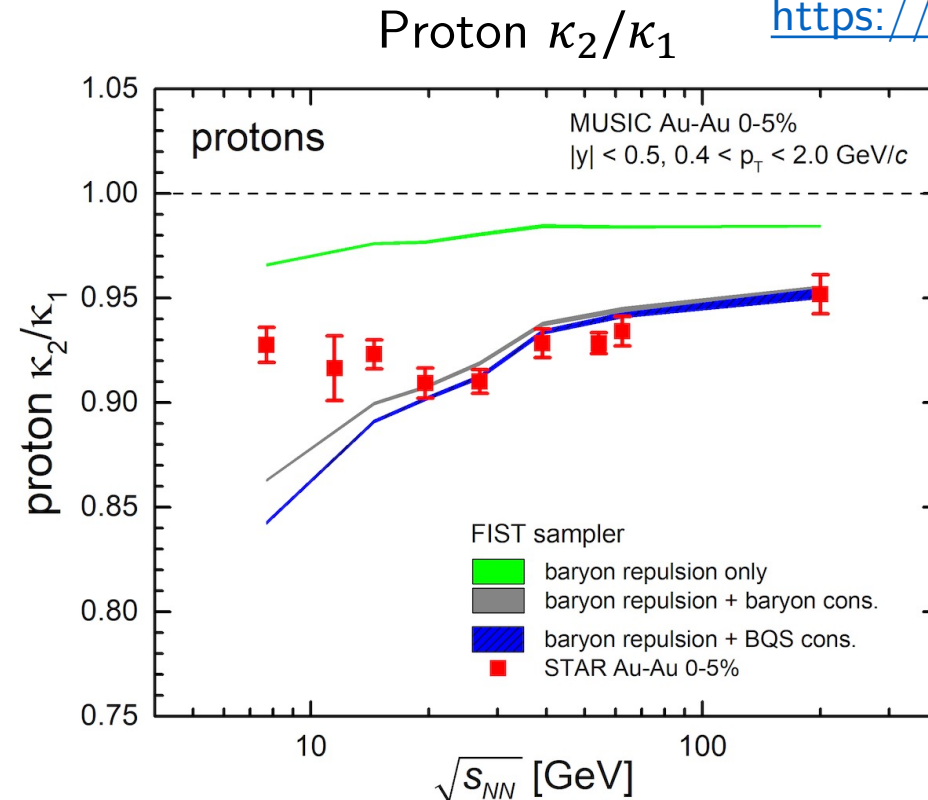


- LHC: Dominant effect from baryon conservation
- Very low energies: net-p \approx net-Q \Rightarrow electric charge conservation dominates
- Simultaneous treatment of B and Q conservation is important

Quantitative calculation for BES energies

Canonical sampling of HRG over MUSIC Cooper-Frye hypersurfaces – FIST sampler

<https://github.com/vlvovch/fist-sampler>



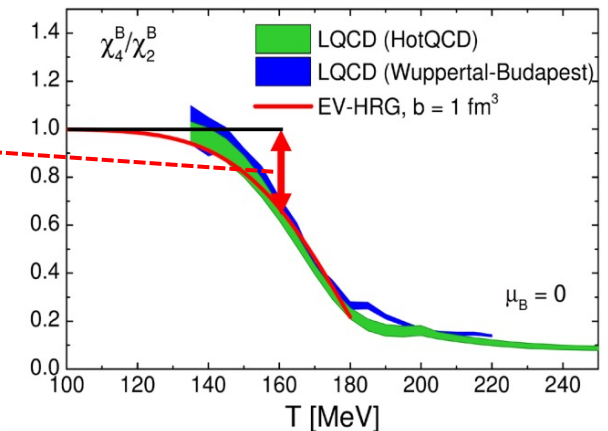
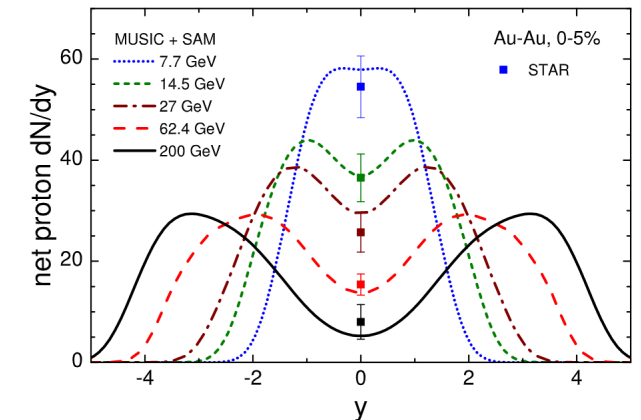
- Effect of charge conservation in addition to baryon conservation becomes visible as energy is decreased
- Becomes dominant effect at 2.4 & 3 GeV?

**Non-critical baseline from
hydrodynamics**

Calculation of non-critical contributions

VV, V. Koch, C. Shen, Phys. Rev. C 105, 014904 (2022)

- (3+1)-D viscous hydrodynamics evolution (MUSIC-3.0)
 - Collision geometry-based 3D initial state [Shen, Alzhrani, PRC 102, 014909 (2020)]
 - Crossover equation of state based on lattice QCD [Monnai, Schenke, Shen, Phys. Rev. C 100, 024907 (2019)]
 - Cooper-Frye particlization at $\epsilon_{sw} = 0.26 \text{ GeV}/\text{fm}^3$
- Non-critical contributions are computed at particlization
 - Cumulants matched to QCD at $\mu_B = 0$ via excluded volume $b = 1 \text{ fm}^3$ [VV, V. Koch, Phys. Rev. C 103, 044903 (2021)]
 - Exact global baryon conservation* (and other charges)
 - SAM-2.0 [VV, Phys. Rev. C 105, 014903 (2022)]
 - or FIST sampler (Monte Carlo) [VV, Phys. Rev. C 106, 064906 (2022)]
<https://github.com/vlvovch/fist-sampler>

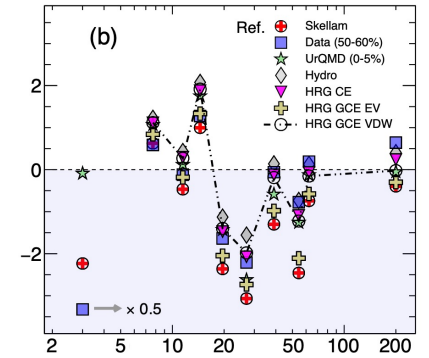
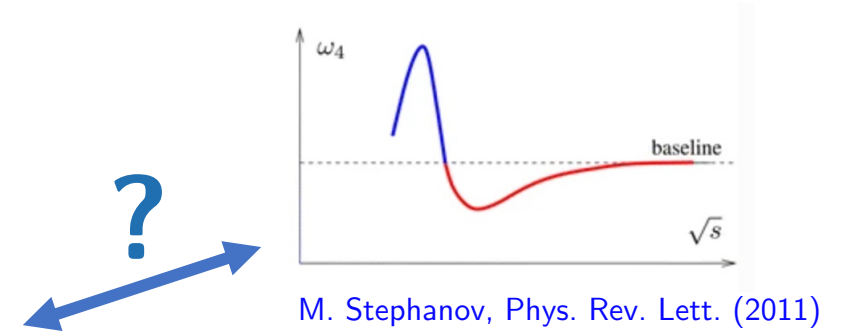
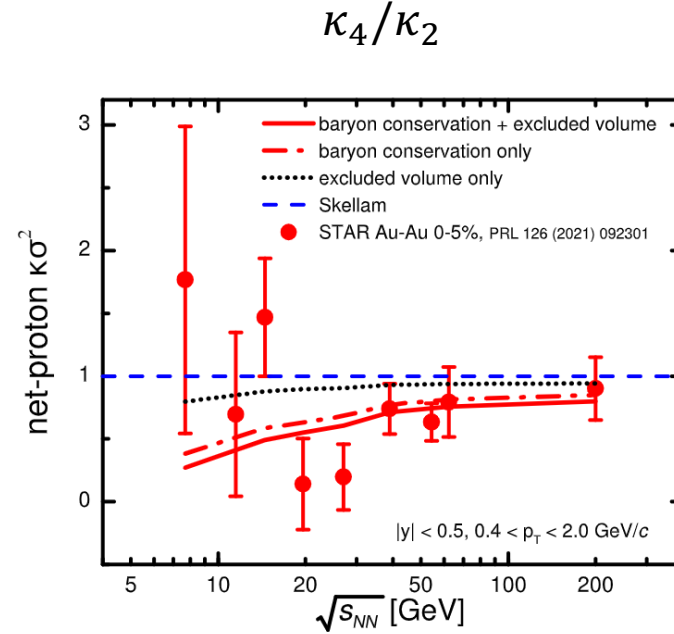
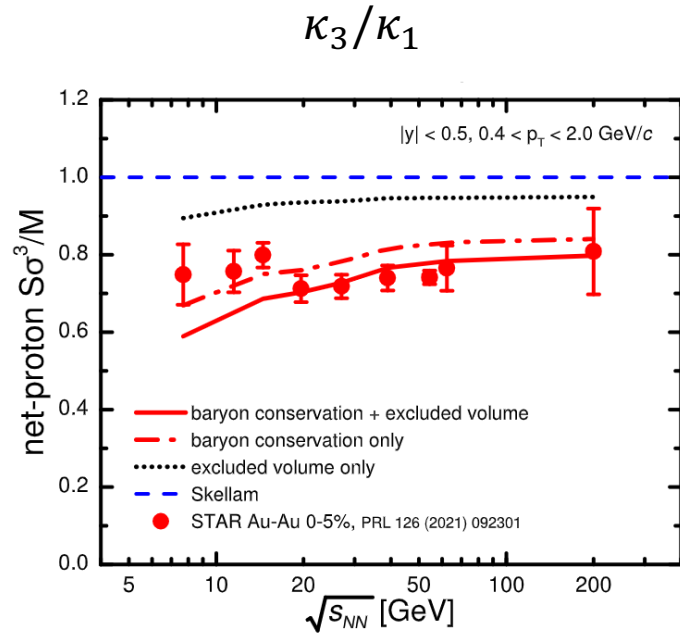


- **Absent:** critical point, local conservation, initial-state/volume fluctuations

*If baryon conservation is the only effect (no other correlations), non-critical baseline can be computed without hydro

Braun-Munzinger, Friman, Redlich, Rustamov, Stachel, NPA 1008, 122141 (2021)

RHIC-BES: Net proton cumulant ratios (MUSIC)



INT-20r-c1 talk by A. Pandav

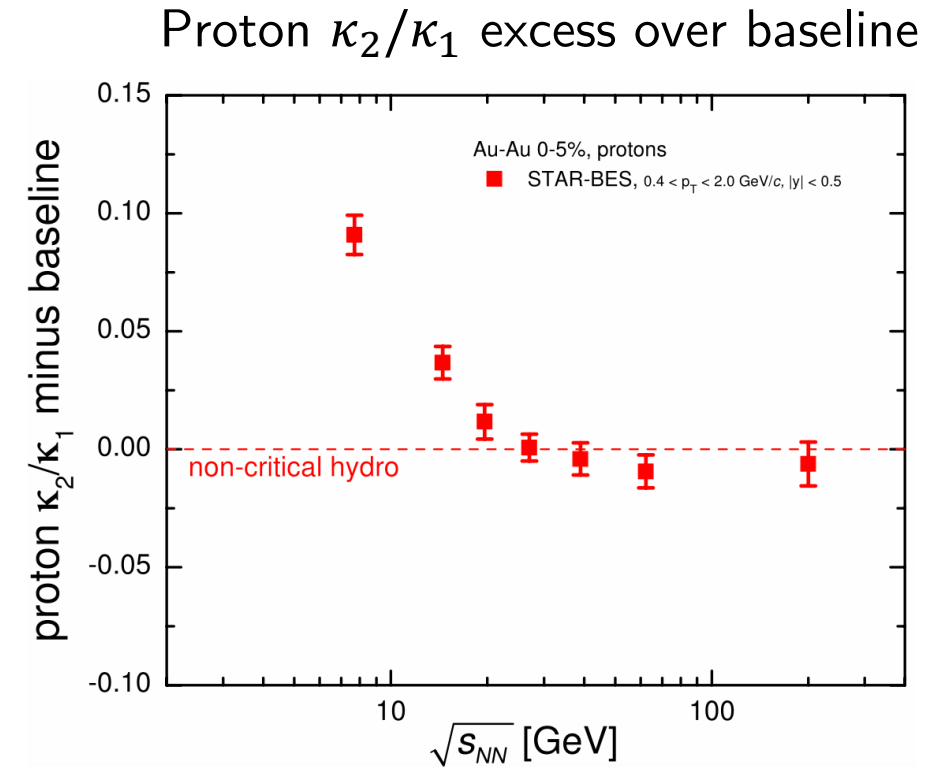
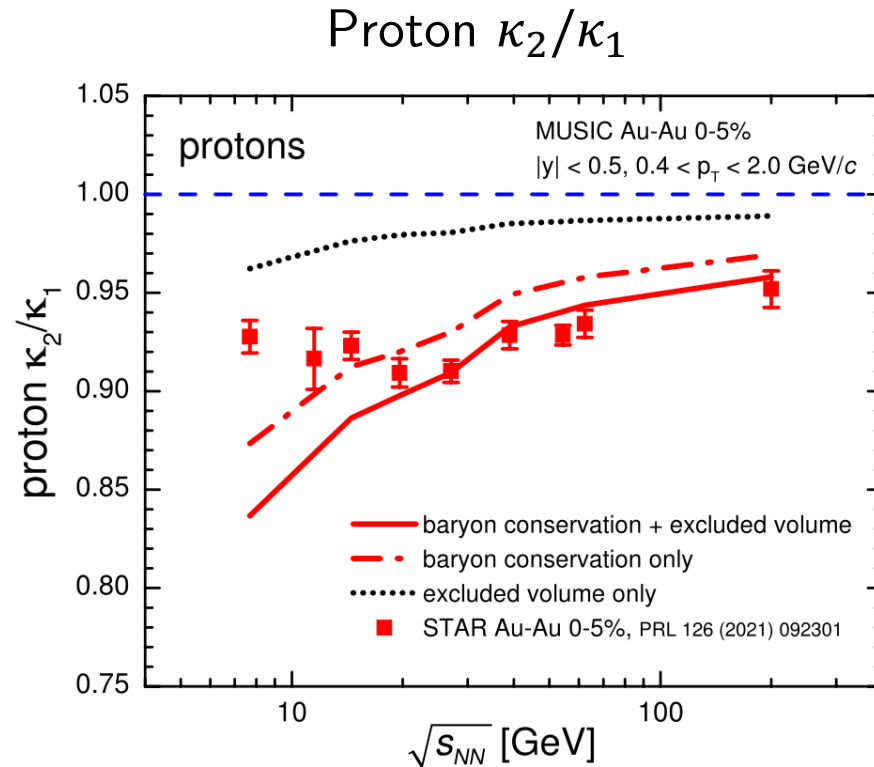
- Data at $\sqrt{s_{NN}} \geq 20$ GeV consistent with non-critical physics (BQS conservation and repulsion)
- Effect from baryon conservation is stronger than repulsion but both are required at $\sqrt{s_{NN}} \geq 20$ GeV
- Reduced errors to come from BES-II

Can we learn more from the more accurate data available for κ_2 and κ_3 ?

Removing the “net” part: Proton variance

Net-proton $\kappa_2/\kappa_1 \sim \frac{\langle p+\bar{p} \rangle}{\langle p-\bar{p} \rangle} \sim \coth\left(\frac{\mu_B}{T}\right)$ in free gas

Proton $\kappa_2/\kappa_1 \sim \frac{\langle p \rangle}{\langle \bar{p} \rangle} = 1$ in free gas



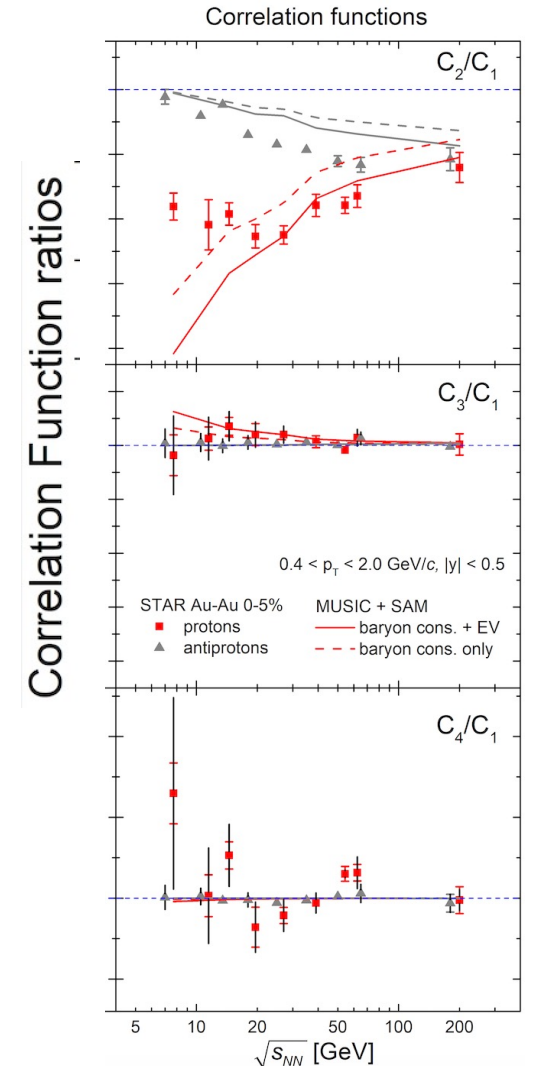
- Data at $\sqrt{s_{NN}} \geq 20$ GeV consistent with non-critical physics (BQS conservation and repulsion)
- Clear excess of proton variance at $\sqrt{s_{NN}} < 20$ GeV – *hint of attractive interactions?*

Correlation Functions (factorial cumulants)

- **Factorial cumulants \hat{C}_n** [Bzdak, Koch, Strodthoff, PRC 95, 054906 (2017)]
 - Remove the Poisson contribution and probe genuine correlations

$$\hat{C}_1 = \kappa_1, \quad \hat{C}_3 = 2\kappa_1 - 3\kappa_2 + \kappa_3,$$

$$\hat{C}_2 = -\kappa_1 + \kappa_2, \quad \hat{C}_4 = -6\kappa_1 + 11\kappa_2 - 6\kappa_3 + \kappa_4.$$
- **Expectation:** High-order ($n > 3$) factorial cumulants
 - have small contributions from non-critical effects (baryon cons. or excluded volume) [Bzdak, Koch, Skokov, EPJC '17; VV et al, PLB '17]
 - are as singular as ordinary cumulants near the critical point [Ling, Stephanov, PRC '16]
- **Observations from STAR data:**
 - \hat{C}_3 & \hat{C}_4 are largely consistent with zero within (large) errors
 - Reanalyze (non-)monotonic energy dependence for \hat{C}_4/\hat{C}_1 instead of κ_4/κ_2 ?
 - Statistically significant effects appear to be driven by two-proton correlations in \hat{C}_2

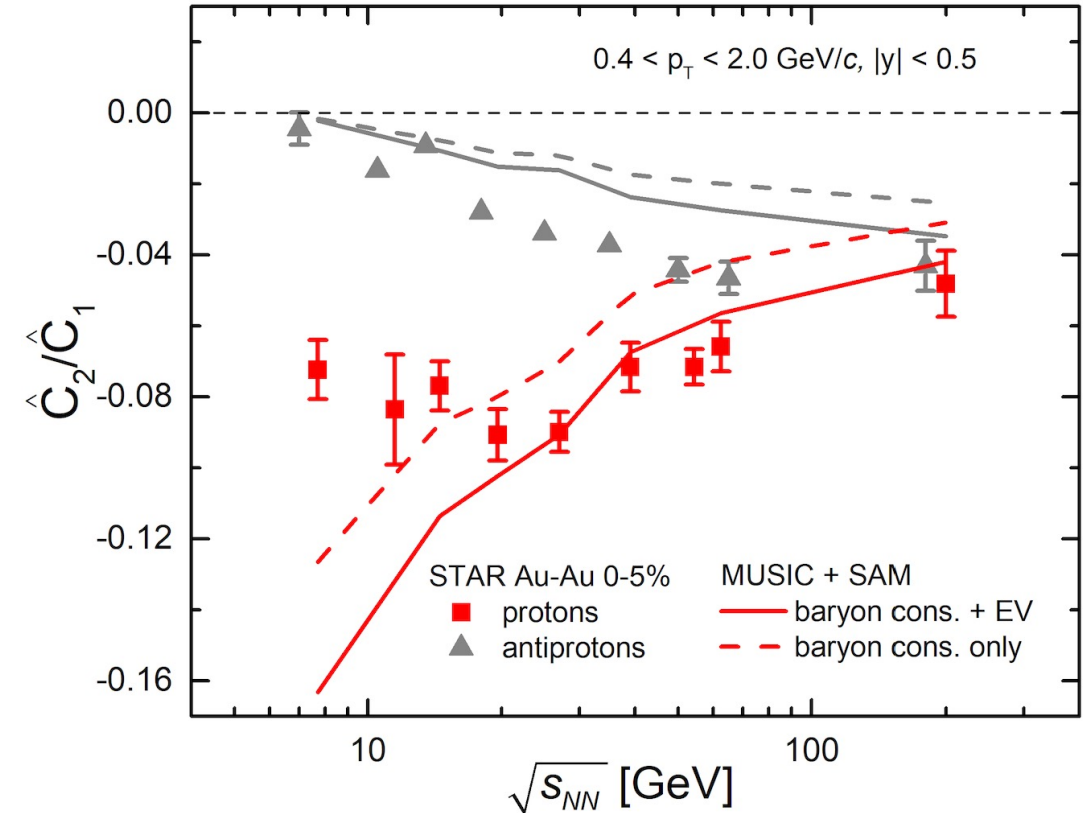


Two-proton correlations

- Protons
 - Consistent with non-critical physics above 20 GeV
 - Enhancement at lower energies
- Antiproton description has issues
 - Correlations in data underestimated by \sim factor 2

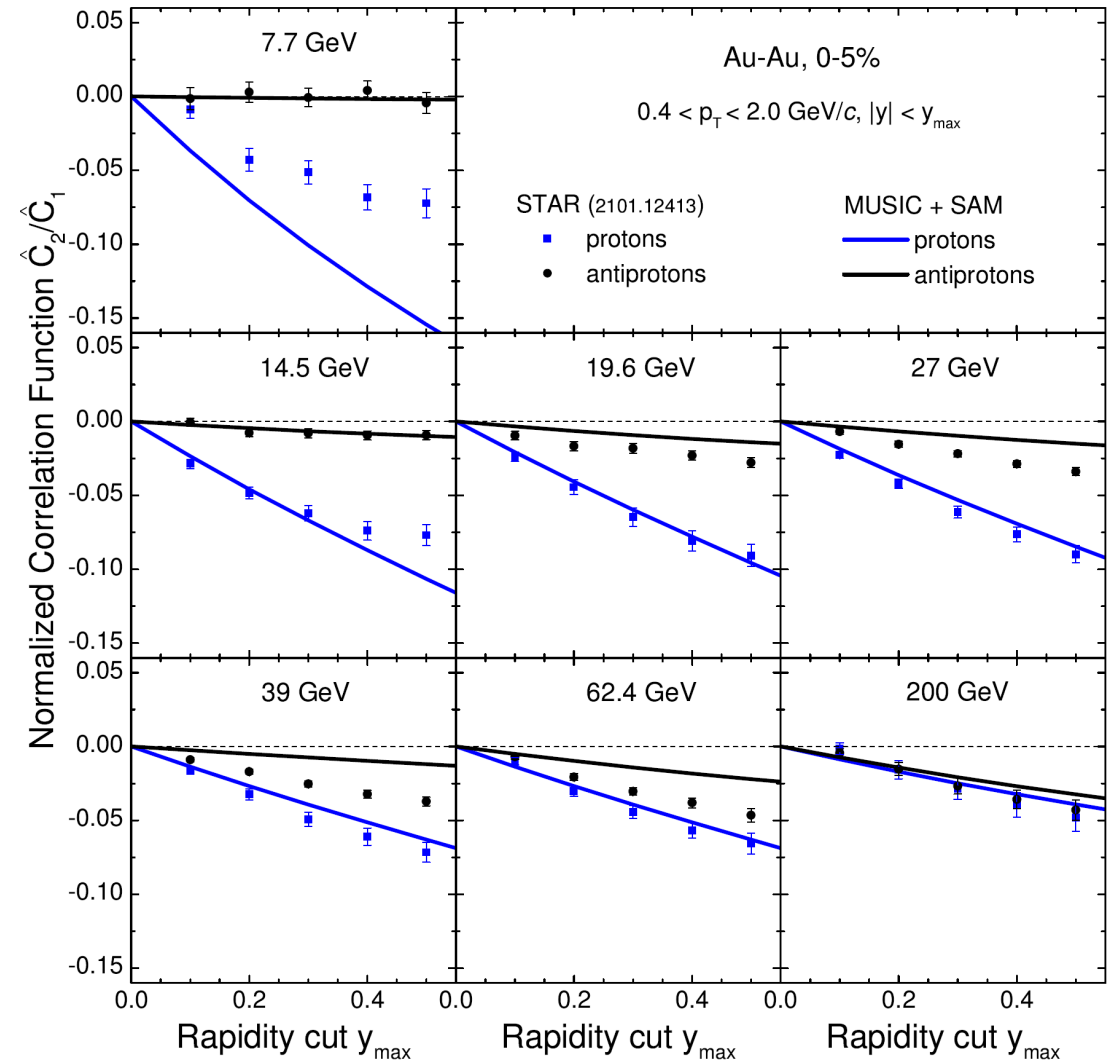
Differences:

- Produced vs stopped?
- Annihilation?



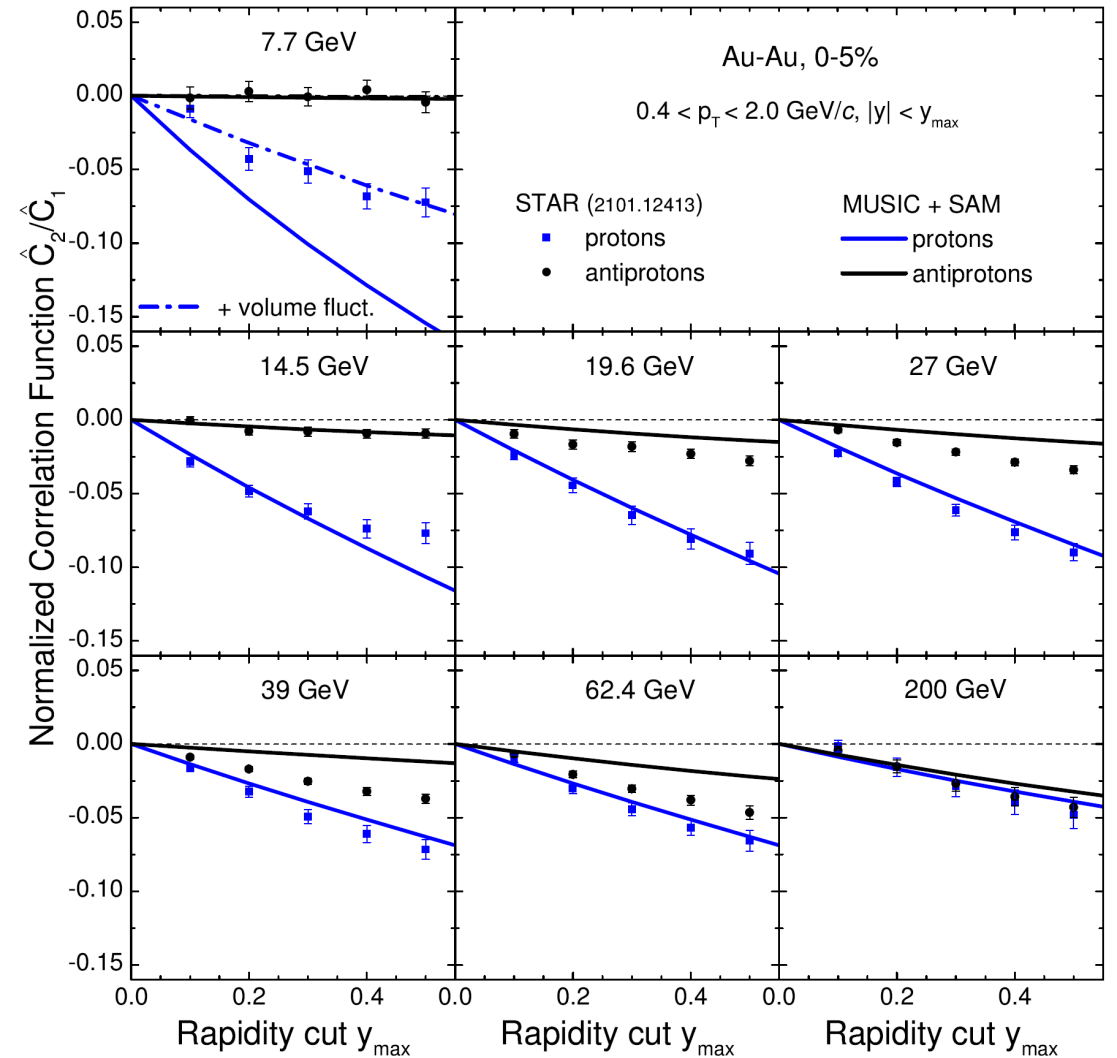
Acceptance dependence of two-particle correlations

- Changing y_{max} slope at $\sqrt{s_{NN}} \leq 14.5$ GeV?



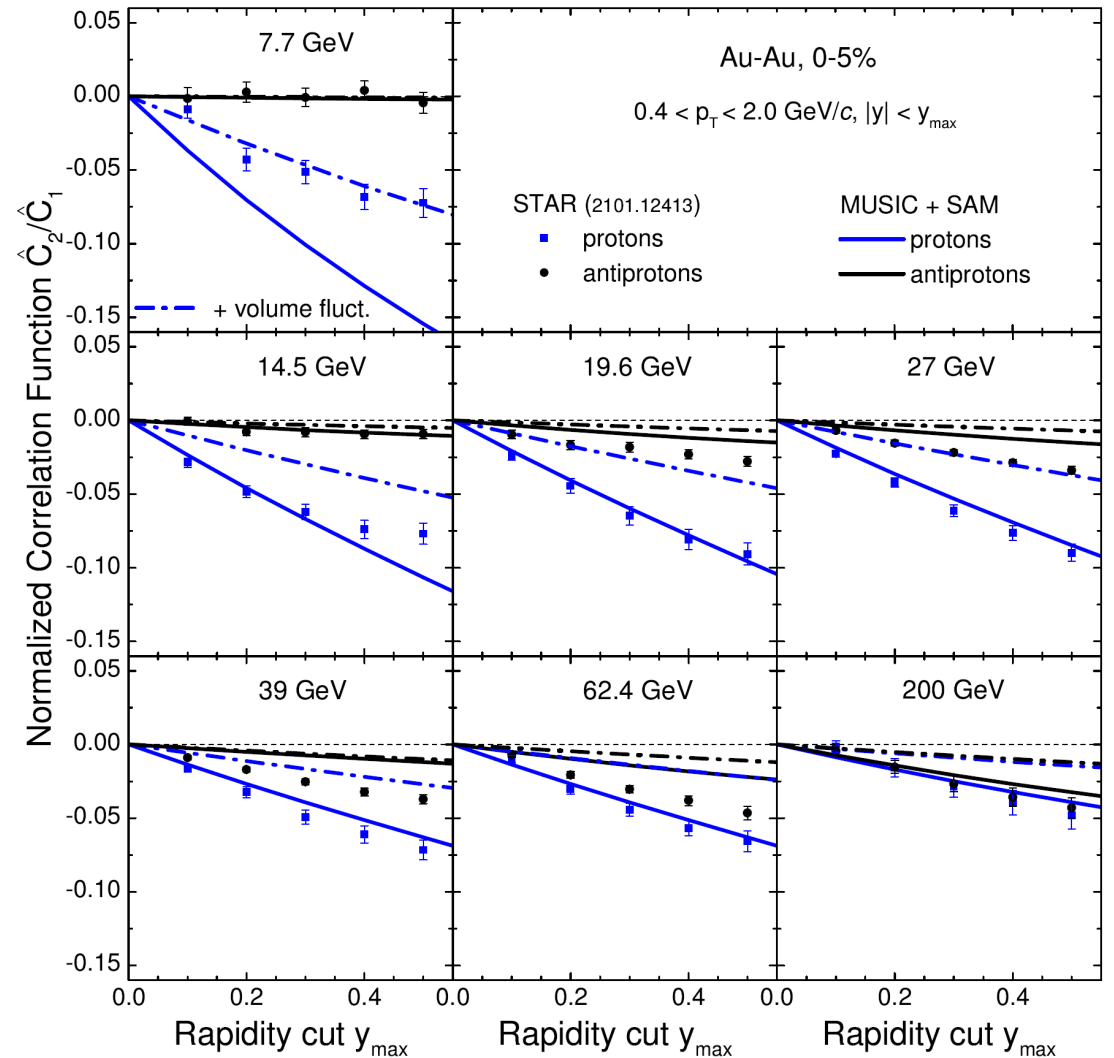
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- **Volume fluctuations?** [Skokov, Friman, Redlich, PRC '13]
 - $C_2/C_1 \neq C_1 * \Delta v^2$



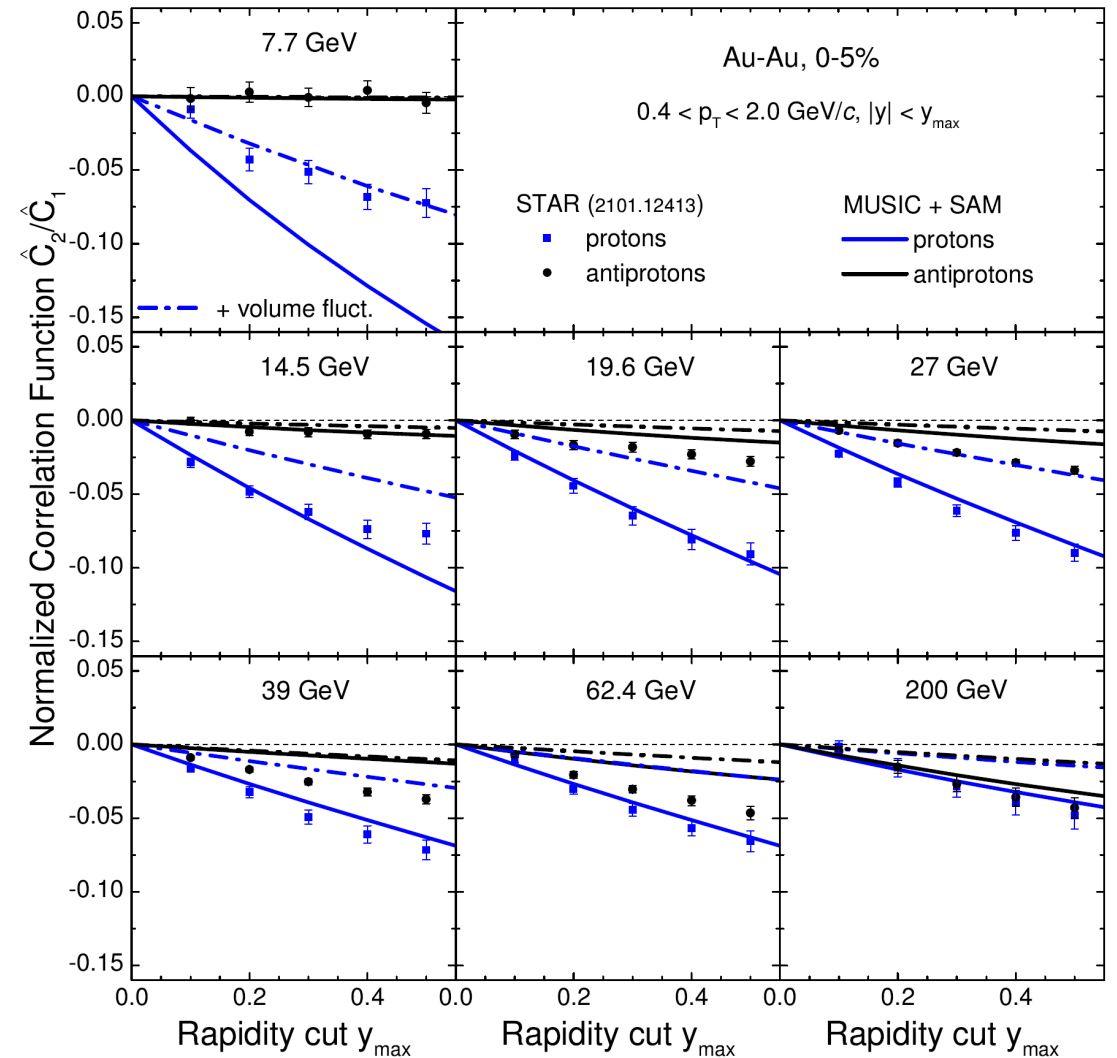
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 - $C_2/C_1 \neq C_1 * \Delta v^2$
 - Can improve low energies but spoil high energies?



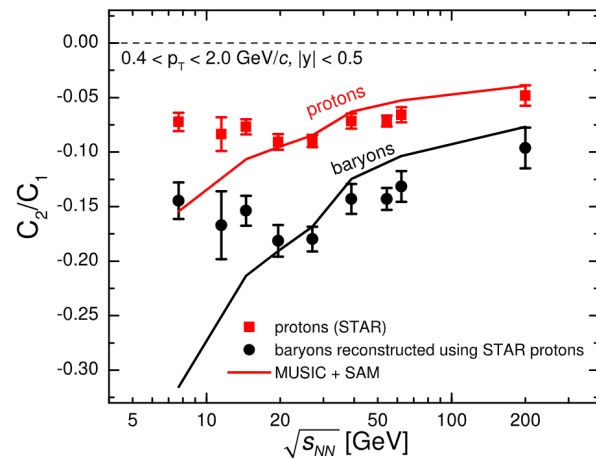
Acceptance dependence of two-particle correlations

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- Volume fluctuations? [Skokov, Friman, Redlich, PRC '13]
 - $C_2/C_1 \neq C_1 * \Delta v^2$
 - Can improve low energies but spoil high energies?
- **Attractive interactions?**
 - Could work if baryon repulsion turns into attraction in the high- μ_B regime
 - **Critical point?**



Baryon cumulants from protons

- **net baryon \neq net proton**
 - protons are a **subset** of all baryons
 - effectively amounts to additional efficiency correction
 - “Poissonizer” of proton cumulants relative to baryons
 - **loss:** $\sim 50\%$ in \hat{C}_2 , $\sim 75\%$ in \hat{C}_3 , $\sim 87.5\%$ in \hat{C}_4
- Baryon cumulants can be reconstructed from proton cumulants based on isospin randomization [Kitazawa, Asakawa, Phys. Rev. C 85 (2012) 021901]
 - Requires the use of joint factorial moments



$$\frac{\hat{C}_2^B}{\hat{C}_1^B} \approx 2 \frac{\hat{C}_2^P}{\hat{C}_1^P}$$

